

Precessional Memory Effect

A New Cosmological Probe

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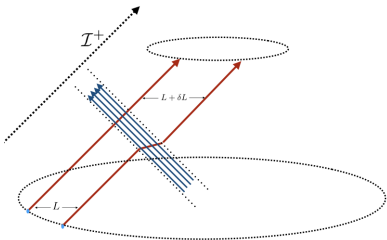


Motivation



Gravitational Wave Memory Effect

The permanent relative displacement of a pair of test mass particles upon the passage of gravitational wave is called **gravitational memory effect**. (Zel'dovich, Braginski, Thorne, Christodoulou...)



Different kinds

- Displacement Christodoulou91,Thorne92
- Electromagnetic memory Bieri 10
- Velocity Grishchuk 89
- Spin-memory Pasterski 15

Bigger Picture!

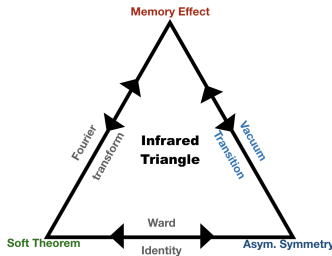


Figure: Strominger et. al.



Fundamentals of Displacement Memory

Geodesic Deviation Equation(GDE) of particles of trajectories $x^\mu(\tau)$ and $x^\mu(\tau) + \xi^\mu(\tau)$

$$\frac{D^2 \xi^\mu}{d\tau^2} = -R^\mu_{\alpha\lambda\beta} \frac{dx^\alpha}{d\tau} \xi^\lambda \frac{dx^\beta}{d\tau}. \quad (1)$$

Weak field approximation: $g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$, where $|h_{\mu\nu}| \ll 1$.

The transverse traceless(TT) gauge condition:

$$h_{0\mu} = 0, \quad h_{ij,j} = 0, \quad h^\mu_\mu = 0 \quad (2)$$

In TT gauge: $\frac{D}{d\tau} \rightarrow \frac{d}{dt}$ and $\frac{dx^\alpha}{d\tau} = (1, 0, 0, 0)$

The Riemann curvature tensor:

$$R_{i0j0} = -\frac{1}{2} h_{ij,00}^{TT} \quad (3)$$

For particles in laboratory frame of reference:

$$\frac{d^2 \xi^i}{d\tau^2} = \frac{1}{2} h_{ij,00}^{TT} \xi^j. \quad (4)$$

Displacement memory in the leading order:

$$\Delta \xi^i(t \rightarrow +\infty, t \rightarrow -\infty) = \frac{1}{2} \Delta h_{ij}^{TT} \xi^j(t_i) = \frac{1}{2} h_{+,x}(t \rightarrow +\infty) - h_{+,x}(t \rightarrow -\infty) \xi^j(t_i).$$



Formulation of Gravitational Thomas Precession



Thomas Precession in a Relativist's Perspective

Thomas Precession is due to the non-commutative nature of Lorentz groups.

$$\left(\frac{d\vec{S}}{dt}\right)_{\text{Non-Rot}} = \left(\frac{d\vec{S}}{dt}\right)_{\text{R}} + \vec{\omega}_{\text{T}} \times \vec{S}, \quad (6)$$

Where ω_{T} is the frequency of precession, it is given by

$$\omega_{\text{T}} = \frac{1}{2} \frac{\vec{a} \times \vec{v}}{c^2}$$

For an electron, the energy associated with Thomas Precession is

$$U_{\text{Non-Rot}} = U_{\text{R}} + \vec{S} \cdot \vec{\omega}_{\text{T}}$$

For a Bohr atom with $n = 1$, $r = r_0$,

$$\frac{|\vec{v}|}{c} \equiv \alpha = \frac{e^2}{4\pi\epsilon_0\hbar c}, \quad r_0 = \frac{\hbar}{m\alpha c}$$

$$|\vec{a}| = \frac{e^2}{4\pi\epsilon_0 r_0^2 m} = \frac{\hbar\alpha c}{mr_0^2}, \quad E_0 = \frac{1}{2} mc^2 \alpha^2$$

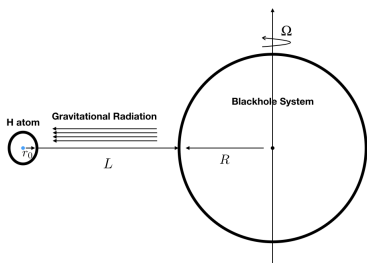
TP energy for Hydrogen atom is due to the Spin-Orbit coupling. This causes fine structure of hydrogen energy levels.

$$U_{\text{TP}} = \vec{S} \cdot \frac{1}{2} \frac{\vec{a}}{c} \times \frac{\vec{v}}{c}$$

$$\simeq \frac{|\vec{S}|}{\hbar} \alpha^2 E_0.$$



Gravitational Thomas Precession for a Toy Model



L : distance from the source to observer
 Ω is the angular frequency of the orbit.

