# Dark Matter & Reactor Neutrino Physics with sub-keV Germanium Detectors

- > Facilities: KSNL & CJPL
- >Neutrinos: Physics with Ge at Reactor
- > Dark Matter: Light WIMPs Searches





中国锦屏地下实验室 China Jinping Underground Laboratory

Henry T. Wong /王子敬 Academia Sinica, Taiwan /中央研究院

2nd International Workshop on

Particle Physics and Cosmology after Higgs and Planck

後希格斯與普朗克粒子物理與宇宙學國際研討會

October 8-9, 2014 - National Center for Theoretical Sciences, Hsinchu, Taiwan October 10-11, 2014 - Fo-Guang-Shan, Kaohsiung, Taiwan

### **TEXONO-CDEX** Collaboration

Taiwan EXperiment On NeutrinO

[since 1997]

- Neutrino Physics at Kuo-Sheng Reactor Neutrino Laboratory (KSNL)
  - > Taiwan (AS, INER, KSNPS)
  - **► India (BHU)**
  - **► Turkey (METU,DEU)**



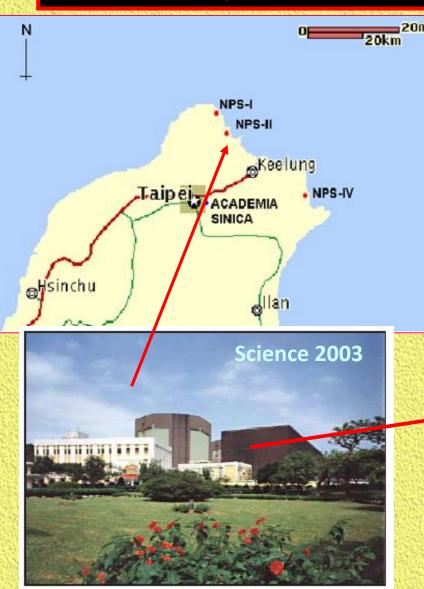
CDEX China Dark Matter EXperiment

[birth 2009]

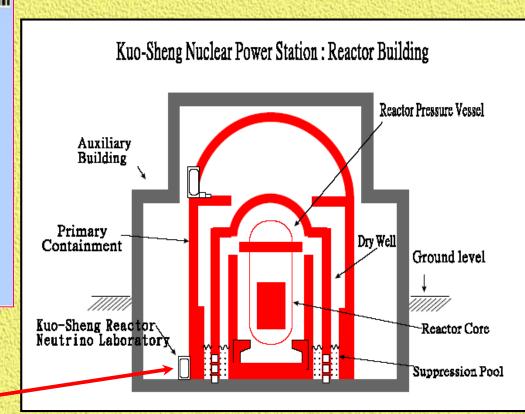


- O Dark Matter Searches at China Jin-Ping **Underground Laboratory (CJPL)** 
  - > China (THU, CIAE, NKU, SCU, YLJHD)

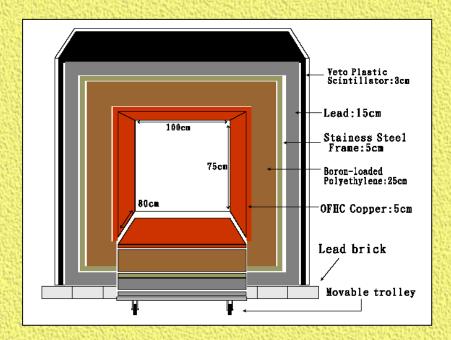
# Kuo Sheng Reactor Neutrino Laboratory [KSNL]



**Powerful collaboration**. Scientists from Taiwan and mainland China are studying neutrino emissions from this nuclear power plant outside Taipei.



- 28 m from core#1 @ 2.9 GW
- Shallow site: ~30 mwe overburden
- ~10 m below ground level







Front View (cosmic vetos, shieldings, control room ....)

