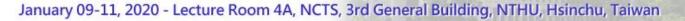
Confronting dark matter co-annihilation of Inert two Higgs Doublet Model with a compressed mass spectrum



Collaborators : Yue-Lin Sming Tsai, Van Que Tran Ref : arXiv:1912.08875





- 1. Introduction
- 2. Brief review of inert two higgs doublet model (i2HDM)
- 3. Constraints on i2HDM
- •4. Results & Discussions
- •5. Summary and Future Prospect



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Introduction – Dark Matter

Evidences for Dark Matter (DM)

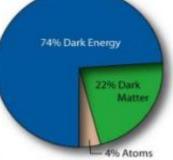
- WMAP measurement (Ω_m=0.25)
- rotation curves of galaxies
- the "bullet" cluster

Open Problems

- DM nature
- DM interactions
- DM formation mechanism

Detection techniques

- signals from colliders
- direct detection
- indirect detection of annihilation products such as neutrinos, antiprotons or gamma-rays



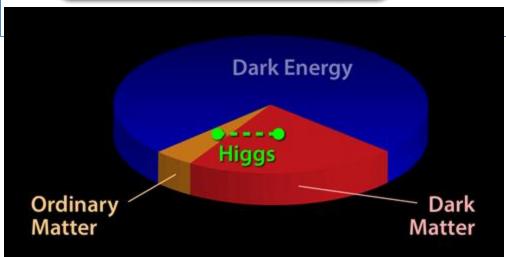


Chandra photo album: X-ray image of 1E0657-558



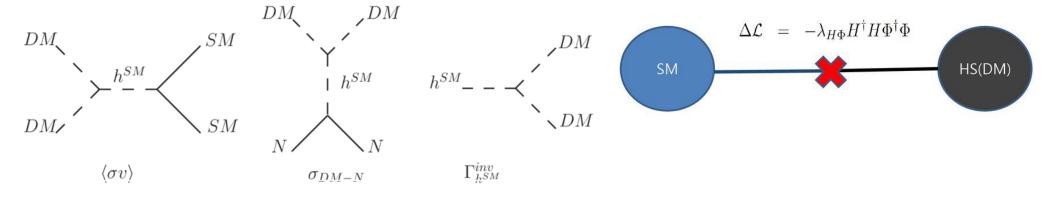
What is Dark Matter?

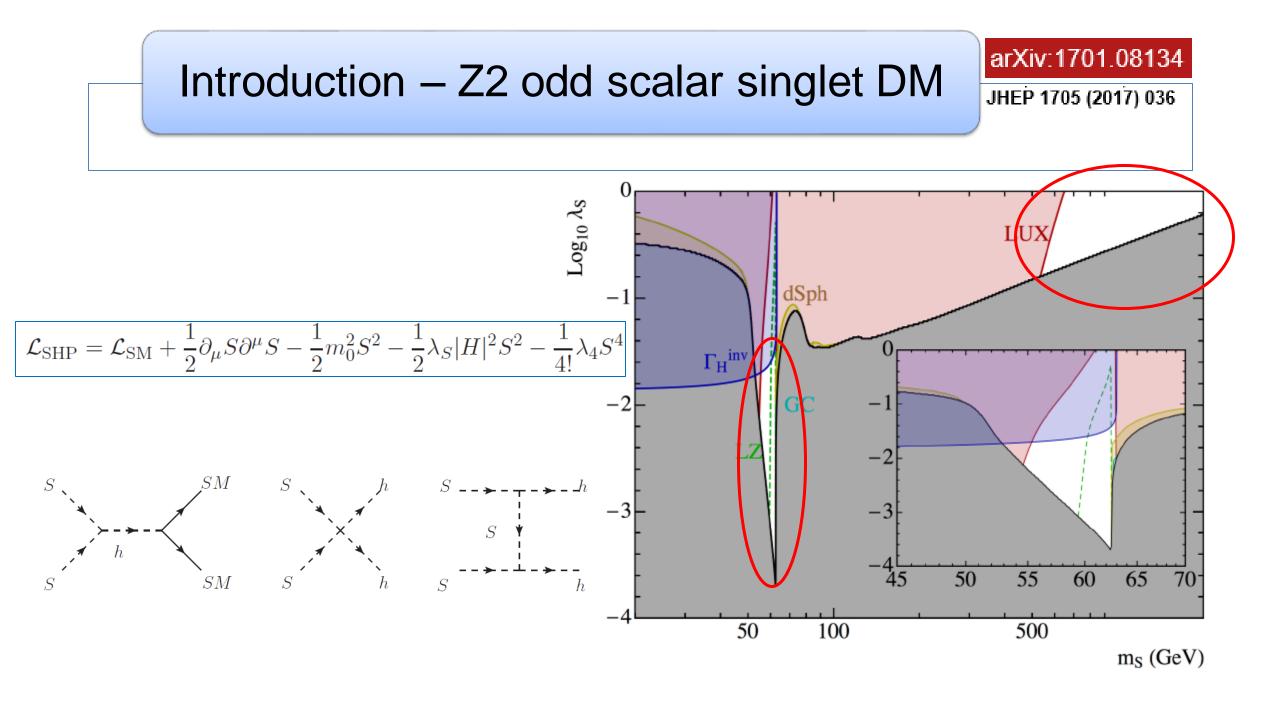
Introduction – Higgs Portal



Hidden Sector DM and Higgs Portal

• The renormalizable Higgs can mediate the interaction between the SM and hidden sector





Introduction – Z2 odd scalar singlet DM

arXiv:1701.08134

JHEP 1705 (2017) 036

65

m_S (GeV)

500

 70°

Even though this simple model has not been totally ruled out, but it's relatively boring ... (1) Is it possible to extend this simple model with co-annihilation processes ?

(2) Is it possible to extend this simple model with non-trivial DM direct detections & collider searches ?