Case: Binary system stars of mass M ,
 revolving in a circular orbital of radius R

Quadrupole moment: $I \sim MR^2$

Upon the passage of Gravitational waves
 there exists an acceleration:

$$a^j = \frac{d^2 \xi^j}{d\tau^2} = \frac{1}{2} h_{ij,00}^{TT} \xi^j.$$

Using quadrupole approximation :

$$a^j = \frac{G}{c^4} \Omega^4 \frac{I_{ij}^{TT}}{L} \xi^j.$$

This acceleration causes Thomas precession in Bohr atom. The energy of spin-orbit coupling is:

$$U_{GTP} \simeq \frac{|\vec{S}|}{2\hbar} \cdot \alpha^2 E_0 \left[\frac{\phi_s R}{L} \left(\frac{v_s}{c} \right)^4 \left(\frac{r_0}{R} \right)^2 \cdot \frac{1}{\alpha^4} \right]. \quad (8)$$

ϕ_s : dimensionless gravitational potential
 v_s : surface velocity of the source

$$\phi_s = \frac{GM}{c^2 R}, \quad v_s = \Omega R.$$



Gravitational Fine Structure and Memory Effect

- Gravitational waves cause Thomas Precession which is encoded in the fine structure of hydrogen. Compared with standard spin-orbit interaction energy:

$$\frac{U_{\text{GTP}}}{U_{\text{TP}}} \sim \phi_s \left(\frac{v_s}{c}\right)^4 \left(\frac{R}{L}\right) \left(\frac{r_0}{R}\right)^2 \cdot \frac{1}{\alpha^4}.$$

- Event's signature can be found on the hydrogen fine structure spectra, which is unique, hence termed it as **Precession Memory Effect(PME)**.

If ν is the frequency of the radiation without memory and $\nu(M)$ is with PME, then

$$\frac{\delta\nu(M)}{\nu} \simeq \left[\phi_s \left(\frac{v_s}{c}\right)^4 \left(\frac{r_0^2}{3 \times 10^3 \text{mL}}\right) \left(\frac{M_\odot}{M}\right) \cdot \frac{1}{\alpha^4} \right] \quad (9)$$



Few simple scenarios:

■ Pulsars:

Mass M_{\odot}
Radius $\sim 10\text{km}$

Distance between hydrogen atom and
pulsar $L \sim 100R$

$$\frac{U_{\text{GTP}}}{U_{\text{TP}}} \sim 10^{-31} \quad (10)$$

Too small to be observed!

■ Blackholes with $M \geq M_{\odot}$

For a Schwarzschild BH, the radius of the source is equal to Schwarzschild radius:

$$R = \frac{2GM}{c^2} = 3\text{km} \times \left(\frac{M}{M_{\odot}} \right). \quad (11)$$

Too small to be observed since $\left(\frac{r_0}{R} \right)^2$ is still suppressing the energy contribution!



Perfect candidates: Primordial Black Holes(PBHs)

Masses of PBHs $\geq M_{\odot}$ are not suitable.

Consider lower mass PBHs, say $\left(\frac{M}{M_{\odot}}\right) \sim 10^{-n}$. Then the radius of the source is

$$R \sim 3\text{km} \times 10^{-n} \sim 10^{-(n-3)}\text{m}$$

The suppressing factor: $\left(\frac{r_0}{R}\right)^2 \sim 10^{2n-27}$. For a source of $\phi_s \approx 1$ and the distance between source to atom is $L = 10^3 R$ the energy of GTP is

$$\frac{U_{\text{GTP}}}{U_{\text{TP}}} \sim 10^{-2} \times 10^{2(n-14)}. \quad (12)$$

To estimate the spectral splitting, quantum mechanical perturbation is used. To apply it, the mass range of PBH source is $10^{-14} \leq \left(\frac{M}{M_{\odot}}\right) < 10^{-13}$.

For sources in this range, the spectral line observed should have the frequency of the order of $\delta\nu \sim 10^7\text{Hz}$ (Radio Waves). As we go farther away the spectral line becomes redder.



As a Cosmological probe



Perfect candidates: Primordial Black Holes(PBHs)

Advantage: They come in wide range of masses from Planck mass to supermassive! Zeldovich67,Hawking74...

