

Warped AdS/CFT Correspondence

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AdS/CFT Correspondence

$\mathcal{N} = 4$ SYM in 4D \simeq IIB String in $AdS_5 \times S^5$

$$S^5: \quad X_1^2 + X_2^2 + X_3^2 + X_4^2 + X_5^2 + X_6^2 = R^2 \quad \text{in } R^6$$

$$AdS_5: \quad -X_{-1}^2 - X_0^2 + X_1^2 + X_2^2 + X_3^2 + X_4^2 = -R^2 \quad \text{in } R^{2,4}$$

- | | | |
|------------|----------------------------------|-----------------------------|
| N=4 SYM | Conformal Sym. in 4D: $SO(4, 2)$ | R-sym.: $SU(4)$ |
| IIB String | Isometry in AdS_5 : $SO(4, 2)$ | Isometry in S^5 : $SO(6)$ |

't Hooft coupling: $\lambda = g_{YM}^2 N = g_s N$, $R = (4\pi g_s N)^{1/4} l_s \sim \lambda^{1/4} l_{pl}$

g_s : string coupling, $2\pi l_s^2$: inverse string tension.

For $\lambda \gg 1$, $R \gg l_{pl}$, SUGRA limit.

- The holographic principle.
- CFT calculation of BH entropy, BH greybody factor.
- PP wave limit.

Let $X_{-1} = R \cosh \rho \cos \tau,$

$X_0 = R \cosh \rho \sin \tau,$

$X_1 = R \sinh \rho \sin \theta_1 \sin \theta_2 \cos \phi, \quad X_2 = R \sinh \rho \sin \theta_1 \sin \theta_2 \sin \phi, \quad (1)$

$X_3 = R \sinh \rho \sin \theta_1 \cos \theta_2, \quad X_4 = R \sinh \rho \cos \theta_1.$

Global coordinate: $ds^2 = R^2 \left[-\cosh^2 \rho d\tau^2 + d\rho^2 + \sinh^2 \rho d\Omega_3^2 \right] \quad (2)$

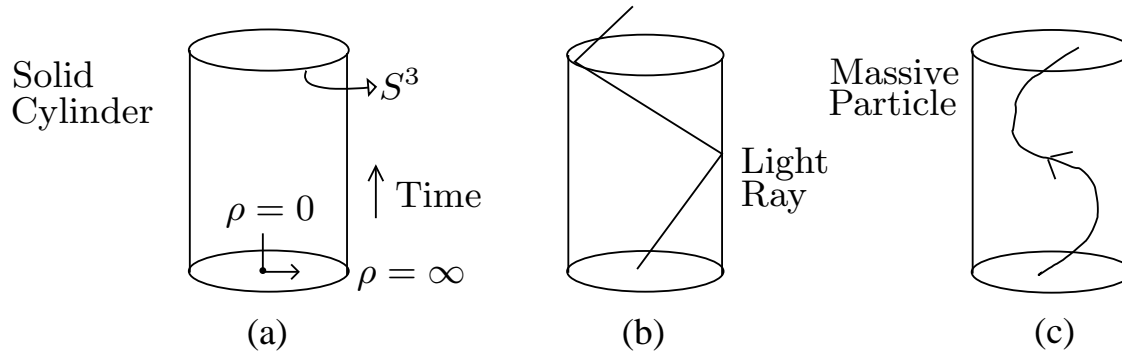


Figure 1. Penrose diagram and the AdS boundary ($S^3 \times R$).

Define $dx = d\rho / \cosh \rho, \quad \tan(x/2) = \tanh(\rho/2), \quad \rho \rightarrow \infty \sim x = \pi/2.$

$\Rightarrow -\cosh^2 \rho d\tau^2 + d\rho^2 = \sec^2 x (-d\tau^2 + dr^2).$

Let $X_{-1} + X_4 = R/z$, $X_\mu = Rx_\mu/z$ for $\mu = 0, \dots, 3$.

Poincare coordinate: $ds^2 = R^2 \left(\frac{-dx_0^2 + dx_1^2 + dx_2^2 + dx_3^2 + dz^2}{z^2} \right)$ (3)