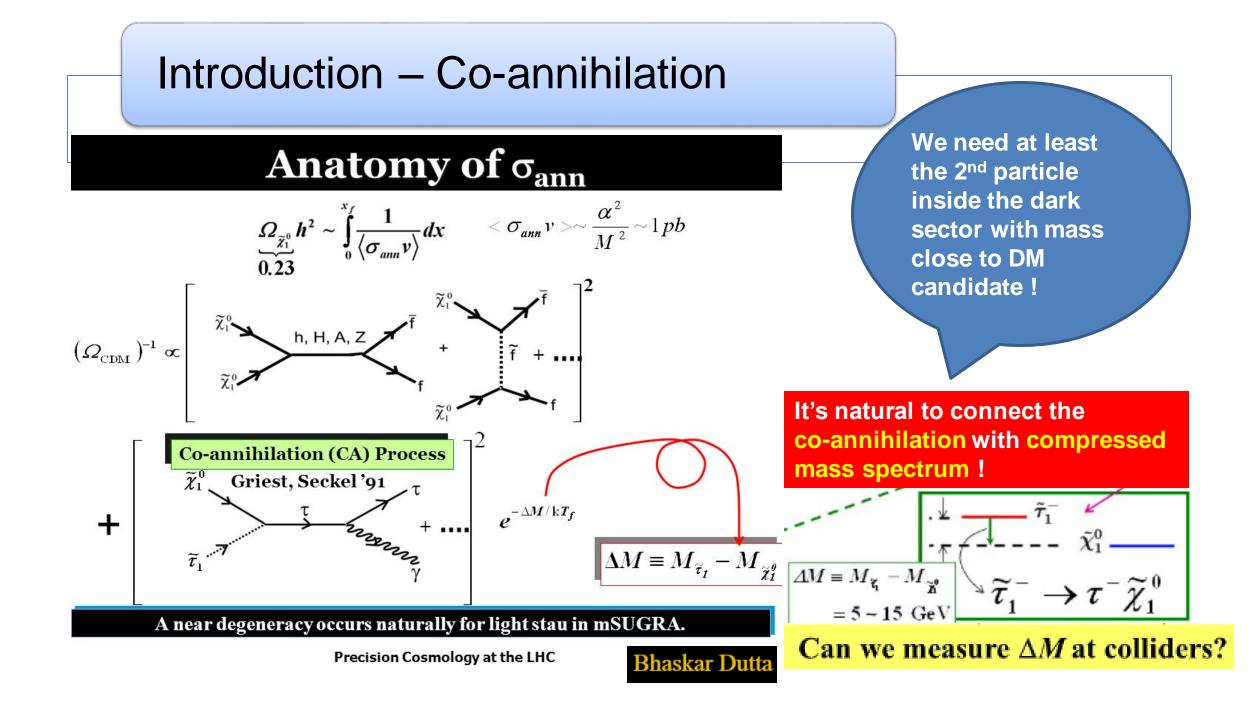
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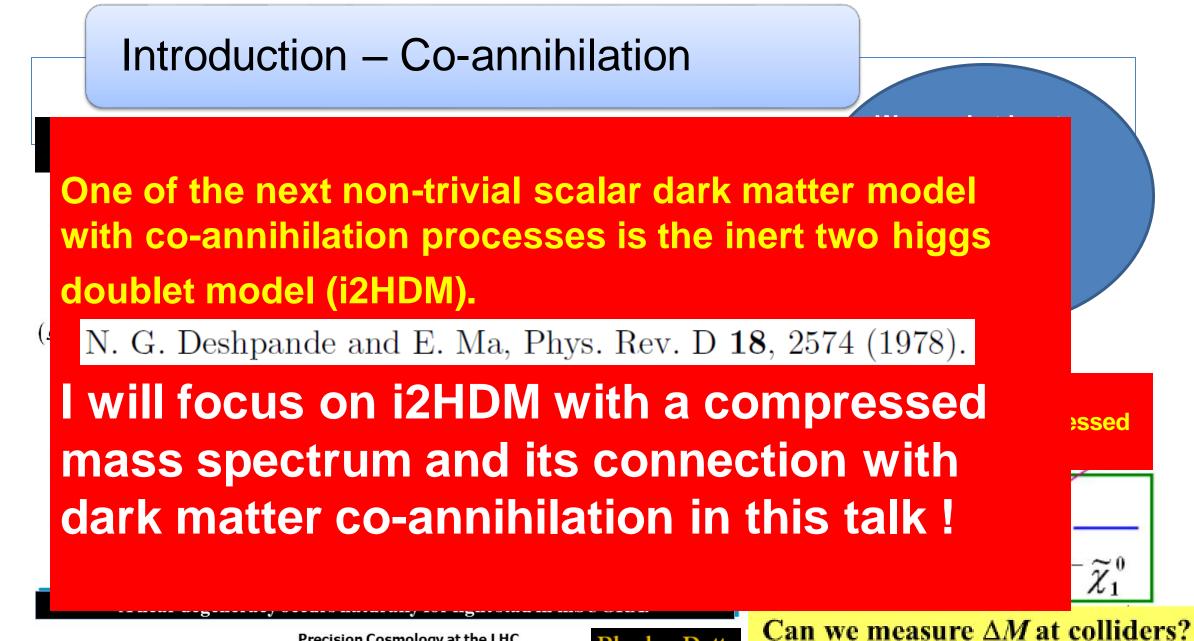
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 $\mathcal{L}_{\text{SHP}} =$

S

,* ;*







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$$H_{1} = \begin{pmatrix} G^{+} \\ \frac{1}{\sqrt{2}} (v + h + iG^{0}) \end{pmatrix}, \text{ and } H_{2} = \begin{pmatrix} H^{+} \\ \frac{1}{\sqrt{2}} (S + iA) \end{pmatrix}.$$
 (1)

Here, G^{\pm} and G^{0} are charged and neutral Goldstone bosons respectively. The symmetry breaking pattern for the doublets are $\langle H_{1}^{T} \rangle = (0, v/\sqrt{2})$ and $\langle H_{2}^{T} \rangle = (0, 0)$. In the end, we have five physical mass eigenstates: two CP-even neutral scalar h and S, one CP-odd neutral scalar A, and a pair of charged scalars H^{\pm} .