Why are they interesting?

- Prime dark matter candidates
- Candidates of Supermassive BHs in galaxy centers and AGNs
- Possible contribution to GRBs and cosmic rays
- Possible contribution to BBN synthesis



Interesting Recent Developments

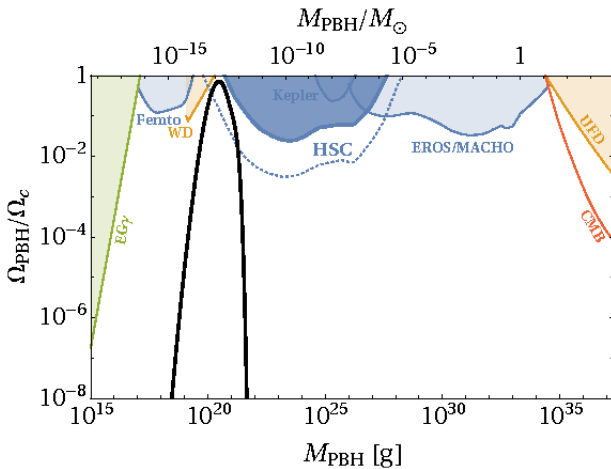


Figure: Inflationary PBHs as all dark matter (Inomata et. al. 2017)



Preliminaries: Mass and Abundances of PBHs

For simplicity, radiation dominant era is considered.

During the time of formation, the mass of PBH is proportional to Horizon mass.

$$\begin{aligned}
 M_{\text{PBH}} &= \gamma \frac{4\pi\rho}{3} H^{-3} \Big|_{\text{Formation}} = \frac{\gamma}{2G} H^{-1} \Big|_{\text{Formation}} \\
 &\simeq 10^{20} \text{g} \left(\frac{\gamma}{0.2} \right) \left(\frac{g_*}{106.75} \right)^{-\frac{1}{6}} \left(\frac{k}{7 \times 10^{20} \text{MPC}^{-1}} \right)^{-2} \quad (13)
 \end{aligned}$$

The abundance of PBHs are estimated using mass fraction defined as $\beta(M) := \frac{\rho_{\text{PBH}}(M)}{\rho}$.

Mass Fraction depends on the profile of density perturbations.

$$\beta = \gamma \int_{\delta_c}^1 P(\delta) d\delta \quad (14)$$

For a gaussian Distribution:

$$\begin{aligned}
 \beta(M) &= \gamma \int_{\delta_c}^1 \frac{d\delta}{\sqrt{2\pi\sigma_{\text{PBH}}^2(M)}} e^{-\frac{\delta^2}{2\sigma_{\text{PBH}}^2(M)}} \\
 &\simeq \frac{3\gamma}{\sqrt{2\pi}} \sigma_{\text{PBH}}(M) e^{-\frac{1}{18\sigma_{\text{PBH}}^2(M)}}. \quad (15)
 \end{aligned}$$



PBH Background in Early Universe

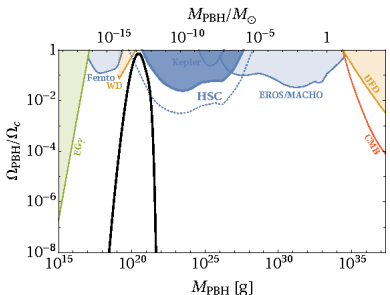


Figure: Inflationary PBHs as all dark matter (Inomata et. al. 2017)

Number density as monochromatic mass

$$\text{function: } n_{\text{PBH}}(M) = \frac{\beta(M)\rho_{\text{total}}}{M}$$

Cumulative number density

$$dN(M) = n_{\text{PBH}}(M)d\ln M$$

If P is the probability of mergers per unit time, event rate of PBH mergers of a certain mass at fixed z per unit time is

$$dE = dN(M)P$$

Total number of events:

$$E = \int \frac{\beta(M)\rho(t_i)}{M} P d\ln M$$

- For $30M_\odot$, $E \rightarrow 2\text{Gpc}^{-3}\text{yr}^{-1}$ (Sasaki16)
- $10M_\odot$, $E \rightarrow 5\text{Gpc}^{-3}\text{yr}^{-1}$
- $2M_\odot$, $E \rightarrow 2000\text{Gpc}^{-3}\text{yr}^{-1}$ (Chen16)



Our Proposal: As a new Cosmological Probe



Our Proposal

- After the end of radiation dominant era, light elements were formed. Lightest and the most abundant element during that epoch was hydrogen.
- If there was a large abundance of PBHs in the early universe, then the number of merger events should be large.
- Gravitational waves from these merger events would be stronger near the source, therefore it should leave strong Gravitational Thomas Precession Signatures in the hydrogen cloud in the vicinity.
- These local GTP effect should have a distinct $\frac{1}{L}$ profile.
- Assuming the abundance of PBHs of desired mass range and also with considerable event rate, we can map GTP fluctuations in the entire sky.
- Event rate could be used to constrain mass fraction and PBH contribution to dark matter.