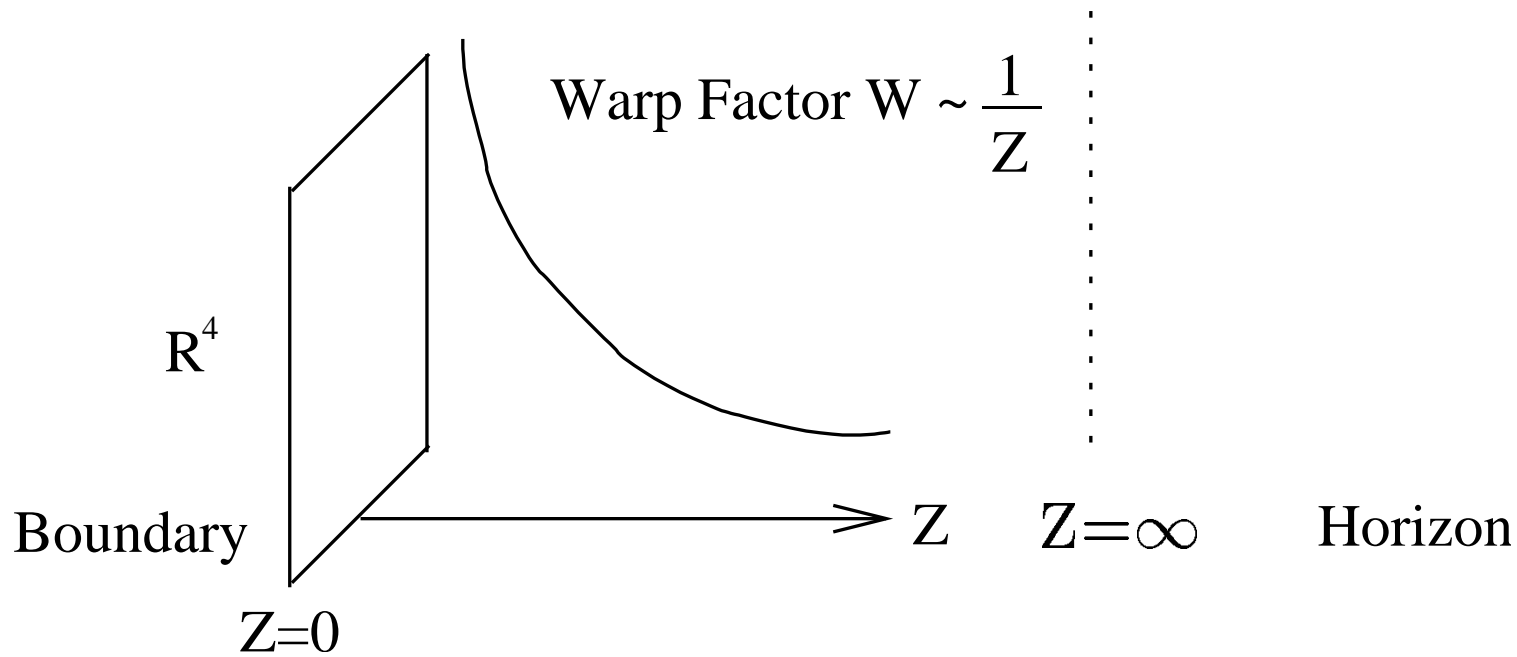
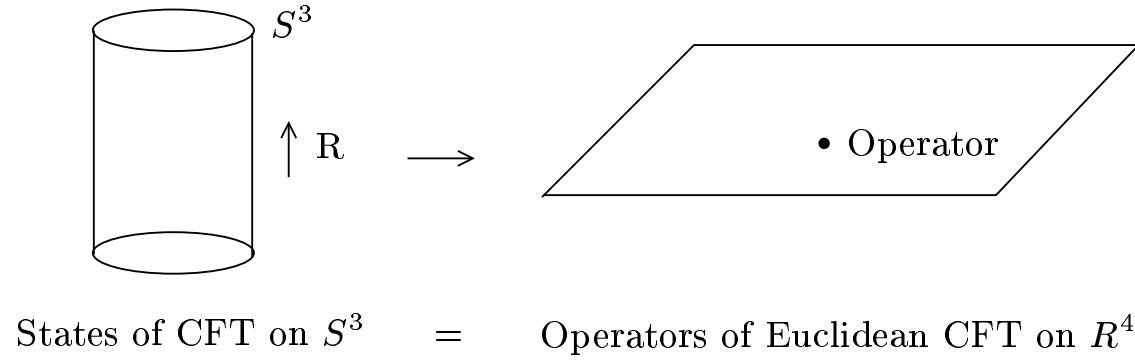


Figure 2. AdS boundary: $S^3 \times R$.



$$\mathcal{Z}_{bulk} [\phi(\vec{x}, z)|_{z=0} = \phi_0(\vec{x})] = \langle e^{\int d^4x \phi_0(\vec{x}) \mathcal{O}(\vec{x})} \rangle_{\text{Field Theory}} \quad (4)$$

$$\text{Scalar field of mass } m: S = N^2 \int \frac{dx^4 dz}{z^5} [z^2 (\partial\phi)^2 + m^2 R^2 \phi^2] \quad (5)$$

$$\Rightarrow z^3 \partial_z \left(\frac{1}{z^3} \partial_z \phi \right) - p^2 \phi - \frac{m^2 R^2}{z^2} \phi = 0 \quad (6)$$

$$\phi = z^2 \left[A_+ I_\nu(pz) + A_- K_\nu(pz) \right], \text{ with } \nu = \sqrt{4 + m^2 R^2}. \quad (7)$$

$$\text{Near } z = 0, \phi \sim B_+ z^{2+\nu} + B_- z^{2-\nu}. \quad (8)$$

Boundary condition: $\phi(x, z)|_{z=\epsilon} = \epsilon^{2-\nu} \phi_0^r(x)$, $\phi_0^r(x)$ fixed for $\epsilon \rightarrow 0$.

The rescaling $x_\mu \rightarrow \lambda x_\mu$, $z \rightarrow \lambda z$ is an isometry in AdS,
 ϕ does not get rescaled, ϕ_0^r has dimension $2 - \nu$,
 \Rightarrow the corresponding operator O has dimension $\Delta \equiv 2 + \nu$.

$$\text{Bulk Green function in } AdS_5: G_\Delta(z, x^\mu, x'^\mu) = \frac{z^\Delta}{[(x - x')^2 + z^2]^\Delta} \quad (9)$$

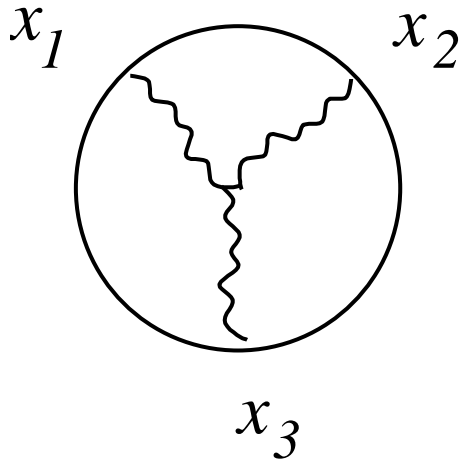


Figure 4. Three point functions.

