

# A Geometric Approach to CP Violation: Applications to the MCPMFV SUSY Model

## JAE SIK LEE

National Center for Theoretical Sciences, Hsinchu, TAIWAN

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\* based on JHEP 1010:049,2010, arXiv:1006.3087 [hep-ph] and arXiv:1009.1151 [math.OC] with J. Ellis and A. Pilaftsis

- The SUSY models such as the MSSM contain many possible sources of flavour and CP violation in the soft SUSY-breaking sector:
  - Gaugino mass terms:  $3 \oplus 3 = 6$

$$\mathcal{L}_{\text{soft}} \supset \frac{1}{2} (M_3 \,\widetilde{g}\widetilde{g} + M_2 \,\widetilde{W}\widetilde{W} + M_1 \,\widetilde{B}\widetilde{B} + \text{h.c.})$$

 $30 \oplus 33 \oplus 46 = 109$ 

- Trilinear a terms  $\mathbf{a}_{\mathbf{f}ij} \equiv \mathbf{h}_{\mathbf{f}ij} \cdot \mathbf{A}_{\mathbf{f}ij}$ :  $3 \times (3 \oplus 6 \oplus 9) = 54$ 

$$-\mathcal{L}_{\text{soft}} \supset \left(\widetilde{u}_R^* \, \mathbf{a_u} \, \widetilde{Q} H_2 - \widetilde{d}_R^* \, \mathbf{a_d} \, \widetilde{Q} H_1 - \widetilde{e}_R^* \, \mathbf{a_e} \, \widetilde{L} H_1 + \text{h.c.}\right)$$

- Sfermion mass terms:  $5 \times (3 \oplus 3 \oplus 3) = 45$ 

 $-\mathcal{L}_{\text{soft}} \supset \widetilde{Q}^{\dagger} \mathbf{M}_{\widetilde{\mathbf{Q}}}^{\mathbf{2}} \widetilde{Q} + \widetilde{L}^{\dagger} \mathbf{M}_{\widetilde{\mathbf{L}}}^{\mathbf{2}} \widetilde{L} + \widetilde{u}_{R}^{*} \mathbf{M}_{\widetilde{\mathbf{u}}}^{\mathbf{2}} \widetilde{u}_{R} + \widetilde{d}_{R}^{*} \mathbf{M}_{\widetilde{\mathbf{d}}}^{\mathbf{2}} \widetilde{d}_{R} + \widetilde{e}_{R}^{*} \mathbf{M}_{\widetilde{\mathbf{e}}}^{\mathbf{2}} \widetilde{e}_{R}$ 

– Higgs mass terms:  $3 \oplus \mathbf{1} = 4$ 

$$-\mathcal{L}_{\text{soft}} \supset M_{H_u}^2 H_2^{\dagger} H_2 + M_{H_d}^2 H_1^{\dagger} H_1 - (m_{12}^2 H_1 H_2 + \text{h.c.})$$

• Recently, we have suggested MCPMFV framework with the maximal set of flavour-singlet mass scales: J. Ellis, JSL and A. Pilaftsis, Phys. Rev. D **76** (2007) 115011, [arXiv:0708.2079 [hep-ph]]

 $13 \oplus 6 = 19$  Parameters !

For related approaches, see,

- M. Argyrou, A. B. Lahanas and V. C. Spanos, JHEP 0805 (2008) 026; [arXiv:0804.2613 [hep-ph]]
- G Colangelo, E. Nikolidakis and C. Smith, Eur. Phys. J. C 59 (2009) 75; [arXiv:0807.0801 [hep-ph]]
- W. Altmannshofer, A. J. Buras and P. Paradisi, Phys. Lett. B 669 (2008) 239; [arXiv:0808.0707 [hep-ph]]
- L. Mercolli and C. Smith, Nucl. Phys. B 817 (2009) 1; [arXiv:0902.1949 [hep-ph]]
- A. L. Kagan, G. Perez, T. Volansky and J. Zupan, Phys. Rev. D 80 (2009) 076002; [arXiv:0903.1794 [hep-ph]]
- R. Zwicky and T. Fischbacher, Phys. Rev. D 80 (2009) 076009; [arXiv:0908.4182 [hep-ph]]
- J. Ellis, R. N. Hodgkinson, JSL and A. Pilaftsis, JHEP 1002 (2010) 016; [arXiv:0911.3611 [hep-ph]]

• Then, who ordered "more" CP violation beyond the SM CKM phase? A. D. Sakharov, JETP Letters 5(1967)24



*CP violation in the SM is too weak to explain the matter dominance of the Universe* J. Cline, arXiv:hep-ph/0609145

The matter-dominated Universe did!

