

Generic dark matter signature for gamma-ray telescopes

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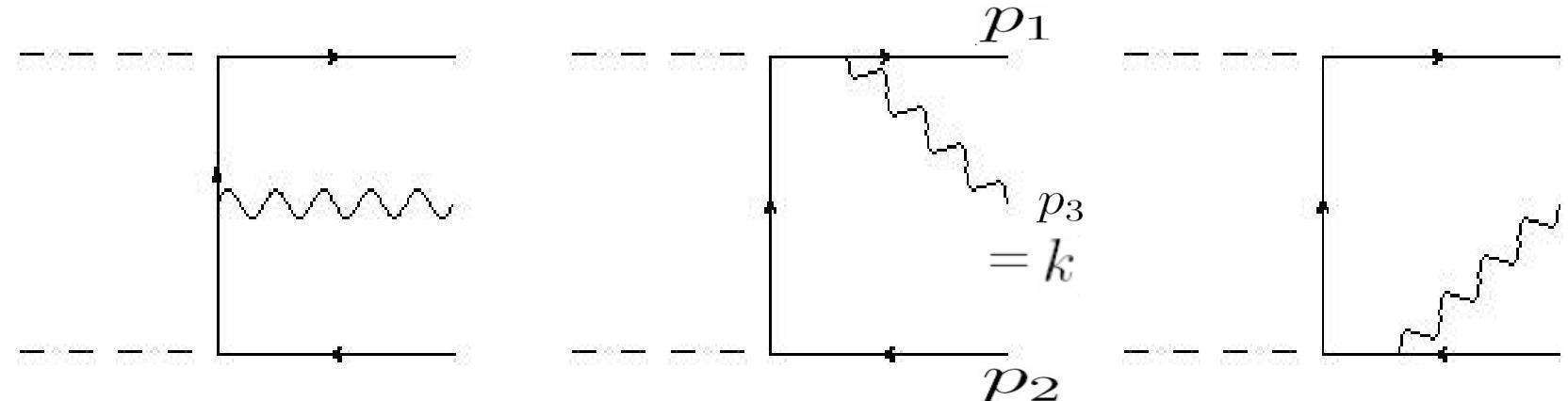
arXiv:0906.3009

In collaboration with V. Barger, Y. Gao, and D. Mafatia

Flores, Olive, Rudaz PLB (1989)

Bringmann, Bergstrom, Edsjo, arXiv: 0710.3169

$$\phi + \phi \rightarrow e(p_1) + \bar{e}(p_2) + \gamma(p_3)$$



$$\ell^T=(\nu,e^-)_L \qquad \qquad L_{L,R}^T=(N^0,E^-)_{L,R} \qquad {\cal L}\supset y\overline{\ell_L}L_R\phi$$

$${\cal M}_1=\frac{y^2\bar u_L(p_1)e\gamma_\mu(\not{\hbox{\kern-2.3pt p}}_1+\not{\hbox{\kern-2.3pt k}})[\frac{1}{2}(\not{\hbox{\kern-2.3pt p}}_1-\not{\hbox{\kern-2.3pt p}}_2+\not{\hbox{\kern-2.3pt k}})+m_E]v_L(p_2)}{(p_1+k)^2[\frac{1}{2^2}(p_1-p_2+k)^2-m_E^2]}\times 2$$

$$\phi + \phi \rightarrow e(p_1) + \bar{e}(p_2) + \gamma(p_3)$$

$$\mathcal{M}_1 = \frac{y^2 \bar{u}_L(p_1) e \gamma_\mu (\not{p}_1 + \not{k}) [\frac{1}{2}(\not{p}_1 - \not{p}_2 + \not{k}) + m_E] v_L(p_2)}{(p_1 + k)^2 [\frac{1}{2^2}(p_1 - p_2 + k)^2 - m_E^2]} \times 2$$