With a cubic term in the action,

$$\langle \mathcal{O}_1 \mathcal{O}_2 \mathcal{O}_3 \rangle \sim \int \frac{d^4 x dz}{z^5} G_\Delta(z, x, x_1) G_\Delta(z, x, x_2) G_\Delta(z, x, x_3). \quad (10)$$

$$\begin{cases} 0 > m^2 R^2 \geq -4, & 2 - \nu < 0, & \phi \text{ induces relevant perturbation;} \\ m^2 R^2 > 0, & 2 - \nu > 0, & \phi \text{ induces irrelevant perturbation.} \end{cases}$$

$$\begin{cases} (E_{\text{FT}}) = \frac{1}{z}(E_{\text{proper}}); \\ (\text{size})_{\text{FT}} = z(\text{proper size}). \end{cases} \quad \begin{cases} z \rightarrow 0 \sim \text{spatial infinity, UV;} \\ z \rightarrow \infty \sim \text{horizon, IR.} \end{cases}$$

IR/UV Correspondence: UV in field theory \sim IR in gravity .

Blackhole \Rightarrow Hawking temperature(T_H) \Rightarrow thermal effect.

Simplest black hole in Poicare coordinate (S^5 suppressed):

$$ds^2 = \frac{R^2}{z^2} \left[- \left(1 - \frac{z^4}{z_0^4} \right) dt^2 + d\vec{x}^2 + \left(1 - \frac{z^4}{z_0^4} \right)^{-1} dz^2 \right]. \quad (11)$$

$$\text{For } z \simeq z_0, ds^2 \propto \left[-16 \left(1 - \frac{z}{z_0} \right)^2 dt^2 + dz^2 + \dots \right]. \quad (12)$$

Period in imaginary time: $\beta = \pi z_0/2$.

Bekenstein Hawking entropy: $S_{BH} = \frac{(\text{Area})}{4G_N} \Rightarrow \frac{S_{BH}}{V} = \frac{\pi^2}{2} N^2 T^3. \quad (13)$

Weakly coupled field theory entropy: $\frac{S_{FT}}{V} = \frac{4\pi^2}{3 \cdot 2} N^2 T^3. \quad (14)$

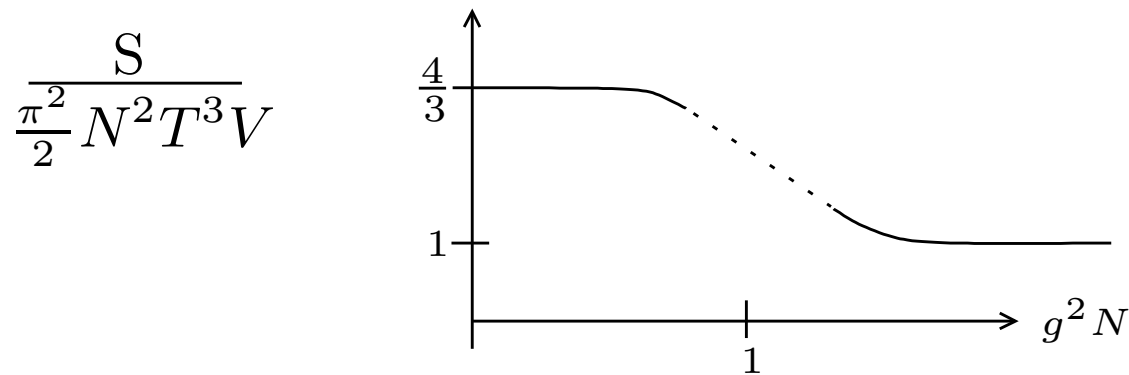


Figure 5. Field theory entropy vs. Bekenstein Hawking entropy.

$g^2 N$ correction to S_{FT} and R^4 to S_{BH} go in the right direction.

AdS₃/CFT₂ Correspondence

CFT interpretation of BH QMN

The BTZ blackhole ($R = 1$), part of SUGRA:

$$ds^2 = - \left[\frac{(r^2 - r_+^2)(r^2 - r_-^2)}{r^2} \right] dt^2 + \left[\frac{r^2}{(r^2 - r_+^2)(r^2 - r_-^2)} \right] dr^2 + r^2 \left[d\theta - \frac{r_+ r_-}{r^2} dt \right]^2. \quad (15)$$

$$M = (r_+^2 + r_-^2), \quad J = 2r_+ r_-.$$

$$T_H = (r_+^2 - r_-^2)/(2\pi r_+), \quad \mathcal{A}_H = 2\pi r_+, \quad \Omega_H = J/(2r_+^2).$$

Locally equivalent to AdS_3 :

$$ds^2 = - \sinh^2 \mu (r_+ dt - r_- d\phi)^2 + d\mu^2 + \cosh^2 \mu (r_+ d\phi - r_- dt)^2. \quad (16)$$

The dual CFT on the boundary is (1+1)-D.