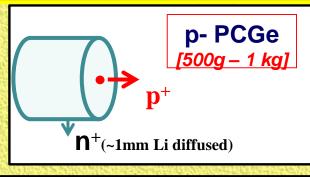
**Configuration: Modest yet Unique** 

Flexible "Baseline" Design:

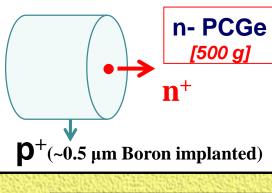
Allows different detectors conf. for different physics

### Baseline Hardware Design

[Both TEXONO@KSNL & CDEX-1@CJPL]

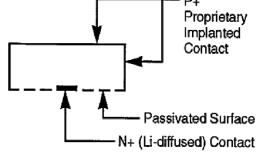




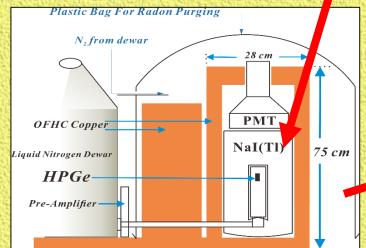


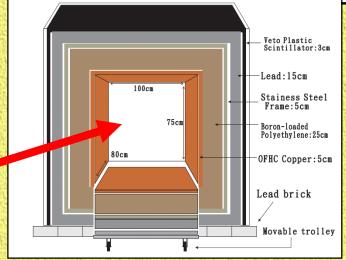




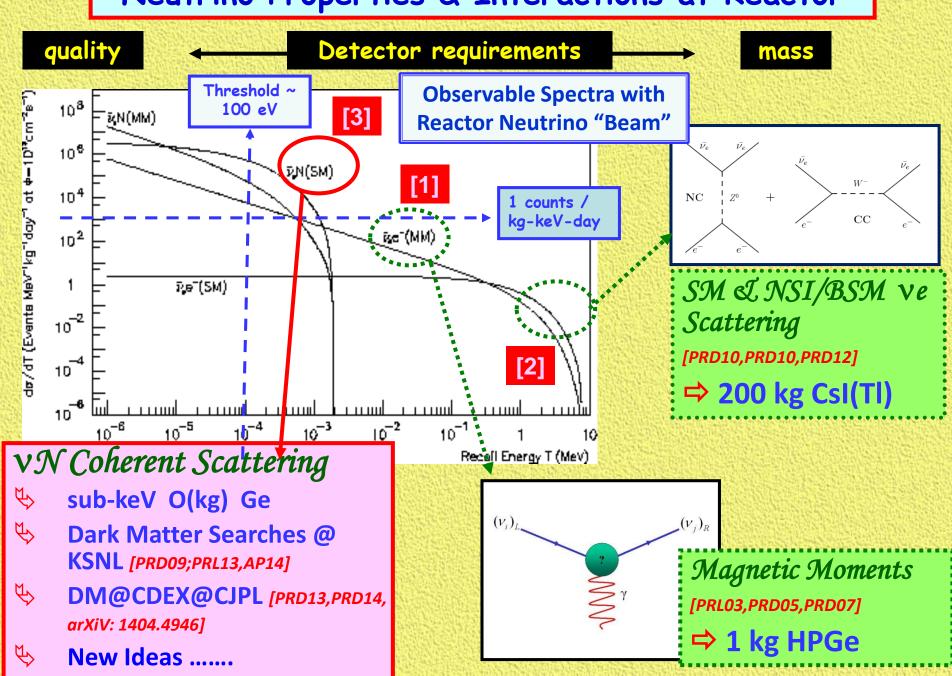








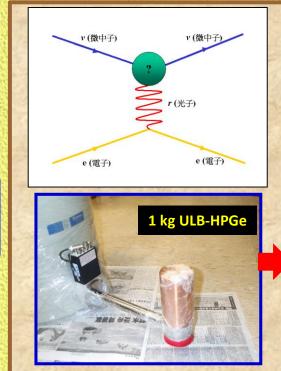
### Neutrino Properties & Interactions at Reactor