$$\mathcal{M}_2 = \frac{y^2 \bar{u}_L(p_1) [\frac{1}{2}(\not{p}_1 - \not{p}_2 - \not{k}) + m_E] (-\not{p}_2 - \not{k}) e \gamma_\mu v_L(p_2)}{[\frac{1}{2^2}(p_1 - p_2 - k)^2 - m_E^2] (p_2 + k)^2} \times 2$$

$$\mathcal{M}_3 = \frac{y^2 \bar{u}_L(p_1) [\frac{1}{2}(\not{p}_1 - \not{p}_2 - \not{k}) + m_E] e \gamma_\mu [\frac{1}{2}(\not{p}_1 - \not{p}_2 + \not{k}) + m_E] v_L(p_2)}{[\frac{1}{2^2}(p_1 - p_2 - k)^2 - m_E^2] [\frac{1}{2^2}(p_1 - p_2 + k)^2 - m_E^2]} \times 2$$

$$\not{p}_2 \gamma_\mu \not{p}_1 = -2(p_1 \cdot p_2) \gamma_\mu$$

$$\mathcal{M} = -2 \times \frac{4y^2 e \bar{u}_L(p_1) (\not{p}_2 \gamma_\mu \not{k} + \not{k} \gamma_\mu \not{p}_1) v_L(p_2)}{[(p_1 - p_2 + k)^2 - 4m_E^2][(p_1 - p_2 - k)^2 - 4m_E^2]}.$$

$$-\text{Tr}(\not{p}_1\not{p}_2\gamma_\mu\not{k}\not{p}_2\not{k}\not{\gamma}^\mu\not{p}_2)=4(2p_2\cdot k)^2(2p_1\cdot p_2)$$

$$\mathfrak{z}_{\mathrm{c}}$$

$$\begin{aligned}v_{\text{rel}} \frac{d\sigma}{dx_3 dx_1} &= \frac{1}{4\pi^3} \frac{(y^2 e)^2}{m_\phi^2} \frac{[(1-x_1)^2 + (1-x_2)^2](1-x_3)}{(1-2x_1-r)^2(1-2x_2-r)^2} \\r &= m_E^2/m_\phi^2 \quad x_i = E_i/m_\phi \quad x_1+x_2+x_3 = 2 \\z &= x_3 \\v_{\text{rel}} \frac{d\sigma}{dz} &= \frac{(y^2 e)^2/m_\phi^2}{32\pi^3} \frac{1-z}{(1+r-z)^2} \left(2z \frac{z^2 + (1+r-z)^2}{(1+r)(1+r-2z)} \right. \\&\quad \left. - \frac{(1+r)(1+r-2z)}{1+r-z} \ln \frac{1+r}{1+r-2z} \right)\end{aligned}$$

$$\frac{e}{\Lambda_L^3}\Phi F^{\mu\nu}\partial_\nu \big(\bar{\psi}_{eL}\gamma_\mu\psi_{eL}\big)$$

$${\cal M}=\tfrac{e}{\Lambda_L^3}(k^\mu\epsilon^\nu-k^\nu\epsilon^\mu)(p_1+p_2)_\nu\bar u_L(p_1)\gamma_\mu v_L(p_2)$$

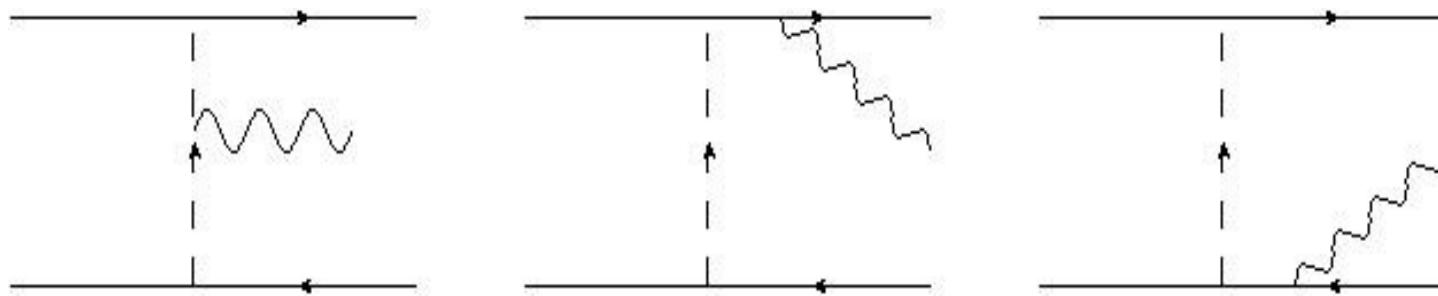
$${\cal M}=\tfrac{e}{2\Lambda_L^3}\bar u_L(p_1)({\not p}_2\not\epsilon\not k+\not k\not\epsilon\not p_1)v_L(p_2)$$