Let us briefly recap the main features of the i2HDM First \mathbb{Z}_2 -odd particles S, A and H^{\pm} are not directly couple to SM fermions while \mathbb{Z}_2 -even Higgs h plays the role of the SM Higgs with mass ~ 125 GeV. Second owing to the exact \mathbb{Z}_2 symmetry, there is no tree-level flavor changing neutral current. Finally the DM candidate can be either S or A depending on their mass, but it is hard to phenomenologically distinguish one from the other, as pointed out in Ref. [18]. Hence, we restrict ourselves to focus on the CP-even scalar S as DM candidate rather than CP-odd pseudo scalar A.

$$V = \mu_1^2 |H_1|^2 + \mu_2^2 |H_2|^2 + \lambda_1 |H_1|^4 + \lambda_2 |H_2|^4 + \lambda_3 |H_1|^2 |H_2|^2 + \lambda_4 |H_1^{\dagger} H_2|^2 + \frac{\lambda_5}{2} \left\{ (H_1^{\dagger} H_2)^2 + \text{h.c.} \right\} .$$

Conventionally, it is more intuitive to adopt the physical mass basis as inputs -

$$\begin{split} m_h^2 &= -2\mu_1^2 = 2\lambda_1 v^2, \\ m_S^2 &= \mu_2^2 + \frac{1}{2}(\lambda_3 + \lambda_4 + \lambda_5)v^2 = \mu_2^2 + \lambda_S v^2, \\ m_A^2 &= \mu_2^2 + \frac{1}{2}(\lambda_3 + \lambda_4 - \lambda_5)v^2 = \mu_2^2 + \lambda_A v^2, \\ m_{H^{\pm}}^2 &= \mu_2^2 + \frac{1}{2}\lambda_3 v^2, \end{split}$$

where we denote

$$\lambda_S = \frac{1}{2}(\lambda_3 + \lambda_4 + \lambda_5), \text{ and } \lambda_A = \lambda_S - \lambda_5 = \lambda_S + \frac{m_A^2 - m_S^2}{v^2}.$$

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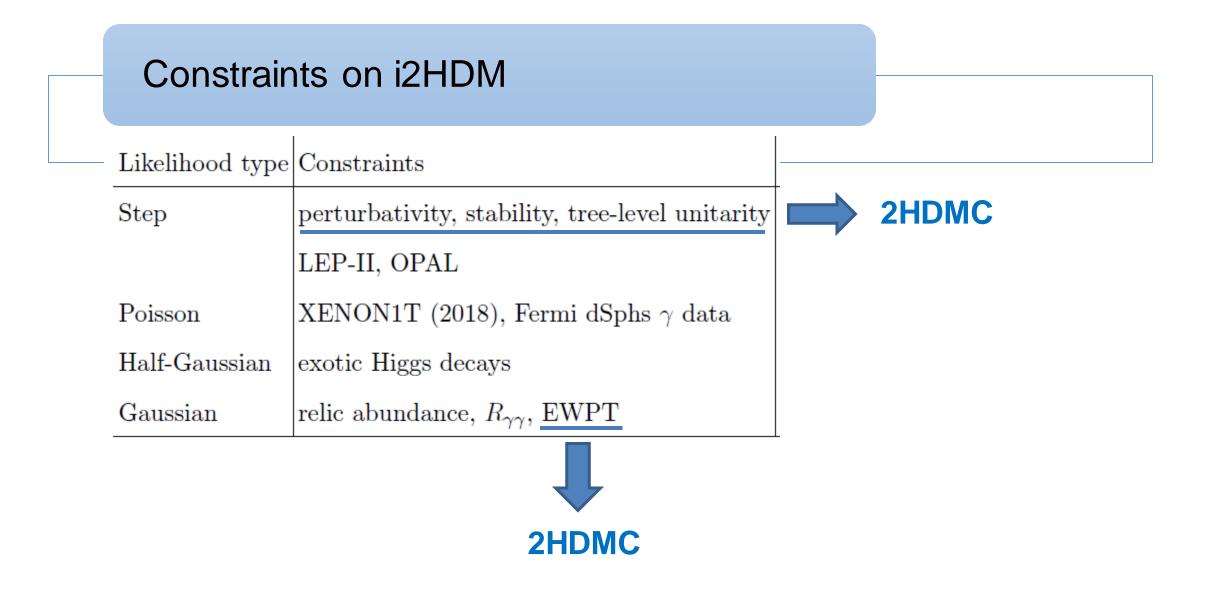
For the scenario with the compressed mass spectra, the mass splitting parameters $\Delta^0 = m_A - m_S$ and $\Delta^{\pm} = m_{H^{\pm}} - m_S$ instead of m_A and $m_{H^{\pm}}$ are more useful. Hence, our input parameters are $\{m_S, \Delta^0, \Delta^{\pm}, \lambda_2, \lambda_S\}.$ (6)

where we denote

$$\lambda_S = \frac{1}{2}(\lambda_3 + \lambda_4 + \lambda_5), \text{ and } \lambda_A = \lambda_S - \lambda_5 = \lambda_S + \frac{m_A^2 - m_S^2}{v^2}.$$