Two independent copies of CFT with

$$T_L = (r_+ - r_-)/(2\pi), \quad T_R = (r_+ + r_-)/(2\pi), \quad 1/T_L + 1/T_R = 2/T_H.$$

Operators \mathcal{O} in (1+1)-D CFT are characterized by the conformal weights: (h_L, h_R) :

$$h_L + h_R = \Delta, \quad h_R - h_L = \pm s. \quad (17)$$

$$\begin{cases} \text{Scalar:} & \Delta = \sqrt{1 + m^2}; \\ \text{Spinor and vector:} & \Delta = 1 + |m|. \end{cases}$$

Retarded Green function: $D_{\text{ret}}(x, x') = i\theta(t - t') \langle [\mathcal{O}(x), \mathcal{O}(x')] \rangle_T$, (18)

$$= i\theta(t - t') \bar{\mathcal{D}}(x, x').$$

$$\bar{\mathcal{D}}(x, x') = \mathcal{D}_+(x, x') - \mathcal{D}_-(x, x'), \quad x^\pm = t \pm \sigma.$$

$$\mathcal{D}_+(x) = \frac{(\pi T_R)^{2h_R}}{\sinh^{2h_R}(\pi T_R x^- - i\epsilon)} \frac{(\pi T_L)^{2h_L}}{\sinh^{2h_L}(\pi T_L x^+ - i\epsilon)}; \quad (19)$$

$$\mathcal{D}_-(x) = \frac{(\pi T_R)^{2h_R}}{\sinh^{2h_R}(\pi T_R x^- + i\epsilon)} \frac{(\pi T_L)^{2h_L}}{\sinh^{2h_L}(\pi T_L x^+ + i\epsilon)}. \quad (20)$$

With $k_{\pm} = (\omega \mp k)/2$, Fourier transform of $\bar{D}(x)$:

$$\bar{D}(k_+, k_-) = \Gamma\left(h_L + \frac{ik_+}{2\pi T_L}\right) \Gamma\left(h_R + \frac{ik_-}{2\pi T_R}\right) \Gamma\left(h_L - \frac{ik_+}{2\pi T_L}\right) \Gamma\left(h_R - \frac{ik_-}{2\pi T_R}\right). \quad (21)$$

Two sets of poles in the lower complex plane of ω :

$$\omega_L = k - 4\pi iT_L(n + h_L); \quad \omega_R = -k - 4\pi iT_R(n + h_R). \quad (22)$$

$n = 0, 1, 2, \dots$. They characterize the decay of the perturbation.

These poles coincide precisely with the QMN in the BTZ background!

Scalar perturbation satisfies
$$\left[\frac{1}{\sqrt{|g|}} \partial_{\mu} (\sqrt{|g|} g^{\mu\nu} \partial_{\nu}) - m^2 \right] \Phi = 0. \quad (23)$$

Let $\Phi = e^{-i(k_+ x^+ + k_- x^-)} f(\mu)$, $x^{\pm} = r_{\pm} t - r_{\mp} \phi$, $z = \tanh^2 \mu$,

$$(k_+ + k_-)(r_+ - r_-) = \omega - k; \quad (k_+ - k_-)(r_+ + r_-) = \omega + k.$$

$$z(1-z)\tilde{f}''(z) + (1-z)\tilde{f}'(z) + \left[\frac{k_+^2}{4z} - \frac{k_-^2}{4} - \frac{m^2}{4(1-z)} \right] \tilde{f}(z) = 0. \quad (24)$$

$$\tilde{f}(z) = z^\alpha (1-z)^{\beta_s} {}_2F_1(a_s, b_s; c_s; z), \quad (25)$$

$$\alpha = -ik_+/2, \quad \beta_s = \left(1 - \sqrt{1+m^2}\right)/2$$

$$a_s = -i(k_+ - k_-)/2 + \beta_s, \quad b_s = -i(k_+ + k_-)/2 + \beta_s, \quad c_s = 1 + 2\alpha.$$

QNM: vanishing Dirichlet condition at infinity.

$$\text{Flux: } \mathcal{F} = \sqrt{g} \frac{1}{2i} (f^* \partial_\mu f - f \partial_\mu f^*) \propto \left| \frac{\Gamma(c_s) \Gamma(c_s - a_s - b_s)}{\Gamma(c_s - a_s) \Gamma(c_s - b_s)} \right|^2. \quad (26)$$

\mathcal{F} vanishes if

$$\begin{cases} c_s - a_s = -n \\ c_s - b_s = -n \end{cases} \Rightarrow i(k_+ \pm k_-)/2 = n + \left(1 + \sqrt{1+m^2}\right)/2. \quad (27)$$

For spinors,

$$\beta_f = -(m + 1/2)/2, \quad c_f = 1/2 + 2\alpha.$$

$$a_f = -i(k_+ - k_-)/2 + \beta_f + 1/2, \quad b_f = -i(k_+ + k_-)/2 + \beta_f.$$

For vectors,

$$\beta_v = m/2, \quad c_v = 1 + 2\alpha.$$

$$a_v = -i(k_+ - k_-)/2 + \beta_v, \quad b_v = -i(k_+ + k_-)/2 + \beta_v.$$

CFT interpretation of BH Greybody Factor

$$ds^2 = - \left[\frac{(r^2 - r_+^2)(r^2 - r_-^2)}{r^2} \right] dt^2 + \left[\frac{r^2}{(r^2 - r_+^2)(r^2 - r_-^2)} \right] dr^2 + r^2 \left[d\theta - \frac{r_+ r_-}{r^2} dt \right]^2. \quad (28)$$