• Electric Dipole Moments (EDMs): T violation  $\Rightarrow$  CP violation (under CPT )



 $|d_{\rm Tl}| < 9 \times 10^{-25} {\rm e\,cm}$ ,  $|d_{\rm Hg}| < 3.1 \times 10^{-29} {\rm e\,cm}$ ,  $|d_{\rm n}| < 3 \times 10^{-26} {\rm e\,cm}$ B. C. Regan, E. D. Commins, C. J. Schmidt and D. DeMille, PRL**88** (2002) 071805; W. C. Griffith, M. D. Swallows, T. H. Loftus, M. V. Romalis, B. R. Heckel and E. N. Fortson, PRL**102** (2009) 101601; C. A. Baker *et al.*, PRL**97** (2006) 131801

Question: No large CP phases?

<u>A scan method</u>: See, for example, V. D. Barger, T. Falk, T. Han, J. Jiang, T. Li and T. Plehn, Phys. Rev. D 64 (2001) 056007 [arXiv:hep-ph/0101106].

 $\diamond$  15 parameter MSSM:



I				
0	$\leq$	$ heta_{\mu}, \phi_{1,3},  heta_{Ad,e,u,t}$	$\leq$	$2\pi$
$100~{ m GeV}$	$\leq$	$\mu$	$\leq$	$1 \mathrm{TeV}$
$100~{ m GeV}$	$\leq$	$2M_1, M_2, M_3$	$\leq$	$1 \mathrm{TeV}$
$0~{ m GeV}$	$\leq$	A	$\leq$	$1 \mathrm{TeV}$
$0~{ m GeV}$	$\leq$	$m_{ ilde{e}_R}, m_{ ilde{u}_R}$	$\leq$	$1 { m TeV}$
2	<	$\tan \beta$	<	10

Open circles suffer from parameter tuning  $\Delta X/X$  worse than 1%. Green dots correspond to configurations with a light Higgs  $m_h < 113~GeV.$ 

<u>A scan method</u>: See, for example, V. D. Barger, T. Falk, T. Han, J. Jiang, T. Li and T. Plehn, Phys. Rev. D 64 (2001) 056007 [arXiv:hep-ph/0101106].



• A scan method: J. R. Ellis, JSL and A. Pilaftsis, JHEP 0810 (2008) 049 [arXiv:0808.1819 [hep-ph]]



• <u>A scan method</u> is like "shooting in the dark" ...

blind, time consuming, no quiding principle, etc

Any analytic, exact, and more effective method?



• A linear approximation: We consider the case with N CP-violating phases

In the N-dimensional CP-phase space, we define

*N*-D phase vector  $\boldsymbol{\Phi} = (\Phi_1, \Phi_2, \cdots, \Phi_N)$ 

and then any CP-odd observable O and any EDM E can be expanded as

$$O = \mathbf{\Phi} \cdot \mathbf{O} + \cdots; \quad E = \mathbf{\Phi} \cdot \mathbf{E} + \cdots$$

Formally, we define

$$\mathbf{O} \equiv \nabla O; \qquad \mathbf{E} \equiv \nabla E$$

with  $\nabla \equiv (\partial/\partial \phi_1, \partial/\partial \phi_2, \cdots, \partial/\partial \phi_N)$ 

• [Simple 3D example] with 3 CP phases and 1 EDM constraints: EDM-free subspace and Optimal direction in the linear approximation



• THE HIGHER-DIMENSIONAL GENERALIZATION | with

N CP phases and n EDM constraints

The N-dimensional vector of the optimal direction

$$\Phi^*{}_{\alpha} = \epsilon_{\alpha\beta_1\cdots\beta_n\gamma_1\cdots\gamma_{N-n-1}} E^{(1)}_{\beta_1}\cdots E^{(n)}_{\beta_n} B_{\gamma_1\cdots\gamma_{N-n-1}}$$

where (N-n-1)-dimensional B form is

$$B_{\gamma_1\cdots\gamma_{N-n-1}} = \epsilon_{\gamma_1\cdots\gamma_{N-n-1}\sigma\beta_1\cdots\beta_n} O_{\sigma} E_{\beta_1}^{(1)}\cdots E_{\beta_n}^{(n)}$$

The maximum allowed value of O is:

$$O^{\max} = (\phi^*)^{\max} \widehat{\Phi}^* \cdot \phi^*$$

with the

normalized optimal-direction vector  $\widehat{\Phi}^*$  and  $\phi^*$  which may be practically determined by the validity of the small-phase approximation of the EDMs.