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Constraints on i2HDM

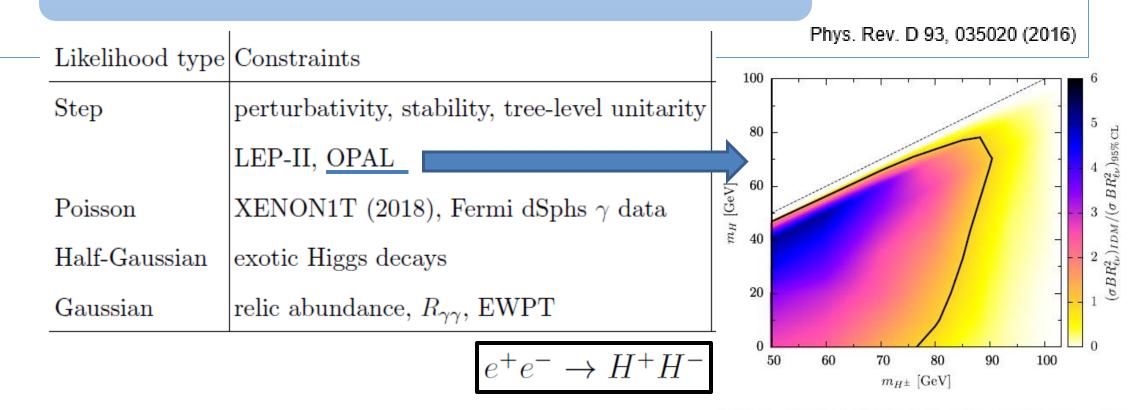
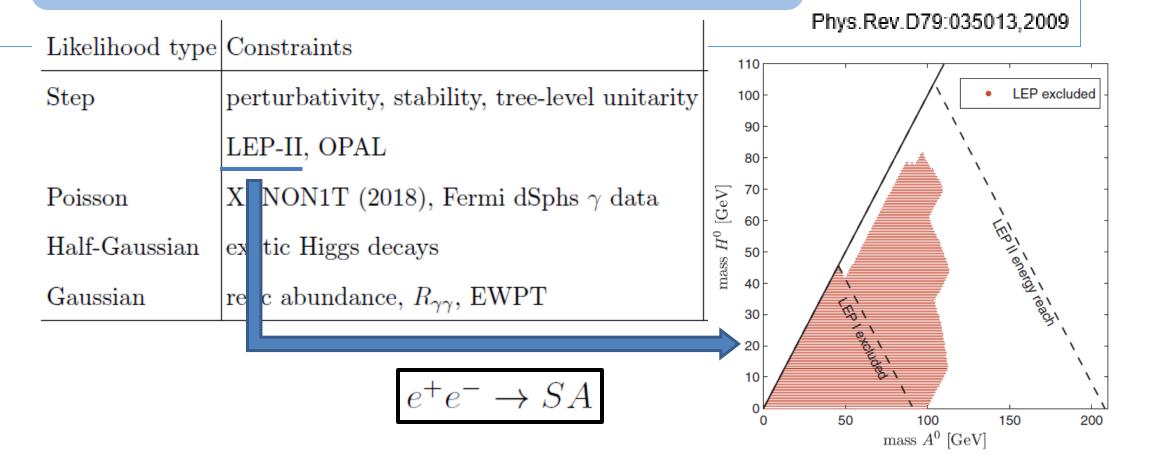
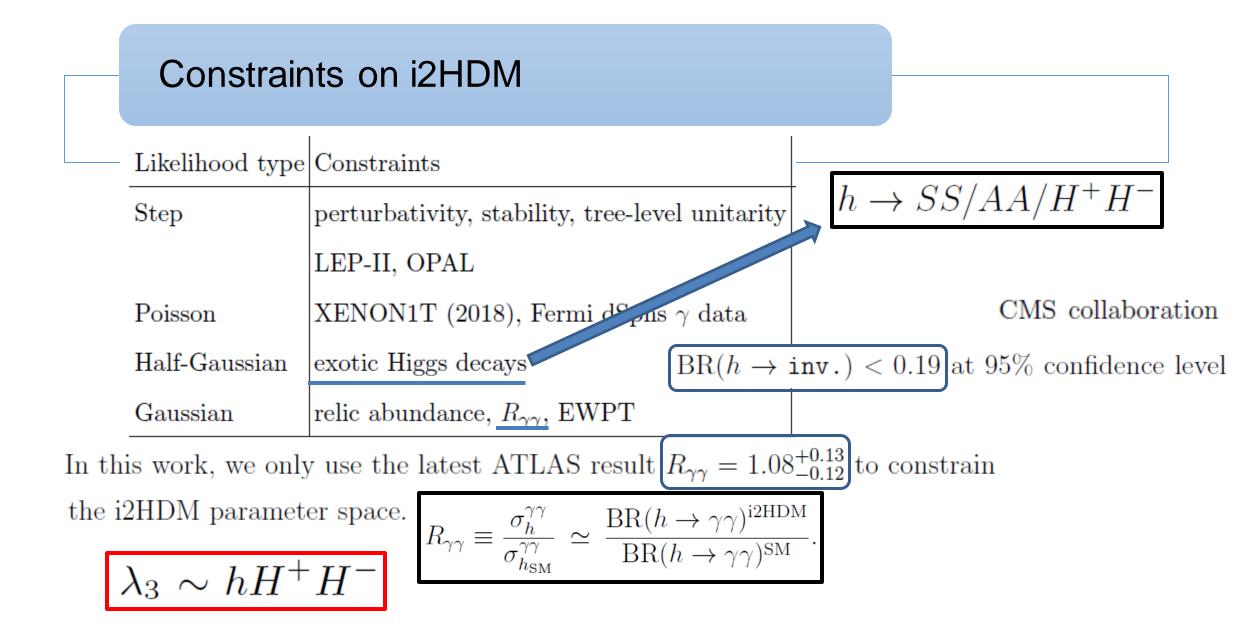


FIG. 5. Excluded region (95% C.L.) in the $m_{H^{\pm}} - m_H$ plane derived from the OPAL search for leptons and missing energy of Ref. [65]. The color map shows the magnitude of $(\sigma BR_{\ell\nu}^2)_{IDM}/(\sigma BR_{\ell\nu}^2)_{95\% C.L.}$.