$$\sum |{\cal M}|^2 = \left(\frac{e}{2\Lambda_L^3}\right)^2 2[(2p_1\cdot p_2)(2p_1\cdot k)^2 + (2p_1\cdot p_2)(2p_2\cdot k)^2]$$

$$2p_1\cdot p_2=M_\Phi^2(1-x_3),\,2k\cdot p_1=M_\Phi^2(1-x_2),\,2k\cdot p_2=M_\Phi^2(1-x_1)$$

$$\frac{1}{\Gamma}\frac{d\Gamma}{dx_3}=20(1-x_3)x_3^3\qquad \frac{1}{\Gamma}\frac{d\Gamma}{dx_2}=15x_2^2-30x_2^3+\tfrac{35}{2}x_2^4$$

SUSY-like DM pair $\chi\chi \rightarrow \ell\bar{\ell}\gamma$



$$i\mathcal{M}_{\text{tree}} = (i)^5 e g_{L,R}^2 T^c \bar{u}(p_1) \cdots [u(K/2)\bar{v}(K/2) - v(K/2)\bar{(K/2})] \cdots v(p_2)$$

$$u(K/2)\bar{v}(K/2) - v(K/2)\bar{u}(K/2) = (m_\chi + K/2)\gamma_5$$

$$A_1 = \frac{\not{\epsilon}^* \frac{1}{\not{p}_1+\not{k}} (K/2)\gamma_5}{(K/2-p_2)^2 - \widetilde{M}^2} \rightarrow \frac{\not{\epsilon}^* \gamma_5/2}{(K/2-p_2)^2 - \widetilde{M}^2}$$

$$A_2 = \frac{(K/2)\gamma_5 \frac{1}{-\not{p}_2-\not{k}} \not{\epsilon}^*}{(-K/2+p_1)^2 - \widetilde{M}^2} \rightarrow -\frac{\not{\epsilon}^* \gamma_5/2}{(-K/2+p_1)^2 - \widetilde{M}^2}$$

$$A_3 = \frac{K\gamma_5/2 \quad (p_1-p_2) \cdot \epsilon^*}{[(-K/2+p_1)^2 - \widetilde{M}^2][(K/2-p_2)^2 - \widetilde{M}^2]}$$

$$A = \frac{1}{2} \frac{(p_2-p_1) \cdot K \not{\epsilon}^* + K(p_1-p_2) \cdot \epsilon^*}{[(-K/2+p_1)^2 - \widetilde{M}^2][(K/2-p_2)^2 - \widetilde{M}^2]} \gamma_5$$

$$\mathcal{M}_{\text{tree}} = \frac{1}{2} e g_{L,R}^2 \bar{u}(p_1) \frac{(p_2-p_1) \cdot k \not{\epsilon}^* + k(p_1-p_2) \cdot \epsilon^*}{[(-K/2+p_1)^2 - \widetilde{M}^2][(K/2-p_2)^2 - \widetilde{M}^2]} \gamma_5 v(p_2)$$

$$\bar{u}(p_1)(p_2 \cdot k \not{\epsilon}^* - k p_2 \cdot \epsilon^*)v(p_2) = -\tfrac{1}{2}\bar{u}(p_1)\not{p}_2\not{\epsilon}^*kv(p_2)$$