• How good is the linear approximation? The quadratic fit to the Thallium EDM that is used to obtain the 6D vector  $\mathbf{E}^{d_{\text{Tl}}} \equiv \nabla(d_{\text{Tl}}/d_{\text{Tl}}^{\text{EXP}})$  in an expansion around  $\widetilde{\varphi}_{\alpha} = 0^{\circ}$  for the MCPMFV scenario:  $|M_{1,2,3}| = 250 \text{ GeV}, M_{H_u}^2 = M_{H_d}^2 = \widetilde{M}_Q^2 = \widetilde{M}_U^2 = \widetilde{M}_D^2 = \widetilde{M}_L^2 = \widetilde{M}_E^2 = (100 \text{ GeV})^2, |A_u| = |A_d| = |A_e| = 100 \text{ GeV}$ , and  $\tan \beta = 40$ .



- We are considering J. Ellis, JSL, and A. Pilaftsis, JHEP 0810:049,2008, arXiv:0808.1819 [hep-ph]; K. Cheung, O. C. W. Kong, and JSL, JHEP 0906:020,2009, arXiv:0904.4352 [hep-ph]; J. Ellis, JSL and A. Pilaftsis, Phys. Rev. D 76 (2007) 115011, [arXiv:0708.2079 [hep-ph]]
  - Thallium EDM
  - Neutron EDM
  - Mercury EDM
  - Deuteron EDM
  - muon EDM
  - $A_{\rm CP}(b \rightarrow s\gamma)$

These are all implemented in CPsuperH2.2

• Thallium EDM;I.B. Khriplovich and S.K. Lamoreaux, *CP Violation Without Strangeness* (Springer, New York, 1997); M. Pospelov and A. Ritz, Annals Phys. **318** (2005) 119,[arXiv:hep-ph/0504231]

$$d_{\rm Tl} [e \,{\rm cm}] = -585 \cdot d_e^E [e \,{\rm cm}] - 8.5 \times 10^{-19} [e \,{\rm cm}] \cdot (C_S \,{\rm TeV}^2) + \cdots$$

$$d_{e}^{E} = (d_{e}^{E})^{\tilde{\chi}^{\pm}} + (d_{e}^{E})^{\tilde{\chi}^{0}} + (d_{e}^{E})^{H}$$
$$C_{S} = (C_{S})^{4f} + (C_{S})^{g}$$

where  $(C_S)^{4f} = C_{de} \frac{29 \operatorname{MeV}}{m_d} + C_{se} \frac{\kappa \times 220 \operatorname{MeV}}{m_s}$  with  $\kappa \equiv \langle N | m_s \bar{s}s | N \rangle / 220 \operatorname{MeV} \simeq 0.50 \pm 0.25$  and  $(C_S)^g = (0.1 \operatorname{GeV}) \frac{m_e}{v^2} \sum_{i=1}^3 \frac{g_{H_i gg}^S g_{H_i \bar{e}e}^P}{M_{H_i}^2}$ 

Neutron EDM [QCD sum rule techniques (QCD)];M. Pospelov and A. Ritz, Phys. Rev. Lett. 83 (1999) 2526, [arXiv:hep-ph/9904483]; M. Pospelov and A. Ritz, Nucl. Phys. B 573 (2000) 177, [arXiv:hep-ph/9908508];
 M. Pospelov and A. Ritz, Phys. Rev. D 63 (2001) 073015, [arXiv:hep-ph/0010037]; D. A. Demir, M. Pospelov and A. Ritz, Phys. Rev. D 67 (2003) 015007,[arXiv:hep-ph/0208257]; D. A. Demir, O. Lebedev, K. A. Olive, M. Pospelov and A. Ritz, Nucl. Phys. B 680 (2004) 339, [arXiv:hep-ph/0311314]; K. A. Olive, M. Pospelov, A. Ritz and Y. Santoso, Phys. Rev. D 72 (2005) 075001, [arXiv:hep-ph/0506106]