Reactor Neutrino

@ KSNL: Summary



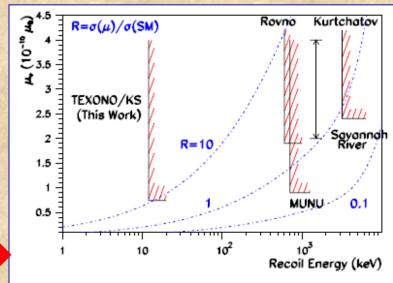
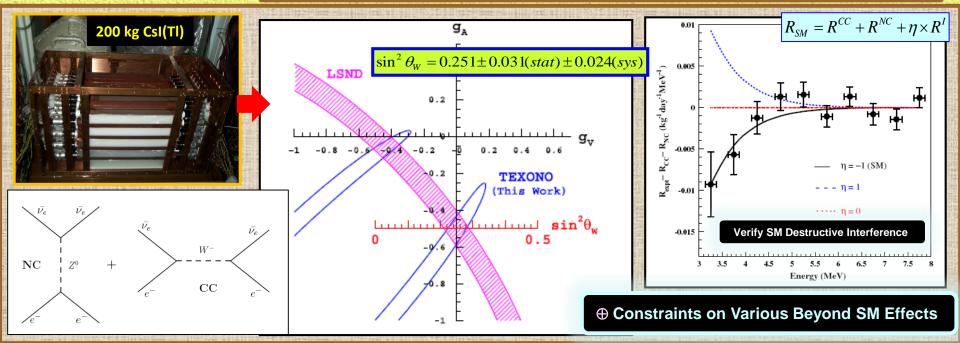


FIG. 14 (color online). Summary of the results in the searches of neutrino magnetic moments with reactor neutrinos. Both the limits and the detection thresholds of the various experiments are shown.



### NEW (!): Neutrino "Milli-charge" [+Chen, Liu, Chi; PRD14]

**Neutrino Electromagnetic Form Factors** 

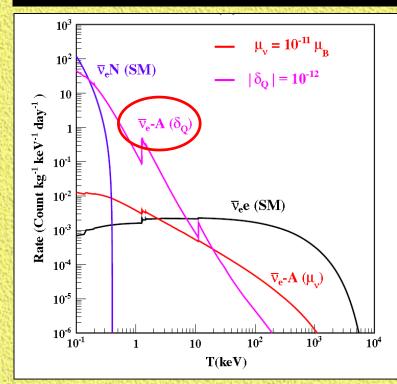
$$\Gamma^{\mu}_{\rm em} \equiv F_1 \cdot \gamma^{\mu} + F_2 \cdot \sigma^{\mu\nu} \cdot q_{\nu}$$

$$F_1 = \delta_Q \cdot e_0 + \frac{1}{6} \cdot q^2 \cdot \langle r_{\nu}^2 \rangle,$$
  $F_2 = (-i) \cdot \frac{\mu_{\nu}}{2 \cdot m_e},$ 

$$F_2 = (-i) \cdot \underbrace{\begin{pmatrix} \mu_{\nu} \\ 2 \cdot m_e \end{pmatrix}}_{,}$$

**Atomic Ionization Differential Cross-Section** with full atomic physics many-body "MCRRPA" calculation [PL13]

$$\overline{\nu_{\rm e}} + {\rm A} \rightarrow \overline{\nu_{\rm e}} + {\rm A}^+ + e^-,$$



- transfer ("minimum ionizing")
- **Smoking-gun signatures for positive** signals: peaks at known K/L binding energy at known ratios [different from cosmic-activation electron-capture background]
- ✓ Present Bound :  $\delta_0$  < 10<sup>-12</sup>
- **Future Sensitivity Goal (100 eVee** threshold):  $\delta_0 \sim 10^{-14}$

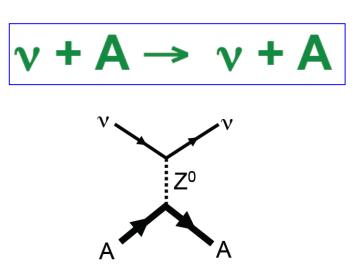
## Current Research Theme:

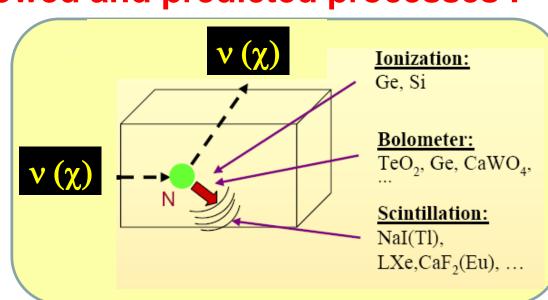
# "sub-keV" Ge Detectors

- Physics Goals for O[100 eV threhold⊕1 kg mass⊕1 cpkkd] detector:
  - N coherent scattering, potential applications to reactor monitoring
  - Low-mass WIMP searches → [CDEX Program @CJPL]
  - Explore v/WIMP electromagnetic properties & interactions
  - Open & Explore new detector window & detection channel & physics parameter space

### Neutrino-Nucleus Coherent Scattering:

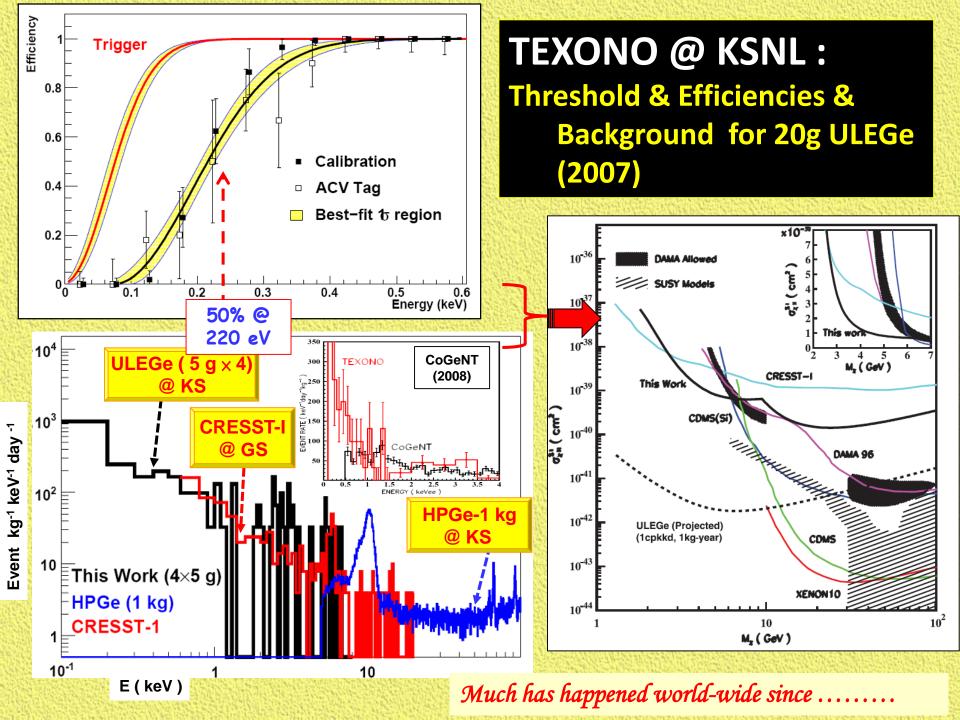
### > Standard Model allowed and predicted processes:





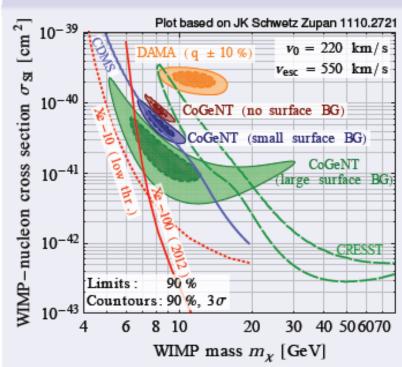
- Neutral current process (same for all v-flavor)
- $\succ \sigma \propto N^2$  @  $E_v < 50 \text{ MeV}$ 
  - ⇒ "Coherent" [probe "sees" the whole nucleus]
- > sensitive probe for BSM; interest in reactor monitoring
- > important process in stellar collapse & supernova explosion
- Irreducible background/analogous interaction in dark matter detection
- ➢ Ge at KSNL @ QF~0.2: cut-off ~ 300 eV ;

