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Dark matter mass and distribution by directional detection

... talk is based on arXiv:1707.05523; KN, T. Ikeda, R. Yakabe, T. Naka, K. Miuchi

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Outline

Introduction
DM Distribution
Numerical Results
Summary

Dark Matter



Stable

- Electrically neutral, weakly interacting
- Non-relativistic
- No candidate in SM

WIMP? Axion (-like particle)? or …?

Direct Detection





Constraints by Direct Detection

Constraint for mass - cross section
Cross section < 10⁻⁴⁶ cm² (SI), 10⁻³⁸cm² (SD)



XENON 1T (2018)



Schumann arXiv:1501.01200

Directional Direct Detection

 Ordinally direct detection Recoil energy E_R
Directional detection Recoil energy E_R + Direction of nuclear recoil (+ time)
Typical target

l ypical target CF₄, CS₂, CHF₃ (gas detector)

> Ag, Br, C, N, O (solid detector)



DRIFT Phys. of the Dark Universe 9–10 (2015)



NEWAGE PTEP (2015), 4

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Velocity distribution 1

Maxwell distribution $f(v) = \frac{1}{(\pi v_0^2)^{3/2}} e^{-(v+v_E)^2/v_0^2}$ $\frac{dR}{dE_R} = \frac{N_T \rho_0}{m_{\chi}} \int^{v_{\text{max}}} d\vec{v} f(\vec{v}) |\vec{v}| \frac{d\sigma(\vec{v})}{dE_R}$



commonly supposed in direct detections

isotropy is also supposed

How can we test it? cosmological observations directional detection

the Galaxy Velocity distribution 2 the Solar system

Some N-body simulations suggest anisotropy



Ling, Nezri, Athanassoula & Teyssier (2009) cf. Kuhlen et al. (2012), David R. Law (2009) ...

Numerical calculation

 $\overline{N(v_{0,\mathrm{ani.}})}$

$$f(v_{\phi}) = \frac{1-r}{N(v_{0,\text{iso.}})} \exp\left[-v^2/v_{0,\text{iso.}}^2\right] + \text{isotropic}$$

Tangential velocity - anisotropy parameter r - r=0.25 is suggested by simulation Goal: Discrimination - isotropic case (r=0) ---anisotropic case (r=0.3)



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 Monte Carlo simulation of scattering supposing f(v) Direction (scattering angle) + Recoil energy Elastic scattering M_{DM}=3M_N Target : F (light target) /Ag (heavy target)

Analysis

depends on resolutions of a detector

Energy resolution :OK Angular resolution :OK



E_R -cos θ energy-angular distribution

Energy resolution :NG Angular resolution :OK



cosθ angular histogram

Isotropic .5 **cos** θ r = 0.2 Strategy for discrimination EXP. r = 0.3 r = 0.4 Which one is the most similar to Exp.? ^{.5} cos⁰ θ r = 0.5 Statistical test to examine the similarity of distributions r = 0.7

✓ Chi-squared test

•••

- Likelihood analysis
- Kolmogorov–Smirnov test

Anisotropic

^{0.5} cos⁰ θ r = 0.8

r = 1

Strategy for discrimination 2



Supposing M_{DM} is known..._{Ethr=50keV (Ag)}

Ethr=50keV (Ag) M_{dm}=300GeV





Required event numbers to exclude isotropic case are 6×10³ (ER-cos) / 5×10³ (cos only) for target F 6×10⁴ (cos only) / 2×10⁴ (cos only) for target Ag.

Note: What if M_{DM} is not known?



cf. Samuel K. Lee and, Annika H.G. Peter arXiv:1202.5035

Summary

- Possibility to figure out DM mass and anisotropy of DM distribution is discussed.
- If DM mass is known by other searches, we can discuss the anisotropy once O(10³-10⁴) event is obtained in directional detection.
- Even if M_{DM} is not known, once both E_R and angular information are obtained we can give constraints for M_{DM} and distribution.