$$d_n = d_n(d_q^E, d_q^C) + d_n(d^G) + d_n(C_{bd}) + \cdots,$$

$$d_n(d_q^E, d_q^C) = (1.4 \pm 0.6) (d_d^E - 0.25 d_u^E) + (1.1 \pm 0.5) e (d_d^C + 0.5 d_u^C) / g_s,$$
  

$$d_n(d^G) \sim \pm e (20 \pm 10) \text{ MeV } d^G,$$
  

$$d_n(C_{bd}) \sim \pm e 2.6 \times 10^{-3} \text{ GeV}^2 \left[ \frac{C_{bd}}{m_b} + 0.75 \frac{C_{db}}{m_b} \right]$$

where  $d^G = d^G(1 \text{ GeV}) \simeq 8.5 d^G(\text{EW})$ 

Mercury EDM; D. A. Demir, O. Lebedev, K. A. Olive, M. Pospelov and A. Ritz, Nucl. Phys. B 680 (2004) 339, [arXiv:hep-ph/0311314]; K. A. Olive, M. Pospelov, A. Ritz and Y. Santoso, Phys. Rev. D 72 (2005) 075001, [arXiv:hep-ph/0506106]

$$\begin{aligned} d_{\rm Hg} &= 7 \times 10^{-3} e \left( d_u^C - d_d^C \right) / g_s \ + \ 10^{-2} d_e^E \\ &- 1.4 \times 10^{-5} e \,{\rm GeV}^2 \, \left[ \frac{0.5 C_{dd}}{m_d} + 3.3 \kappa \frac{C_{sd}}{m_s} + (1 - 0.25 \kappa) \frac{C_{bd}}{m_b} \right] \\ &+ (3.5 \times 10^{-3} \,\,{\rm GeV}) e \, C_S \\ &+ (4 \times 10^{-4} \,\,{\rm GeV}) e \, \left[ C_P + \left( \frac{Z - N}{A} \right)_{\rm Hg} \, C'_P \right] \\ \end{aligned}$$
where  $\mathcal{L}_{C_P} = C_P \, \bar{e} e \, \bar{N} i \gamma_5 N \ + \ C'_P \, \bar{e} e \, \bar{N} i \gamma_5 \tau_3 N$  with

$$C_P = (C_P)^{4f} \simeq -375 \text{ MeV} \sum_{q=c,s,t,b} \frac{C_{eq}}{m_q},$$
  
 $C'_P = (C'_P)^{4f} \simeq -806 \text{ MeV} \frac{C_{ed}}{m_d} - 181 \text{ MeV} \sum_{q=c,s,t,b} \frac{C_{eq}}{m_q}$ 

Deuteron EDM; O. Lebedev, K. A. Olive, M. Pospelov and A. Ritz, Phys. Rev. D 70 (2004) 016003, [arXiv:hep-ph/0402023]

$$d_D \simeq -\left[5^{+11}_{-3} + (0.6 \pm 0.3)\right] e \left(d_u^C - d_d^C\right)/g_s$$
  
-(0.2 \pm 0.1) e  $\left(d_u^C + d_d^C\right)/g_s + (0.5 \pm 0.3)(d_u^E + d_d^E)$   
+(1 \pm 0.2) \times 10^{-2} e \text{ GeV}^2  $\left[\frac{0.5C_{dd}}{m_d} + 3.3\kappa\frac{C_{sd}}{m_s} + (1 - 0.25\kappa)\frac{C_{bd}}{m_b}\right]$   
\pm e (20 \pm 10) MeV  $d^G$ 

The projective sensitivity: Y. K. Semertzidis *et al.* [EDM Collaboration], AIP Conf. Proc. **698** (2004) 200, [arXiv:hep-ex/0308063]; Y. F. Orlov, W. M. Morse and Y. K. Semertzidis, Phys. Rev. Lett. **96** (2006) 214802, [arXiv:hep-ex/0605022]; and also see http://www.bnl.gov/edm/deuteron\_proposal\_080423\_final.pdf

$$|d_D| < 3 \times 10^{-27} - 10^{-29} \ e \,\mathrm{cm}$$

For our numerical study, we take  $3 \times 10^{-27} e \,\mathrm{cm}$  as a representative expected value