Rate ~10 kg-1 day-1 @ threshold~100 eV



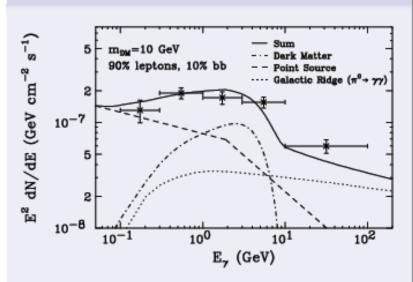
### Hints for light dark matter





- Several intriguing direct detection signals
- But severe tension with null results

#### ...and in the skies



• An tentative  $\gamma$  ray excess from the Galactic Center

Hooper Goodenough 0912.2998, 1010.2752, 1201.1303

- Morphology ≠ point source
- Radio filaments

Linden Hooper Yusef-Zadeh 1106.5493

Isotropic radio background

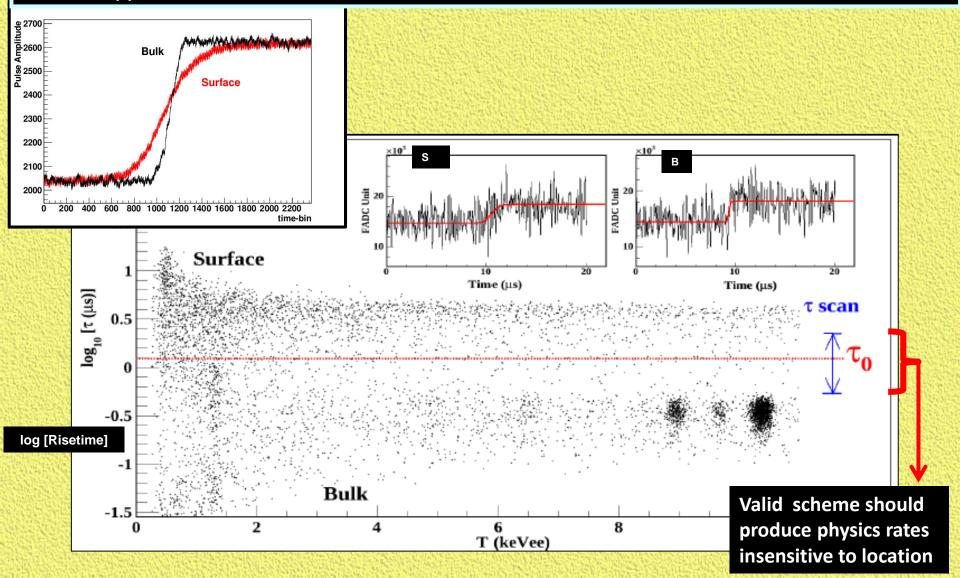
Hooper Belikov Jeltema Linden Profumo Slatyer 1203.3547

# Sub-keV Ge Detector Techniques: Users' R&D Items

- Quenching Factors -- nuclear recoils' lonization Yields
- **\*** Energy Definition & Calibration
- Trigger Efficiencies near threshold
- Bulk Vs Surface Events Selection algorithms & efficiencies
- Physics Vs Noise Pulse-Shape Selection -- algorithms & efficiencies [On-Going....]