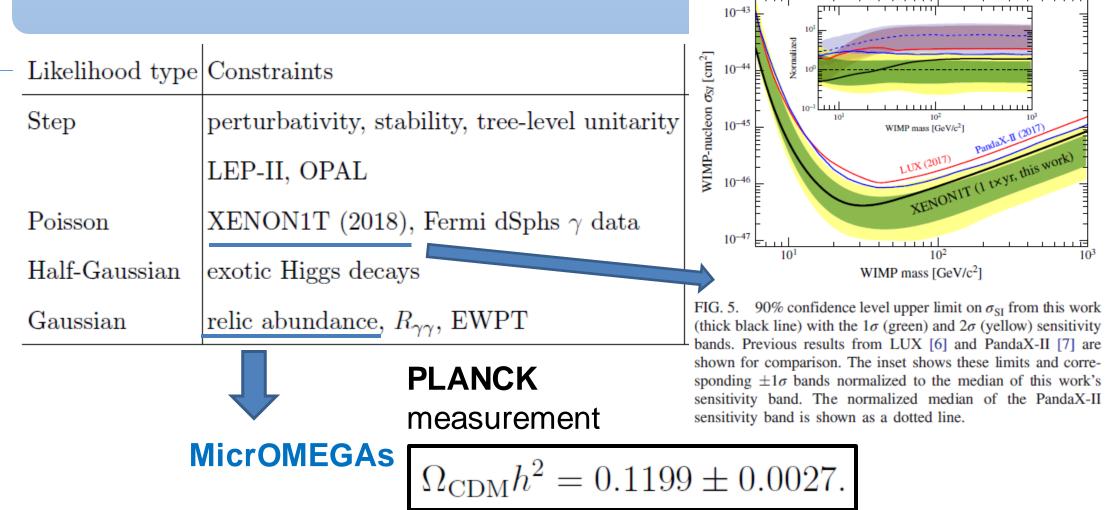
Constraints on i2HDM







Constraints on i2HDM



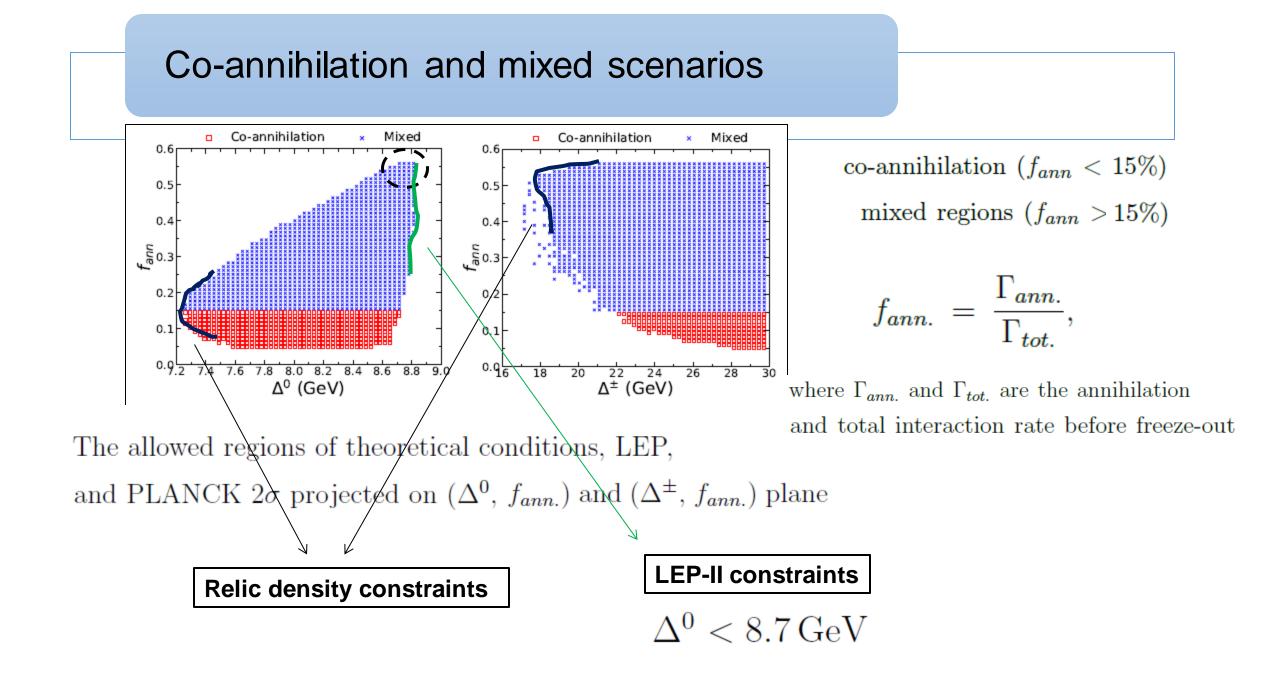
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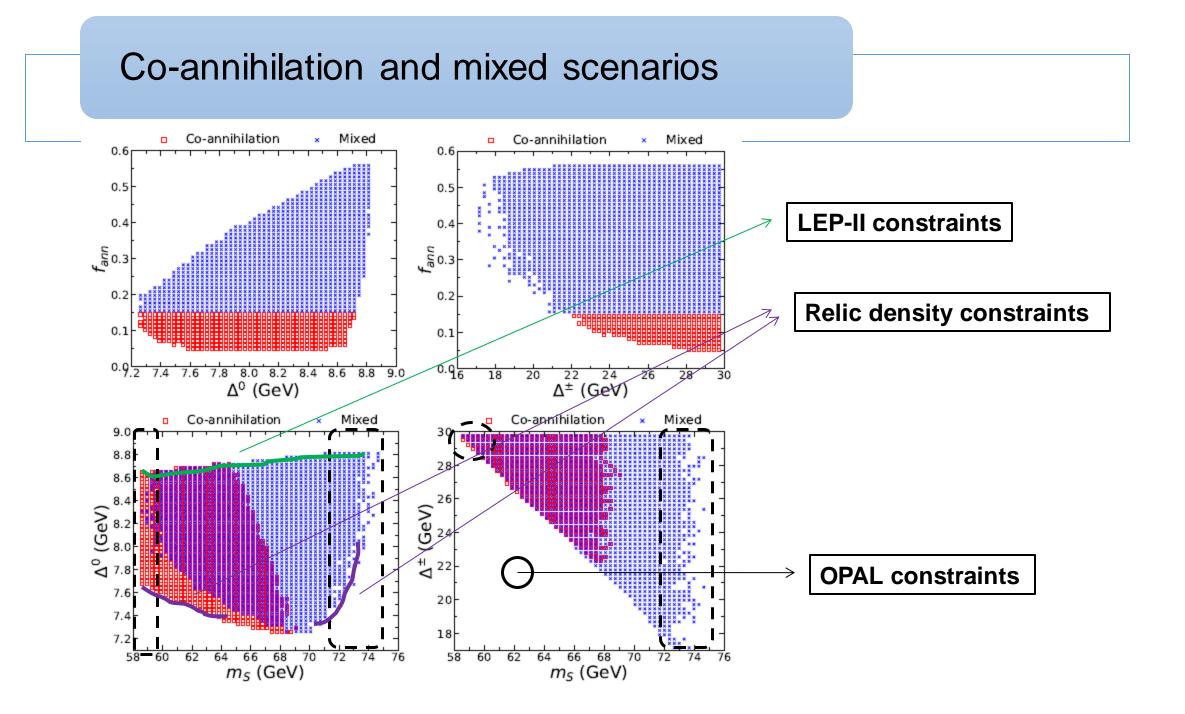