$$\bar{u}(p_1)(p_1 \cdot k \not{\epsilon}^* - k p_1 \cdot \epsilon^*)v(p_2) = -\tfrac{1}{2}\bar{u}(p_1)\not{k}\not{\epsilon}^*\not{p}_1v(p_2)$$

$$\mathcal{M}_{\text{tree}} = \frac{1}{4} e g_{L,R}^2 \bar{u}(p_1) \frac{k \not{\epsilon}^* \not{p}_1 - \not{p}_2 \not{\epsilon}^* k}{[(-K/2+p_1)^2 - \widetilde{M}^2][(K/2-p_2)^2 - \widetilde{M}^2]} \gamma_5 v(p_2)$$

Two effective operators

$$\bar{u}(p_1) \not{p}_2 \not{\epsilon}^* \not{k} v(p_2) = \bar{u}_1(2p_2 \cdot \epsilon^* \not{k} - 2p_2 \cdot k \not{\epsilon}^*)v(p_2) \longleftarrow 2\mathcal{O} \quad \mathcal{O} = F^{\mu\nu}\bar{\psi}\gamma_\nu\partial_\mu\psi$$

$$\bar{u}(p_1) \not{k} \not{\epsilon}^* \not{p}_1 v(p_2) = \bar{u}_1(2p_1 \cdot \epsilon^* \not{k} - 2p_1 \cdot k \not{\epsilon}^*)v_2 \longleftarrow 2\mathcal{O}^\dagger \quad \mathcal{O}^\dagger = F^{\mu\nu}(\partial_\mu\bar{\psi})\gamma_\nu\psi$$

Chisholm identity $\gamma^\alpha\gamma^\beta\gamma^\mu = g^{\alpha\beta}\gamma^\mu - g^{\alpha\mu}\gamma^\beta + g^{\beta\mu}\gamma^\alpha - i\epsilon^{\alpha\beta\mu\nu}\gamma_\nu\gamma_5$

$$\bar{u}(p_1) \not{p}_2 \not{\epsilon}^* \not{k} v(p_2) = \bar{u}_1(p_2 \cdot \epsilon^* \not{k} - p_2 \cdot k \not{\epsilon}^* + \underbrace{k \cdot \epsilon^* \not{p}_2}_{\rightarrow 0} - i\epsilon^{p_2, \epsilon^*, k, \mu}\gamma_\mu\gamma_5)v(p_2)$$

$$\bar{u}(p_1) \not{p}_2 \not{\epsilon}^* \not{k} v(p_2) = -2i\epsilon^{p_2, \epsilon^*, k, \mu}\bar{u}(p_1)\gamma_\mu\gamma_5v(p_2)$$

$$\bar{u}(p_1) \not{k} \not{\epsilon}^* \not{p}_1 v(p_2) = +2i\epsilon^{p_1, \epsilon^*, k, \mu}\bar{u}(p_1)\gamma_\mu\gamma_5v(p_2)$$

$$\mathcal{O} = F^{\mu\nu} \bar{\psi} \gamma_\nu \partial_\mu \psi = -i \tilde{F}^{\mu\nu} \bar{\psi} \gamma_\nu \gamma_5 \partial_\mu \psi$$

$$\mathcal{O}^\dagger = F^{\mu\nu} (\partial_\mu \bar{\psi}) \gamma_\nu \psi = i \tilde{F}^{\mu\nu} (\partial_\mu \bar{\psi}) \gamma_\nu \gamma_5 \psi$$

$$\tilde{F}^{\mu\nu} = \tfrac{1}{2} \epsilon^{\mu\nu\alpha\beta} F_{\alpha\beta} \quad \begin{matrix} \text{dual substitution} \\ \mathbf{E} \leftrightarrow \mathbf{B} \end{matrix}$$

$$\mathcal{O} + \mathcal{O}^\dagger = F^{\mu\nu} \partial_\mu (\bar{\psi} \gamma_\nu \gamma_5 \psi) \equiv \mathcal{F} \;,$$

$$i(\mathcal{O} - \mathcal{O}^\dagger) = \tilde{F}^{\mu\nu} \partial_\mu (\bar{\psi} \gamma_\nu \gamma_5 \psi) \equiv \tilde{\mathcal{F}}$$