CP-violating QCD θ-term: O. Lebedev, K. A. Olive, M. Pospelov and A. Ritz, Phys. Rev. D 70 (2004) 016003, [arXiv:hep-ph/0402023]; M. Pospelov and A. Ritz, Annals Phys. 318 (2005) 119, [arXiv:hep-ph/0504231]; J. Ellis, JSL, and A. Pilaftsis, arXiv:1006.3087 [hep-ph]

$$\mathcal{L} = \frac{\alpha_s}{8\pi} \bar{\theta} \, G^a_{\mu\nu} \tilde{G}^{\mu\nu,a} \quad \text{with} \quad \bar{\theta} = \theta_{\text{QCD}} + \text{Arg Det } M_q$$

$$d_n(\bar{\theta}) = (1 \pm 0.5) \frac{|\langle \bar{q}q \rangle|}{(225 \text{ MeV})^3} \bar{\theta} \times 2.5 \times 10^{-16} e \text{ cm}$$

$$d_{\rm Hg}(\bar{\theta}) \simeq +2.0 \times 10^{-6} \,\bar{\theta} \, e \, {\rm GeV}^{-1} \simeq 3.9 \times 10^{-20} \,\bar{\theta} \, e \, {\rm cm}$$
$$d_D(\bar{\theta}) \simeq -e \left[ (3.5 \pm 1.4) + (1.4 \pm 0.4) \right] \times 10^{-3} \,\bar{\theta} \, {\rm GeV}^{-1} \simeq -9.7 \times 10^{-17} \,\bar{\theta} \, e \, {\rm cm}$$

• [Summary] EDMs and Observables under consideration

$$d_{\rm Tl}/d_{\rm Tl}^{\rm EXP}, \quad d_{\rm n}/d_{\rm n}^{\rm EXP}, \quad d_{\rm Hg}/d_{\rm Hg}^{\rm EXP},$$
$$d_{\rm D}/d_{\rm D}^{\rm EXP}, \quad d_{\mu}/d_{\mu}^{\rm EXP}, \quad A_{\rm CP}(b \to s\gamma)[\%],$$

where we choose the following normalization factors

$$\begin{split} d_{\rm Tl}^{\rm EXP} &= 9 \times 10^{-25} \, e \, {\rm cm} \,, \ \ d_{\rm n}^{\rm EXP} = 3 \times 10^{-26} \, e \, {\rm cm} \,, \ \ d_{\rm Hg}^{\rm EXP} = 3.1 \times 10^{-29} \, e \, {\rm cm} \,, \\ d_{\rm D}^{\rm EXP} &= 3 \times 10^{-27} \, e \, {\rm cm} \,, \ \ d_{\mu}^{\rm EXP} = 1 \times 10^{-24} \, e \, {\rm cm} \end{split}$$

• The EDMs and Observables under consideration are functions of

# 7 parameters

$$\Phi_1, \quad \Phi_2, \quad \Phi_3, \quad \Phi_{A_u}, \quad \Phi_{A_d}, \quad \Phi_{A_e}, \quad \overline{\theta}$$

and then

$$\nabla_{\alpha} \equiv \left(\frac{\partial}{\partial \Phi_{1}}, \ \frac{\partial}{\partial \Phi_{2}}, \ \frac{\partial}{\partial \Phi_{3}}, \ \frac{\partial}{\partial \Phi_{A_{u}}}, \ \frac{\partial}{\partial \Phi_{A_{d}}}, \ \frac{\partial}{\partial \Phi_{A_{e}}}, \ \frac{\partial}{\partial \widehat{\Phi}_{A_{e}}}, \ \frac{\partial}{\partial \widehat{\theta}}\right)$$

The CP-violating phases  $\Phi_{1,2,3}$  and  $\Phi_{A_u,A_d,A_e}$  are specified in degrees and we normalize  $\bar{\theta}$  in units of  $10^{-10}$ :  $\hat{\theta} \equiv \bar{\theta} \times 10^{10}$ 