Results & Discussions

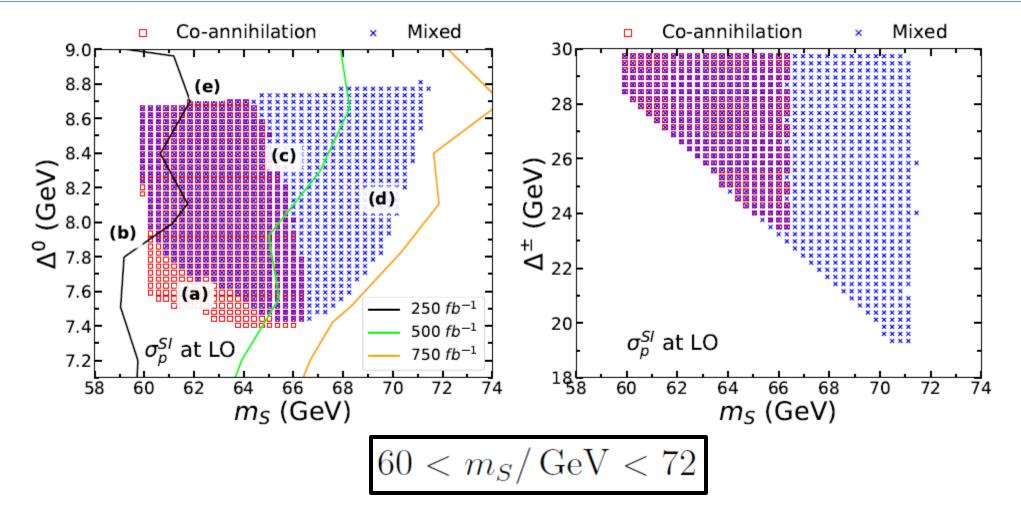
5.0 <	$\leq m_S / \text{GeV} \leq$	£ 100.0 ,		
10-3		20	Likelihood type	Constraints
$10^{\circ} \leq 2$	$\leq \Delta^0 / \text{ GeV} \leq$	20,	Step	perturbativity, stability, tree-level unitarity
1.0 <i>≤</i>	$\leq \Delta^{\pm}/$ GeV \leq	30,		LEP-II, OPAL
0.0		2.0	Poisson	XENON1T (2018), Fermi dSph s γ data
-2.0	$\leq \lambda_S \leq$	2.0,	Half-Gaussian	exotic Higgs decays
0.0	$\leq \lambda_2 \leq$	4.2.	Gaussian	relic abundance, $R_{\gamma\gamma}$, EWPT

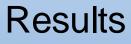












Region (a), the small Δ^0 can provide an efficient co-annihilation to reach the correct relic density even without the annihilation. Generally speaking, once the annihilation process $SS \to W^{\pm}W^{\mp}$ is open, it can be too efficient to reduce the relic density. The only way to get rid of annihilation is via a cancellation between four points interaction and *s*-channel Higgs exchange. Therefore, the coupling λ_S shall be negative. On the other hand, the λ_S can be very small in this region as long as $SS \to W^{\pm}W^{\mp}$ is close.

Region (b), comparing with the lower panel in Fig. 1, involving the Higgs invisible decay constraint lifts the lower limit on DM mass to 60 GeV.

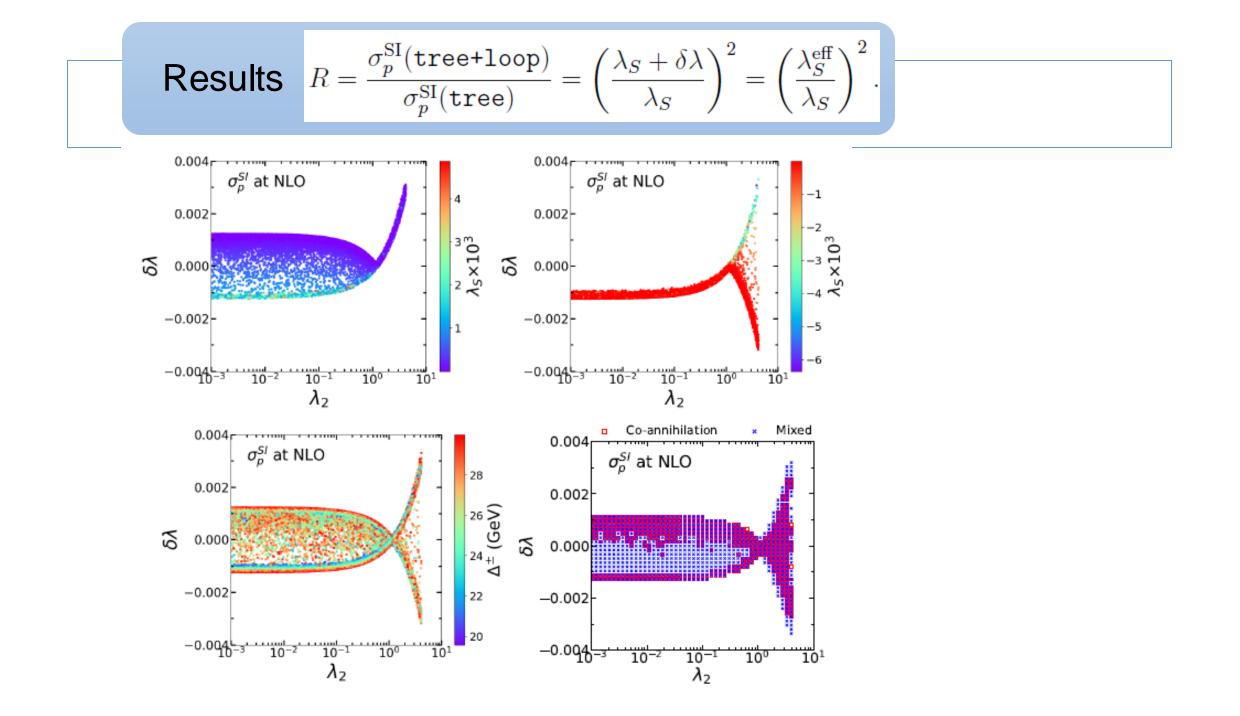


Region (c), the co-annihilation scenario cannot reach the large m_S regions, particularly $m_S < 64 \text{ GeV}$ for $\Delta^0 = 8.7 \text{ GeV}$ and $m_S < 66 \text{ GeV}$ for $\Delta^0 = 7.4 \text{ GeV}$. This is due to the current DM direct detection constraint. In particular, a more negative value of the coupling between the DM and Higgs boson is needed in the larger DM mass region so that the cancellation between $SS \to h^* \to WW^*$ and the four points interaction SSWW can occur to satisfy the correction relic density. However, this gives rise to the DM-proton scattering cross section and hence excluded by the XENON1T measurements. We also note that, due to the OPAL exclusion, a smaller DM mass region results in a larger Δ^{\pm} value. For the co-annihilation scenario, one can see that $\Delta^{\pm} > 23 \,\mathrm{GeV}$ as shown in the right panel of Fig. 2.