LR-like model

non-linear Feynman gauge

$$\Gamma_{W'(q')_\lambda \rightarrow W(q)_\rho \gamma(k)_\mu} = e \left[g^{\lambda\rho} (q' + q)^\mu + g^{\rho\mu} (2k)^\lambda + g^{\lambda\mu} (-2k)^\rho \right]$$

$$\epsilon^*(\not{p}_1 + \not{p}_2) \not{k} - \not{k} (\not{p}_1 + \not{p}_2) \epsilon^*$$

$$\epsilon^* \not{p}_1 \not{k} - \not{k} \not{p}_1 \epsilon^* = + \not{k} \epsilon^* \not{p}_1$$

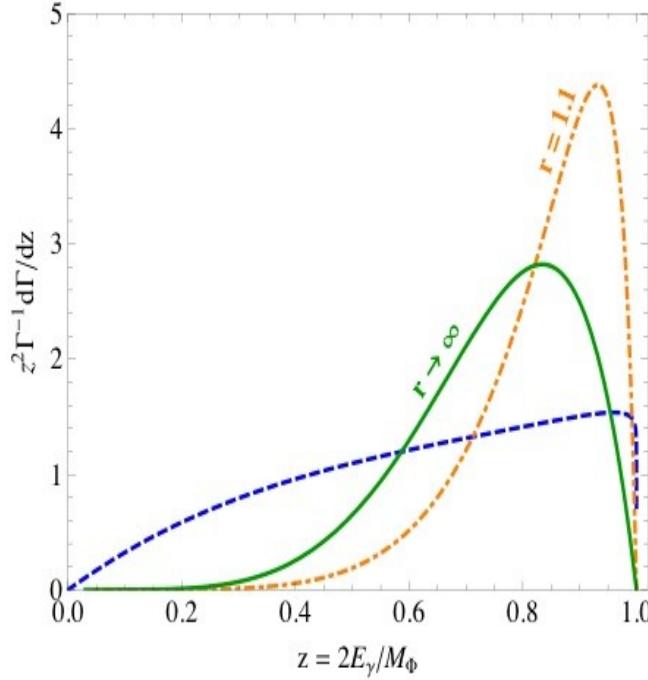
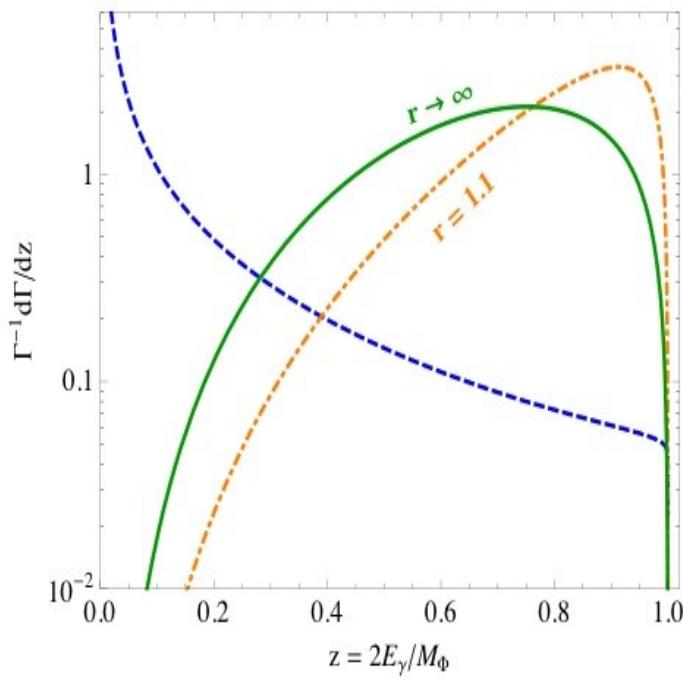
$$\epsilon^* \not{p}_2 \not{k} - \not{k} \not{p}_2 \epsilon^* = - \not{p}_2 \epsilon^* \not{k}$$

$$-2 + 4 + (m_N/m_{W'})^2$$

Model	1	2	3
DM	Majorana χ	Majorana N	Scalar ϕ
Exchange	Scalar S	Vector W'	Fermion E
Interaction \mathcal{L}_I	$g' S^\dagger \bar{\chi} e_R$	$g' W'_\mu^\dagger \bar{N} \gamma^\mu e_R$	$g' \bar{E} \phi e_R$