Consider massless scalar Φ . Let $\Phi = e^{-i\omega t + im\theta} R_{\omega, m}(r)$,

$$\begin{aligned} \partial_r^2 R_{\omega, m}(r) + \left[-\frac{1}{r} + \frac{2r}{r^2 - r_+^2} + \frac{2r}{r^2 - r_-^2} \right] \partial_r R_{\omega, m}(r) \\ + \frac{r^4}{(r^2 - r_+^2)^2 (r^2 - r_-^2)^2} \left[\omega^2 - m^2 + \frac{Mm^2 - J\omega m}{r^2} \right] R_{\omega, m}(r) = 0. \end{aligned} \quad (29)$$

Let $z = (r^2 - r_+^2)/(r^2 - r_-^2)$, $A_1 = \left(\frac{\omega - m\Omega_H}{4\pi T_H} \right)^2$, $B_1 = - \left(\frac{\omega r_-^2 - m\Omega_H r_+^2}{4\pi T_H r_+ r_-} \right)^2$

$$z(1-z)\tilde{R}''_{\omega, m}(z) + (1-z)\tilde{R}'_{\omega, m}(z) + \left[\frac{A_1}{z} + B_1 \right] \tilde{R}_{\omega, m}(z) = 0. \quad (30)$$

$z = 0 \sim$ horizon, $z \rightarrow \infty \sim$ spatial infinity.

In-going into BH at horizon:

$$\tilde{R}_{\omega,m}(z) = z^\alpha {}_2F_1(a, b; c; z), \quad \alpha = i\sqrt{A_1},$$

$$a + b = 2\alpha, \quad ab = -A_1 - B_1, \quad c = 1 + 2\alpha. \quad \text{At horizon, } z = 0,$$

$$\text{Flux: } \mathcal{F} = 2\mathcal{A}_H(\omega - m\Omega_H). \quad (31)$$

BTZ $\sim AdS_3$, hard to tell in-coming from out-going.

$$R_{\omega,m}^{\text{in}} = A_i \left(1 - i\frac{c}{r_2}\right), \quad R_{\omega,m}^{\text{out}} = A_o \left(1 + i\frac{c}{r_2}\right). \quad (32)$$

$$\Rightarrow \mathcal{F}_{\text{in}} = 8\pi c |A_i|^2; \quad \sigma = \frac{\mathcal{F}}{\mathcal{F}_{\text{in}}}. \quad (33)$$

$$\text{In low energy limit } \omega \rightarrow 0, \quad \sigma_{\text{abs}} \simeq \frac{\sigma^{m=0}}{\omega} = \frac{\mathcal{A}_H |\Gamma(a+1)\Gamma(b+1)|^2}{\pi c |\Gamma(a+b+1)|^2}. \quad (34)$$

Choose c so that $\sigma_{\text{abs}}|_{\omega=0} = \mathcal{A}_H$.

Decay rate:

$$\begin{aligned}
 \Gamma &= \frac{\sigma_{\text{abs}}}{e^{\omega/T_H} - 1} \\
 &= 4\pi^2 \omega^{-1} T_L T_R e^{-\omega/(2T_H)} \left| \Gamma \left(1 + i \frac{\omega}{4\pi T_L} \right) \Gamma \left(1 + i \frac{\omega}{4\pi T_R} \right) \right|^2. \quad (35)
 \end{aligned}$$

$$\begin{aligned}
 \text{From CFT, } \Gamma &= \int d\sigma^- e^{-i\omega(\sigma^- - i\epsilon)} \left[\frac{2T_L}{\sinh(2\pi T_L \sigma^-)} \right]^2 \\
 &\quad \times \int d\sigma^+ e^{-i\omega(\sigma^+ - i\epsilon)} \left[\frac{2T_R}{\sinh(2\pi T_R \sigma^+)} \right]^2. \quad (36)
 \end{aligned}$$

The result is identical to that obtained from the gravity side.

Warped AdS₃/CFT₂ Correspondence

Warped BTZ Black Hole

The warped BTZ black hole is a solution of TMG:

$$I_{TMG} = \frac{1}{16\pi G} \int_{\mathcal{M}} d^3x \sqrt{-g} (R + 2/\ell^2) + \frac{\ell}{96\pi G\nu} \int_{\mathcal{M}} d^3x \sqrt{-g} \varepsilon^{\lambda\mu\nu} \Gamma_{\lambda\sigma}^r \left(\partial_\mu \Gamma_{r\nu}^\sigma + \frac{2}{3} \Gamma_{\mu\tau}^\sigma \Gamma_{\nu r}^\tau \right). \quad (37)$$

where $\varepsilon^{\tau\sigma u} = +1/\sqrt{-g}$, $\nu = \frac{m_g \ell}{3}$.