• <u>A scenario</u>: We consider CP-violating variants of a typical CMSSM scenario with

$$\begin{split} |M_{1,2,3}| &= 250 \text{ GeV}, \\ M_{H_u}^2 &= M_{H_d}^2 = \widetilde{M}_Q^2 = \widetilde{M}_U^2 = \widetilde{M}_D^2 = \widetilde{M}_L^2 = \widetilde{M}_E^2 = (100 \text{ GeV})^2, \\ |A_u| &= |A_d| = |A_e| = 100 \text{ GeV}, \end{split}$$

at the GUT scale, varying  $an eta \left( M_{\mathrm{SUSY}} 
ight)$ 

We adopt the convention that  $\Phi_{\mu} = 0^{\circ}$ , and we vary independently the following six MCPMFV phases at the GUT scale:

$$\Phi_1, \quad \Phi_2, \quad \Phi_3, \quad \Phi_{A_u}, \quad \Phi_{A_d}, \quad \Phi_{A_e}$$

in addition to the QCD  $\theta$  term:  $\overline{\theta}$ 

This scenario becomes the SPS1a point when  $\tan \beta = 10$ ,  $\Phi_{1,2,3} = 0^{\circ}$  and  $\Phi_{Au,A_d,Ae} = 180^{\circ}$ 

• Components of the vectors  $\mathbf{E} \equiv \nabla E$ ,  $\mathbf{O} \equiv \nabla O$ ;  $\widehat{\mathbf{\Phi}}^*$ 



 $\widehat{\Phi}^*{}_{\alpha} = \mathcal{N} \varepsilon_{\alpha\beta\gamma\delta\mu\nu\rho} E_{\beta}^{d_{\mathrm{Tl}}} E_{\gamma}^{d_{\mathrm{Hg}}} B_{\mu\nu\rho}$ with the 3-form  $B_{\mu\nu\rho} = \varepsilon_{\mu\nu\rho\lambda\sigma\tau\omega} O_{\lambda} E_{\sigma}^{d_{\mathrm{Tl}}} E_{\tau}^{d_{\mathrm{Hg}}} E_{\omega}^{d_{\mathrm{Hg}}}$  or  $N_{\mu}^{(1)} N_{\nu}^{(2)} N_{\rho}^{(3)}$  for some reference directions

• The products 
$$\Phi^* \cdot \mathbf{O}$$
 Recall the relation  $O^{\max} = (\phi^*)^{\max} \widehat{\Phi}^* \cdot \mathbf{O}$ 



• The maximum values of the observables along the optimal directions  $\tan \beta = 10$ Again, recall the relation  $O^{\max} = (\phi^*)^{\max} \widehat{\Phi}^* \cdot \mathbf{O}$ 



♦ The maximum values of  $\phi^*$  for each EDM-free direction from the figure:  $(\phi^*)^{\max} \sim 25 (d_{\rm D}\text{-optimal}), 25 (d_{\mu}\text{-optimal}), 50 (A_{\rm CP}\text{-optimal}), 40 (\Delta \Phi_{1,A_e} = \hat{\theta} = 0), \text{ and } 45 (\Delta \Phi_{2,3} = \hat{\theta} = 0),$ which are mainly constrained by  $d_{\rm Tl}, d_{\rm Tl}, d_{\rm Tl}, d_{\rm Hg},$  and  $d_{\rm Tl}$ , respectively.

#### ◇ The maximum values of the observables

• The maximum values of the CP phases along the optimal directions  $\tan \beta = 10$ 



In the top panels we see that  $\Phi_1$  and  $\Phi_2$  can be as large as  $\sim 20^{\circ}$  and  $\sim 5^{\circ}$ , respectively, for  $(\phi^*)^{\max} \sim 25$  along the  $d_{\rm D}$ - and  $d_{\mu}$ -optimal directions denoted by the thick solid and dashed lines.

We also note in the middle and bottom panels that the phases of  $A_{d,u,e}$  could be are large, in general, though they are suppressed at the  $M_{SUSY}$  scale.

Finally, we note (not shown) that  $\bar{\theta}$  could be as large as  $\sim 2 \times 10^{-9}$  along the  $d_{\rm D}$ - and  $d_{\mu}$ -optimal directions with  $(\phi^*)^{\rm max} \sim 25$ .

## *♠ Summary*

- We are proposing a geometric method which provides an accurate parametric determination of the optimal cancellation regions where any given physical observable is maximized in the linear approximation
- Our geometric approach is exact, efficient and less computationally-intensive than a naive scan of a multi-dimensional space
- This constitutes an *analytic* solution to the so-called *linear programming problem*
- You may want to apply this method to your problem if you are trying to achieve the best outcome in a given requirements expressed in linear equations