THANK YOU.



A bit about calculations

Gravitational perturbations for the above case can be written as:

$$\begin{aligned} h_{ij,00}^{TT} &= \frac{G}{2c^4} \Omega^4 \frac{MR^2}{L} \xi^j \\ &= \phi_s \left(\frac{v_s}{c}\right)^4 \left(\frac{R}{L}\right) \frac{c^2}{R^2}. \end{aligned} \quad (16)$$

ϕ_s : dimensionless gravitational potential

v_s : surface velocity of the source

$$\phi_s = \frac{GM}{c^2 R}, \quad v_s = \Omega R. \quad (17)$$

Acceleration Approximation

$$a \sim |h_{ij,00}^{TT}| r_0$$

Estimation of Thomas Precession Energy:

$$\begin{aligned} \frac{a}{c} &\sim \phi_s \left(\frac{v_s}{c}\right)^4 \left(\frac{R}{L}\right) \left(\frac{r_0 c}{R}\right) \cdot \frac{E_0 r_0}{\alpha \hbar c} \\ &= \frac{\phi_s R}{L} \left(\frac{v_s}{c}\right)^4 \left(\frac{r_0}{R}\right)^2 \cdot \frac{E_0}{\alpha \hbar}. \end{aligned}$$

Energy:

$$\begin{aligned} U_{\text{GTP}} &\sim \frac{|\vec{S}|}{2\hbar} \phi_s \left(\frac{v_s}{c}\right)^4 \left(\frac{R}{L}\right) \left(\frac{r_0}{R}\right)^2 \cdot \frac{E_0}{\alpha^2} \\ &= \frac{|\vec{S}|}{2\hbar} \cdot \alpha^2 E_0 \left[\frac{\phi_s R}{L} \left(\frac{v_s}{c}\right)^4 \left(\frac{r_0}{R}\right)^2 \cdot \frac{1}{\alpha^4} \right]. \end{aligned}$$



Gravitational Thomas Precession for a Toy Model (contd...)

Energy:

$$\begin{aligned}
 U_{\text{GTP}} &\sim \frac{|\vec{S}|}{2\hbar} \phi_s \left(\frac{v_s}{c}\right)^4 \left(\frac{R}{L}\right) \left(\frac{r_0}{R}\right)^2 \cdot \frac{E_0}{\alpha^2} \\
 &= \frac{|\vec{S}|}{2\hbar} \cdot \alpha^2 E_0 \left[\frac{\phi_s R}{L} \left(\frac{v_s}{c}\right)^4 \left(\frac{r_0}{R}\right)^2 \cdot \frac{1}{\alpha^4} \right].
 \end{aligned}$$

Dimensionless energy parameter:

$$\frac{U_{\text{GTP}}}{U_{\text{TP}}} \sim \phi_s \left(\frac{v_s}{c}\right)^4 \left(\frac{R}{L}\right) \left(\frac{r_0}{R}\right)^2 \cdot \frac{1}{\alpha^4}. \quad (18)$$



Memory Signature on atomic spectra

- After the end of radiation dominant era, light elements were formed. Lightest and the most abundant element during that epoch was hydrogen.
- In case of a merger of PBH, the produced gravitational waves would leave a signature on Hydrogen atomic cloud around them.
- Therefore, it is a unique signature of the event on the hydrogen fine structure spectra, hence called as **Precession Memory Effect(PME)**.

If ν is the frequency of the radiation without memory and $\nu'(M)$ is with PME, then

$$\begin{aligned} \delta\nu'(M) &= \nu \left[\phi_s \left(\frac{v_s}{c} \right)^4 \left(\frac{R}{L} \right) \left(\frac{r_0}{R} \right)^2 \cdot \frac{1}{\alpha^4} \right] \\ \Rightarrow \frac{\delta\nu'(M)}{\nu} &\simeq \left[\phi_s \left(\frac{v_s}{c} \right)^4 \left(\frac{r_0^2}{3 \times 10^3 \text{mL}} \right) \left(\frac{M_\odot}{M} \right) \cdot \frac{1}{\alpha^4} \right] \end{aligned} \quad (19)$$