For $\nu = 1/3$, critical chiral gravity theory.

$$\text{EOM: } G_{\mu\nu} - \frac{1}{\ell^2} g_{\mu\nu} + \frac{\ell}{3\nu} \varepsilon_\mu^{\alpha\beta} \nabla_\alpha \left(R_{\beta\nu} - \frac{1}{4} g_{\beta\nu} R \right) = 0. \quad (38)$$

$$ds^2 = -N^2(r)dt^2 + \ell^2 R^2(r)[d\theta + N_\phi(r)dt]^2 + \frac{\ell^4 dr^2}{4R^2(r)N^2(r)}, \quad (39)$$

$$R^2(r) \equiv \frac{r}{4} \left[3(\nu^2 - 1)r + (\nu^2 + 3)(r_+ + r_-) - 4\nu\sqrt{r_+r_-(\nu^2 + 3)} \right],$$

$$N^2(r) \equiv \frac{\ell^2(\nu^2 + 3)(r - r_+)(r - r_-)}{4R^2}, \quad N_\theta(r) \equiv \frac{2\nu r - \sqrt{r_+r_-(\nu^2 + 3)}}{2R^2}.$$

$$T_H = \frac{\nu^2 + 3}{4\pi} \left\{ \frac{r_+ - r_-}{2\nu r_+ - \sqrt{(\nu^2 + 3)r_+r_-}} \right\}, \quad \mathcal{A}_H = \pi \left\{ 2\nu r_+ - \sqrt{r_+r_-(\nu^2 + 3)} \right\}.$$

$\nu > 1$, stretched AdS_3 , $\nu = 1$, BTZ limit,

$\nu < 1$, squashed AdS_3 , closed time-like curve.

t, r, θ different from the BTZ case.

Warped AdS_3 (locally equivalent to warped BTZ):

$$ds^2 = \frac{\ell^2}{\nu^2 + 3} \left[-\cosh^2 \sigma d\tau^2 + d\sigma^2 + \frac{4\nu^2}{\nu^2 + 3} (du + \sinh \sigma d\tau)^2 \right]. \quad (40)$$

Greybody Factor in Warped BTZ

Consider scalar Φ with mass μ . Let $\Phi = e^{-i\omega t + im\theta} \phi(r)$,

$$\frac{d^2\phi(r)}{dr^2} + \frac{2r - r_+ - r_-}{(r - r_+)(r - r_-)} \frac{d\phi(r)}{dr} - \frac{(\alpha r^2 + \beta r + \gamma)}{(r - r_+)^2(r - r_-)^2} \phi = 0, \quad (41)$$

$$\alpha = -\frac{3\omega^2(\nu^2 - 1)}{(\nu^2 + 3)^2} + \frac{\mu^2\ell^2}{\nu^2 + 3}, \quad \gamma = -\frac{4m \left[m - \omega\sqrt{r_+r_-(\nu^2 + 3)} \right]}{(\nu^2 + 3)^2} + \frac{\mu^2\ell^2 r_+ r_-}{\nu^2 + 3},$$

$$\beta = -\frac{\omega^2(\nu^2 + 3)(r_+ + r_-) - 4\nu \left[\omega^2\sqrt{r_+r_-(\nu^2 + 3)} - 2m\omega \right]}{(\nu^2 + 3)^2} + \frac{\mu^2\ell^2(r_+ + r_-)}{\nu^2 + 3}.$$

$z = \frac{r - r_+}{r - r_-}$, $z = 0 \sim$ horizon, $z = 1 \sim$ spatial infinity.

$$z(1 - z)\tilde{\phi}''(z) + (1 - z)\tilde{\phi}'(z) + \left[\frac{A}{z} + \frac{B}{1 - z} + C \right] \tilde{\phi}(z) = 0; \quad (42)$$

$$A = \frac{4(\omega\Omega_+^{-1} + m)^2}{(r_+ - r_-)^2(\nu^2 + 3)^2}, \quad B = -\alpha, \quad C = -\frac{4(\omega\Omega_-^{-1} + m)^2}{(r_+ - r_-)^2(\nu^2 + 3)^2}.$$

$$\Omega_+^{-1} = \nu r_+ - \frac{\sqrt{r_+ r_- (\nu^2 + 3)}}{2}, \quad \Omega_-^{-1} = \nu r_- - \frac{\sqrt{r_+ r_- (\nu^2 + 3)}}{2}.$$

$$\phi(z) = z^p (1-z)^q u(z), \quad \text{with } p = -i\sqrt{A}, \quad q = \frac{1}{2} (1 - \sqrt{1 + 4\alpha}),$$

$$\Rightarrow z(1-z)u''(z) + \{c - (a+b+1)z\}u'(z) - abu(z) = 0, \quad (43)$$

$$\text{with } a = p + q + \sqrt{C}, \quad b = p + q - \sqrt{C}, \quad c = 2p + 1.$$

$$\text{In-going at horizon: } u(z) = z^p (1-z)^q {}_2F_1(a, b; c; z). \quad (44)$$

$$\text{At } r \rightarrow \infty: \phi \simeq A_{in} \left(r^{-h_-^*} - \frac{i\eta}{\pi} r^{-h_+^*} \right) + A_{out} \left(r^{-h_-^*} + \frac{i\eta}{\pi} r^{-h_+^*} \right). \quad (45)$$

$$h_{\pm}^* = \frac{1}{2}(1 \pm \Delta^*), \quad \Delta^* = \sqrt{1 + 4\alpha} = \sqrt{1 - \frac{12\omega^2(\nu^2 - 1)}{(\nu^2 + 3)^2} + \frac{4\mu^2 \ell^2}{\nu^2 + 3}}.$$

η is chosen so that $\sigma_{abs}|_{\omega=0} = \mathcal{A}_H$.