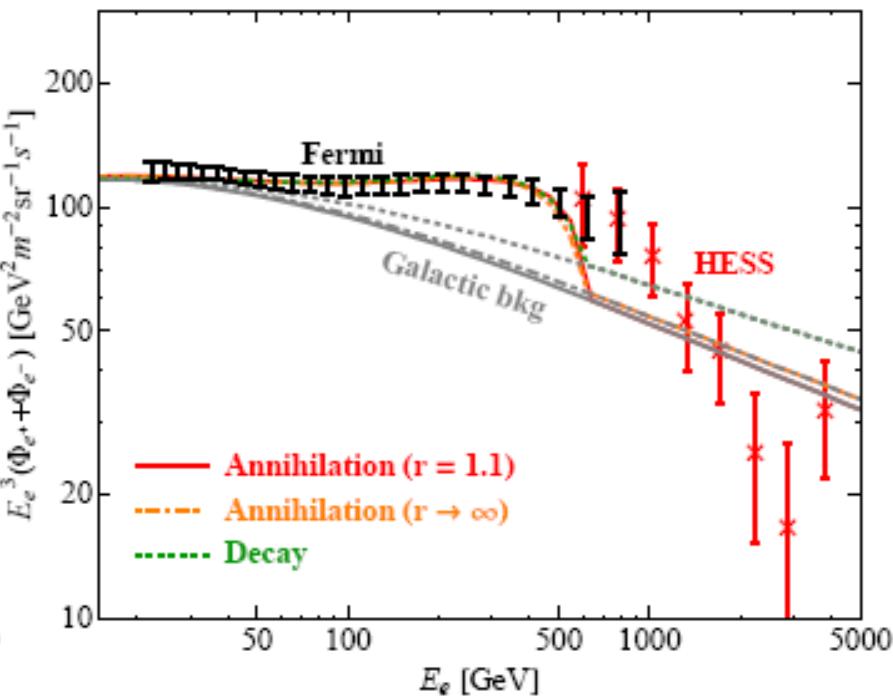
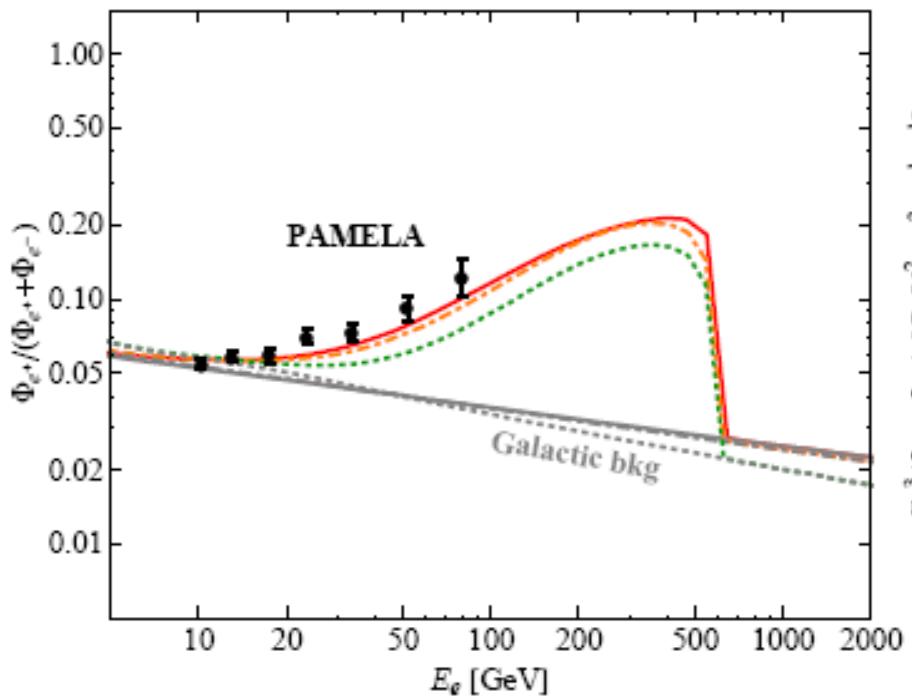
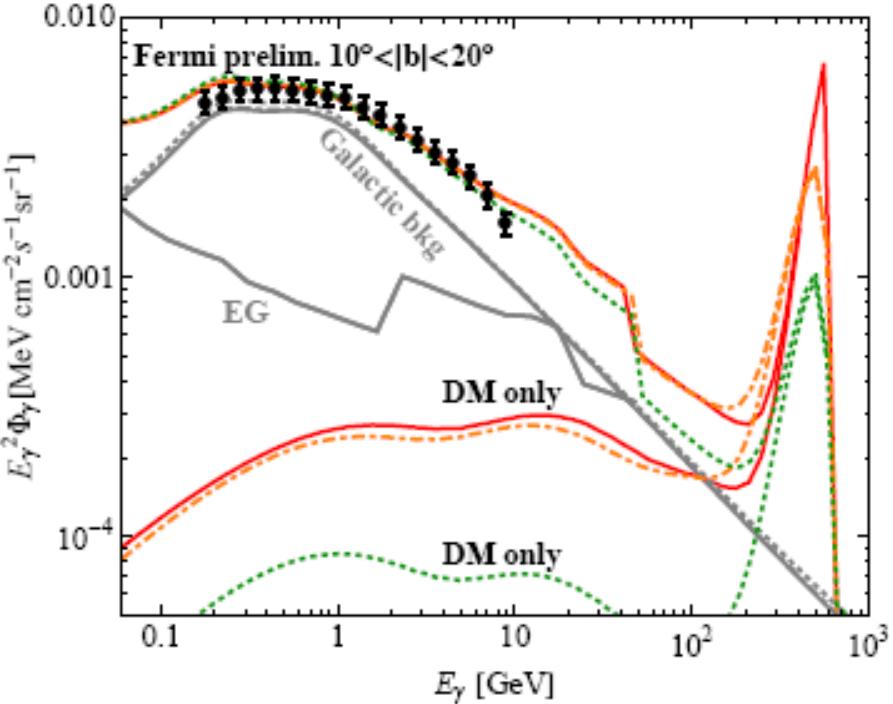
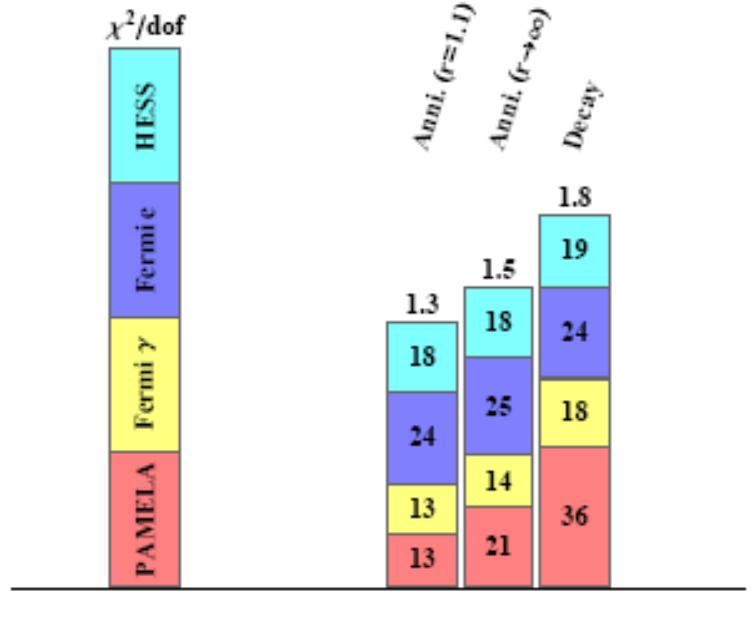
$$\mathcal{M} = -\frac{C}{2} \frac{g'^2 e \bar{u}_R(p_1) (\not{p}_2 \gamma_\mu \not{k} \mp \not{k} \gamma_\mu \not{p}_1) v_R(p_2)}{[\frac{1}{4}(p_1 - p_2 + k)^2 - m_X^2][\frac{1}{4}(p_1 - p_2 - k)^2 - m_X^2]}$$