### PSD for Surface Vs Bulk Events @ PCGe [AP14]

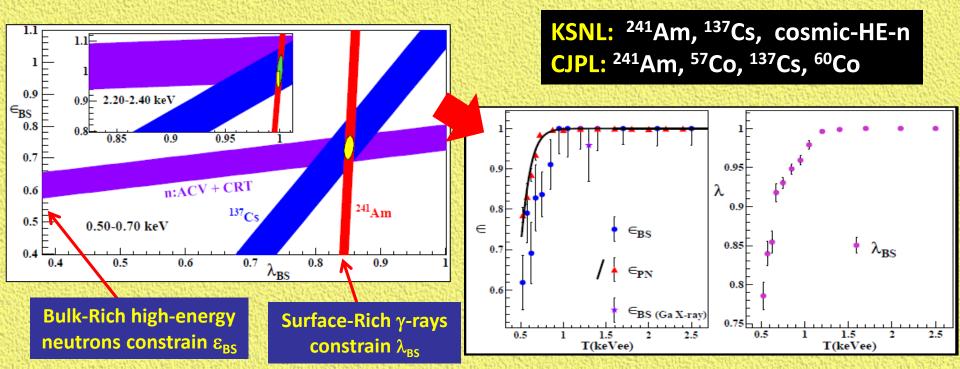
- n+ "inactive layer" is not totally dead; signals finite but slower rise time
- > ACV+CRT tag (cosmic-induced high energy neutrons) ⇒ no surface band
  - n-type PCGe ⇒ no surface band



"Calibration"  $\equiv$  measure energy-dependent signal-retaining ( $\varepsilon_{BS}$ ) & background-suppressing ( $\lambda_{BS}$ ) efficiencies, related by the coupled equations [B,S=real; B'S'=measured]:

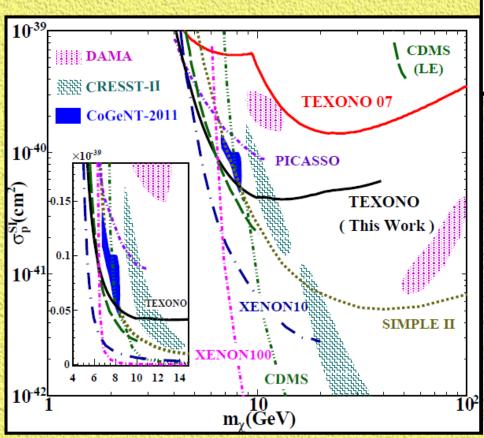
$$B' = \epsilon_{BS} \cdot B + (1 - \lambda_{BS}) \cdot S$$
  
$$S' = (1 - \epsilon_{BS}) \cdot B + \lambda_{BS} \cdot S$$

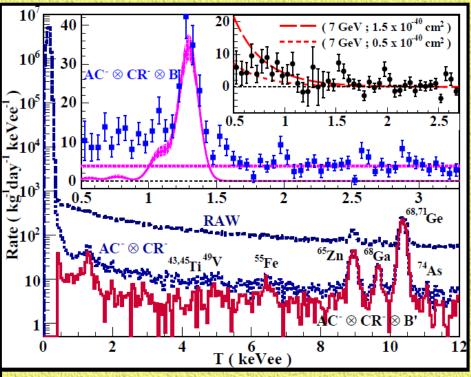
Approach: Identify *THREE(+)* calibration data [low and high energy  $\gamma$ , cosmic-induced neutrons] where (B,S) are known & (B',S') measured  $\oplus$  solve coupled equation for ( $\varepsilon_{BS}$ ,  $\lambda_{BS}$ )



# Light WIMP Searches @ KSNL with Ge

Learn & Establish Techniques
Catalyze CDEX-1 @ CJPL
Produce Physics Results!