Retaining ω dependence,

$$\sigma_{abs} \propto \left| \frac{\Gamma \left[\frac{1}{2}(1 + \sqrt{1 + 4\alpha}) - i\frac{\omega}{4\pi\tilde{T}_R} \right] \Gamma \left[\frac{1}{2}(1 + \sqrt{1 + 4\alpha}) - i\frac{\omega}{4\pi\tilde{T}_L} \right]}{\Gamma \left[1 - i\frac{\omega}{2\pi T_H} \right] \Gamma \left[\sqrt{1 + 4\alpha} \right]} \right|^2, \quad (46)$$

$$\tilde{T}_L = \frac{\nu^2 + 3}{8\pi\nu}, \quad \tilde{T}_R = \frac{(\nu^2 + 3)(r_+ - r_-)}{8\pi \left[\nu(r_+ + r_-) - \sqrt{r_+ r_-}(\nu^2 + 3) \right]}. \quad (47)$$

$$2/T_H = 1/\tilde{T}_L + 1/\tilde{T}_R. \quad (48)$$

Waped AdS_3 , $\Phi = e^{-i\omega_g\tau + iku} \phi_g(\sigma)$:

$$\frac{d^2\phi_g}{d\sigma^2} + \tanh\sigma \frac{d\phi_g}{d\sigma} + \left[(\omega_g \operatorname{sech}\sigma + k \tanh\sigma)^2 - \frac{(\nu^2 + 3)k^2}{4\nu^2} - \frac{\mu^2\ell^2}{\nu^2 + 3} \right] \phi_g = 0. \quad (49)$$

Define $\phi_g = z_g^{(\omega_g+ik)/2} (1-z_g)^{(\omega_g-ik)/2} f(z_g)$, $z_g \equiv \frac{1+i\sinh\sigma}{2}$.

$$z_g(1-z_g)f'' + [c_g - (1+a_g+b_g)z_g]f' - a_gb_gf = 0 \quad (50)$$

$$a_gb_g = \omega_g(\omega_g+1) + \frac{3k^2(\nu^2-1)}{4\nu^2} - \frac{\mu^2\ell^2}{\nu^2+3}, \quad a_g+b_g = 1+2\omega_g,$$

$$c_g = 1 + \omega_g + ik.$$

$$\text{For } z_g \rightarrow \infty, \phi_g(z_g) \rightarrow C_+ z_g^{h_+} + C_- z_g^{h_-}, \quad (51)$$

$$\text{with } h_{\pm} = \frac{1}{2}(1 \pm \Delta), \quad \Delta \equiv \sqrt{1 - \frac{3k^2(\nu^2-1)}{\nu^2} + \frac{4\mu^2\ell^2}{\nu^2+3}}.$$

Boundary of warped AdS_3 : $\sigma \rightarrow \infty$.

$x = e^{-\sigma}$ and $t = \tau/2$, local patch:

$$ds^2 \rightarrow \frac{\ell^2}{(\nu^2+3)x^2} \left[\frac{3(\nu^2-1)}{\nu^2+3} dt^2 + dx^2 + \frac{4\nu^2}{\nu^2+3} x^2 du^2 + \frac{8\nu^2}{\nu^2+3} x dt du \right]. \quad (52)$$

On the boundary,

$$\left(\frac{1}{\sqrt{-g}} \partial_x \sqrt{-g} g^{xx} \partial_x + \partial_u g^{uu} \partial_u - m^2 \right) K(x, u) = 0. \quad (53)$$

$$K(x, u) = e^{iku} x^{h_+} \Rightarrow K(x, u, t) = e^{iku} \left(\frac{x}{|x^2 - t^2|} \right)^{h_+}. \quad (54)$$

$$\text{Bulk field: } \phi(x, t, u) \sim \int dt' du' K_b(x, t, u, t', u') \phi_0(t', u'). \quad (55)$$

$$\text{In the limit } x \rightarrow 0, \quad \frac{\partial}{\partial x} \phi(x, t, u) \sim x^{h_+ - 1} \int dt' du' \frac{e^{ik(u-u')} \phi_0(t', u')}{|t - t'|^{2h_+}}. \quad (56)$$

$$\begin{aligned} S_{eff} &= \lim_{x \rightarrow 0} \left\{ -\frac{1}{2} \int dt du \sqrt{-g} g^{xx} \phi \partial_x \phi \right\} \\ &\sim \frac{1}{2} \int dt du dt' du' \frac{e^{ik(u-u')}}{|t - t'|^{2h_+}} \phi_0(t, u) \phi_0(t', u'). \end{aligned} \quad (57)$$

From the AdS/CFT dictionary: $e^{-S_{eff}(\phi)} = \left\langle e^{\int \phi_0 \mathcal{O}} \right\rangle$. (58)

Correlator: $\left\langle \mathcal{O}_*(u_+, u_-) \mathcal{O}_*(0, 0) \right\rangle^L \sim \frac{e^{i\omega u_+}}{|u_-|^{2h_+^*}}$, (59)

$\left\langle \mathcal{O}_*(u_+, u_-) \mathcal{O}_*(0, 0) \right\rangle^R \sim \frac{e^{i\omega u_-}}{|u_+|^{2h_+^*}}$, $u_{\pm} \equiv u \pm t/2c$. (60)

Assigning \tilde{T}_L and \tilde{T}_R to the left and right sectors.