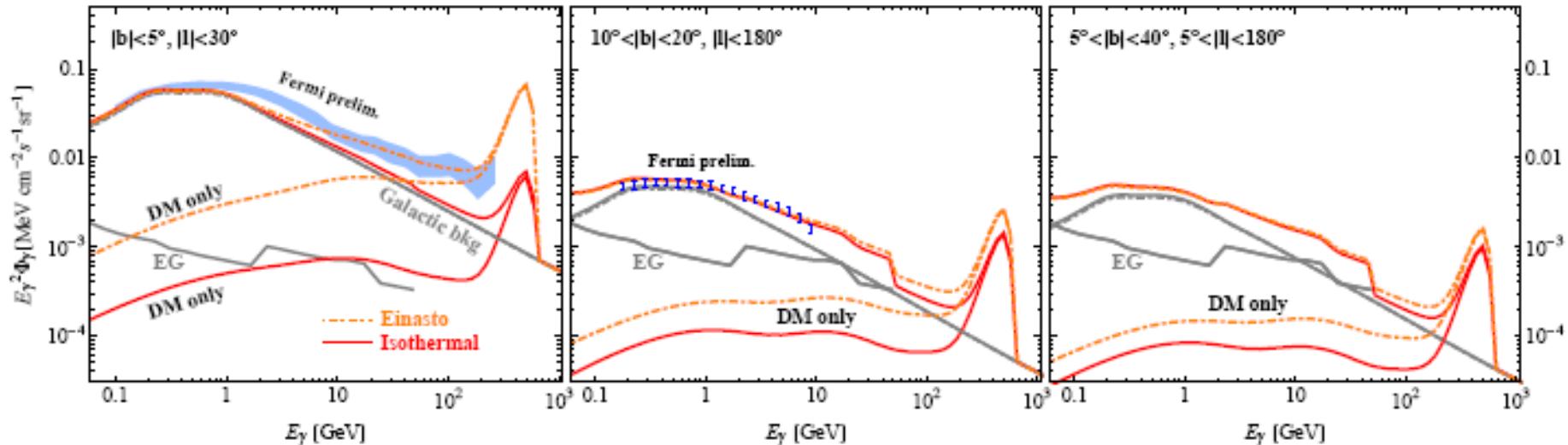
$$C = \begin{cases} \frac{i}{\sqrt{2}} & \text{for Model 1,} \\ \frac{i}{\sqrt{2}} \left(2 + \frac{m_D^2}{m_X^2}\right) & \text{for Model 2,} \\ 1 & \text{for Model 3.} \end{cases}$$



$$\frac{1}{\Gamma} \frac{d\Gamma}{dz} = 20(1-z)z^3,$$

$$\frac{1}{\Gamma} \frac{d\Gamma}{dx_2} = 5(3 - 6x_2 + \frac{7}{2}x_2^2)x_2^2$$





CANGAROO Collab. Of Australia and Nippon for a Gamma Ray Observatory in the Outback

MAGIC Major Atmospheric Gamma-ray Imaging Cerenkov Telescope

CTA Cerenkov Telescope Array

AGIS Advanced Gamma Ray Imaging System

VERITAS Very Energetic Radiation Imaging Telescope Array System

Summary

Unique form of photon/lepton energy spectrum for chirality preserving interaction

Prompt photon / lepton are good messenger of DM