Results

Region (d), it is totally opposite to the region (a). The relic density at this region mainly comes from $SS \to W^{\pm}W^{\mp(*)}$ annihilation. Therefore, Δ^0 shall be large enough in order to suppress co-annihilation contributions. The lower bound of Δ^0 is varied with respect to m_S . The mixed scenario can reach a larger DM mass region as compared with the co-annihilation scenario, hence, due to the OPAL exclusion, the lower limit on Δ^{\pm} can be weaker. As shown in the right panel of Fig. 2, the mass splitting $\Delta^{\pm} > 19$ GeV for the mixed scenario.

Region (e), the LEP-II exclusion is presented. Together with the current XENON1T constraint, we found two important upper limits: $\Delta^0 < 8.8 \,\text{GeV}$ and $m_S < 72 \,\text{GeV}$.



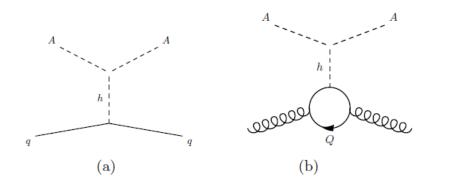




Figure 1. The diagrams which contribute to the spin-independent cross section at the leading order.

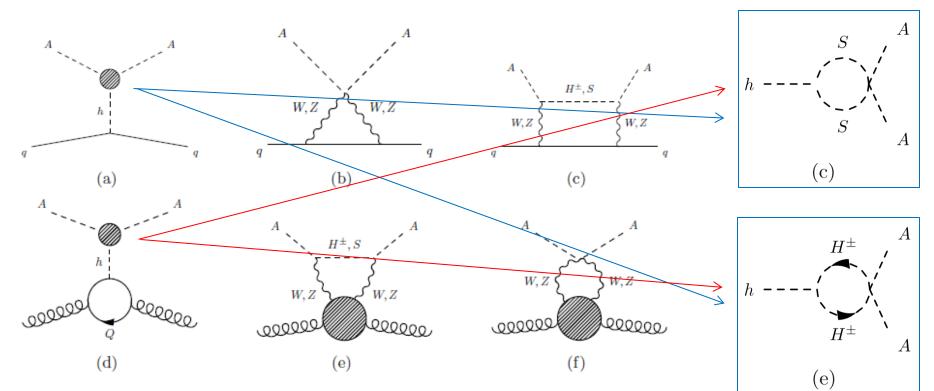
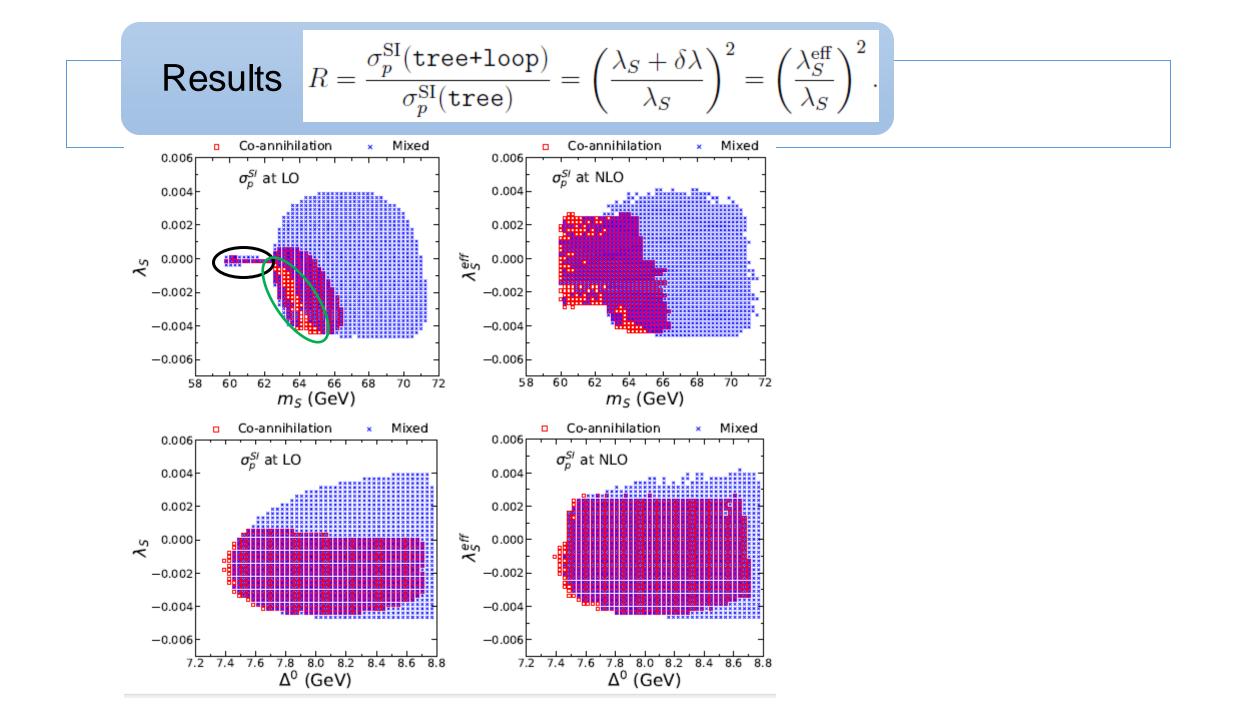
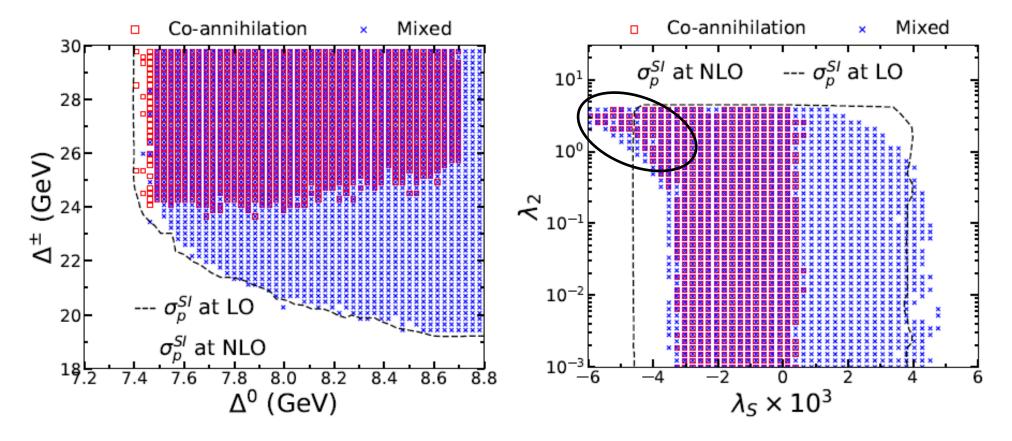