$\left\langle \mathcal{O}_*(u_+, u_-) \mathcal{O}_*(0, 0) \right\rangle_T \sim (2h_+ - 1) e^{i\omega(u_+ + u_-)} \left[\frac{\pi \tilde{T}_R}{\sinh(\pi \tilde{T}_R u_+)} \right]^{2h_+^*} \left[\frac{\pi \tilde{T}_L}{\sinh(\pi \tilde{T}_L u_-)} \right]^{2h_+^*}$. (61)

$\mathcal{O}_*(x_+, x_-)$ with conformal dimension $h_+^* \equiv h_+ \Big|_{k=k_* \equiv \frac{2\nu}{\nu^2+3}\omega}$,

$\sigma_{abs} \sim \int dx_+ dx_- e^{-i\omega(x_+ + x_-)} \left\langle \mathcal{O}_*(x_+, x_-) \mathcal{O}_*(0, 0) \right\rangle_T$. (62)

Discussions

- There is indeed a warped $\text{AdS}_3/\text{CFT}_2$ correspondence.

- Gravity side:

$$\sigma_{abs} \propto \left| \frac{\Gamma \left[\frac{1}{2}(1 + \sqrt{1 + 4\alpha}) - i \frac{\omega}{4\pi \tilde{T}_R} \right] \Gamma \left[\frac{1}{2}(1 + \sqrt{1 + 4\alpha}) - i \frac{\omega}{4\pi \tilde{T}_L} \right]}{\Gamma \left[1 - i \frac{\omega}{2\pi T_H} \right] \Gamma \left[\sqrt{1 + 4\alpha} \right]} \right|^2,$$

$$\tilde{T}_L = \frac{\nu^2 + 3}{8\pi\nu}, \quad \tilde{T}_R = \frac{(\nu^2 + 3)(r_+ - r_-)}{8\pi \left[\nu(r_+ + r_-) - \sqrt{r_+ r_-}(\nu^2 + 3) \right]}, \quad 2/T_H = 1/\tilde{T}_L + 1/\tilde{T}_R.$$

- CFT side: $\left\langle \mathcal{O}_*(u_+, u_-) \mathcal{O}_*(0, 0) \right\rangle^L \sim \frac{e^{i\omega u_+}}{|u_-|^{2h_+^*}},$
 $\left\langle \mathcal{O}_*(u_+, u_-) \mathcal{O}_*(0, 0) \right\rangle^R \sim \frac{e^{i\omega u_-}}{|u_+|^{2h_+^*}}.$

Assigning \tilde{T}_L and \tilde{T}_R to the left and right sectors.

$$\left\langle \mathcal{O}_*(u_+, u_-) \mathcal{O}_*(0, 0) \right\rangle_T \sim (2h_+ - 1) e^{i\omega(u_+ + u_-)} \left[\frac{\pi \tilde{T}_R}{\sinh(\pi \tilde{T}_R u_+)} \right]^{2h_+^*} \left[\frac{\pi \tilde{T}_L}{\sinh(\pi \tilde{T}_L u_-)} \right]^{2h_+^*}.$$

$$\mathcal{O}_*(x_+, x_-) \sim \mathcal{O}_*(u_+, u_-) \text{ with } h_+^* \equiv h_+ \Big|_{k=k_* \equiv \frac{2\nu}{\nu^2+3}\omega},$$

$$\sigma_{\text{abs}} \sim \int dx_+ dx_- e^{-i\omega(x_+ + x_-)} \left\langle \mathcal{O}_*(x_+, x_-) \mathcal{O}_*(0, 0) \right\rangle_T,$$

$$\propto \left| \frac{\Gamma \left[\frac{1}{2}(1 + \sqrt{1 + 4\alpha}) - i \frac{\omega}{4\pi \tilde{T}_R} \right] \Gamma \left[\frac{1}{2}(1 + \sqrt{1 + 4\alpha}) - i \frac{\omega}{4\pi \tilde{T}_L} \right]}{\Gamma \left[1 - i \frac{\omega}{2\pi T_H} \right] \Gamma \left[\sqrt{1 + 4\alpha} \right]} \right|^2.$$

- $\nu > 1$, superradiant modes appear when

$$\omega^2 > \frac{(\nu^2 + 3)^2}{12(\nu^2 - 1)} + \frac{\nu^2 + 3}{3(\nu^2 - 1)}\mu^2\ell^2.$$

Conformal weight h_{\pm}^* becomes complex. $\omega \sim$ angular velocity.

The *effective* mass of scalar is below the B-F bound

$$\mu_{eff}^2 \equiv \frac{4\mu^2}{\nu^2 + 3} - \frac{12(\nu^2 - 1)\omega^2}{(\nu^2 + 3)^2\ell^2} < -\frac{1}{\ell^2}.$$

- Tortoise coordinate r^* : $\phi^*(r^*) \equiv z(r)\phi(r)$, $r^* = f(r)$.

Choose $f(r), z(r)$ s.t. $\left[-\frac{d^2}{dr^{*2}} - \omega^2 + U^*(r^*) \right] \phi(r^*) = 0.$

In spatial infinity, the effective potential

$$\lim_{r \rightarrow \infty} U^*(r) \rightarrow U_{\infty}^* = \frac{(\nu^2 + 3)(\nu^2 + 3 + 4\mu^2\ell^2)}{12(\nu^2 - 1)}.$$

Superradiance: $\omega^2 - U_{\infty}^* > 0.$

- Comparing the poles of bulk QNM and boundary retarded Green function.
- Dual theory of a rotating background.