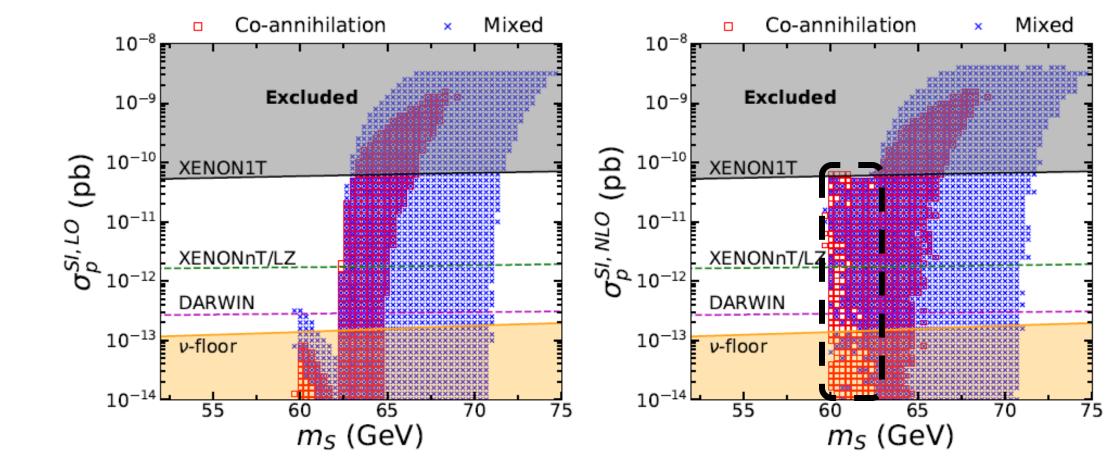
Figure 2. The diagrams we calculate. The shaded region is one-loop correction.

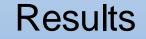












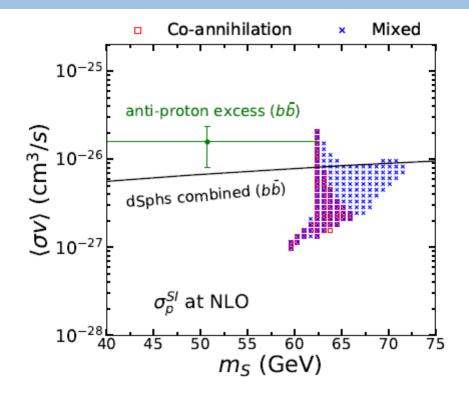


FIG. 7: The scatter plots in 2σ allowed region on the $(m_S, \langle \sigma v \rangle)$ plane. The color scheme for the co-annihilation (red square) and the mixed (blue cross) scenarios is the same as in Fig. 1. The solid black line represents the combined limit for DM annihilating into $b\bar{b}$ from observations of dSphs by Fermi-LAT, HAWC, HESS, MAGIC and VERITAS [79]. The green error bar is the 1σ signal region for anti-proton excess [80].

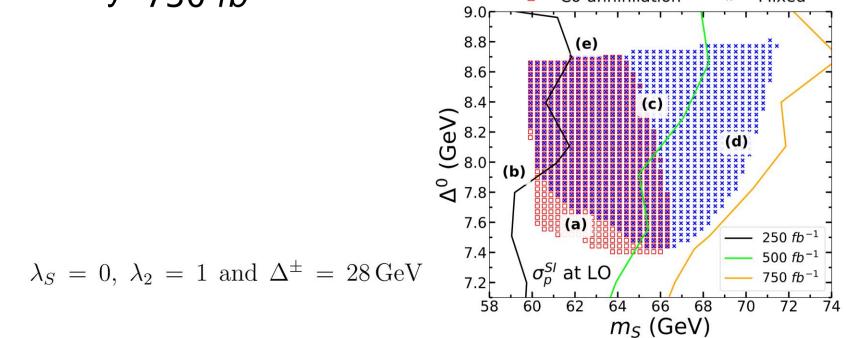
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Conclusions

- In this work, we have performed a global statistical analysis of the i2HDM with five parameters $(m_S, \Delta^0, \Delta^{\pm}, \lambda_2, \lambda_S)$ at the EW scale for compressed mass spectrum.
- By using the profile likelihood method, the survived parameter space were subjected to constraints from the theoretical conditions (perturbativity, stability, and tree-level unitarity), the collider limits (electroweak precision tests, LEP, and LHC), the relic density as measured by PLANCK, the σ_p^{SI} limit from XENON1T, and the ⟨σv⟩ limit from Fermi dSphs gamma ray data.

- We found that the viable parameter space of coannihilation scenario is located at $60 \text{ GeV} \lesssim m_S \lesssim 66 \text{ GeV}$, the $7.4 \text{ GeV} \lesssim \Delta^0 \lesssim 8.8 \text{ GeV}$, and $\Delta^{\pm} \gtrsim 23.0 \text{ GeV}$.
- The allowed parameter space can be tested by the compressed mass spectra searches at the LHC. It can be partially probed with future luminosity $250 \ fb^{-1}$ and mostly probed with luminosity $750 \ fb^{-1}$



Thank you for your attention