### TEXONO@KSNL [PRL13]:

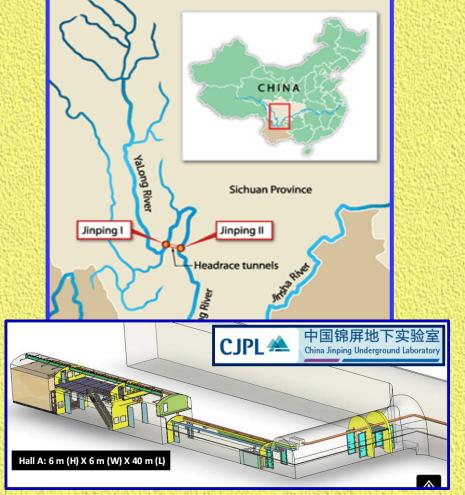
500 eV threshold

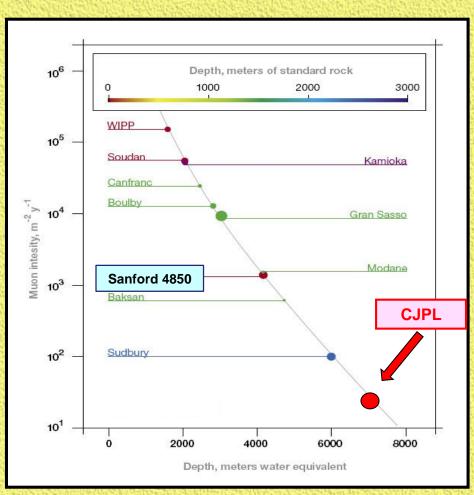
probed and excluded some light WIMP allowed regions implied by other experiments.

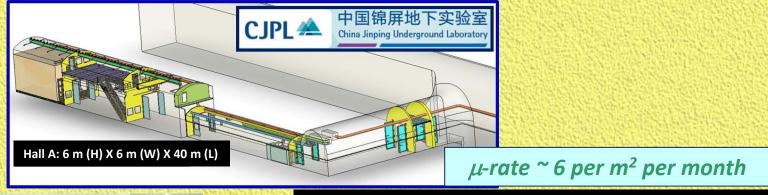
Provide probable explanations to CoGeNT-2011's excess



- 2400+ m rock overburden, drive-in road tunnel access
- 6X6X40 m cavern constructed [managed byTHU & EHDC]
- © CDEX-1 & Panda-X Dark Matter Program





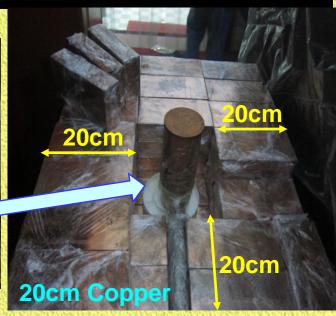




### CDEX-1 Data Taking @ CJPL:

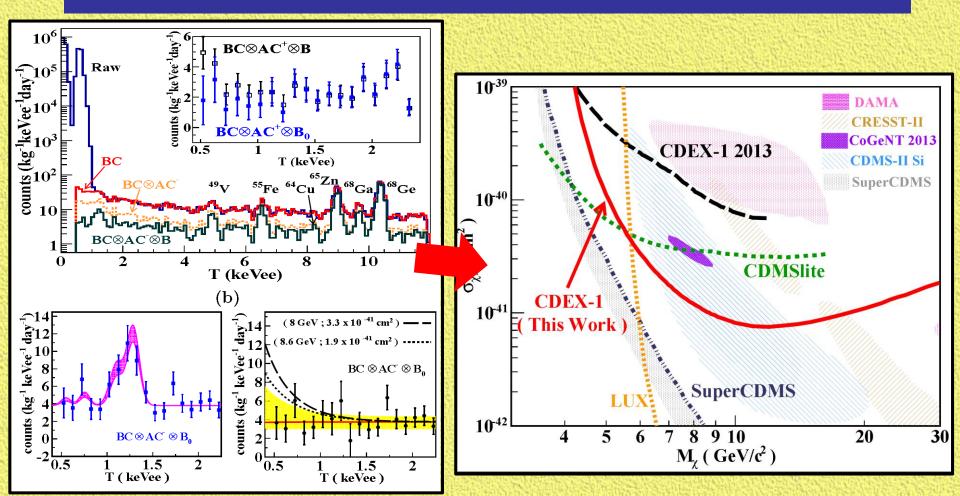
- Adopt KSNL Baseline Design
- Engineering Run 2011
- Physics Run June 2012



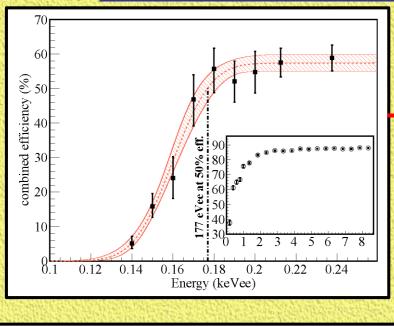


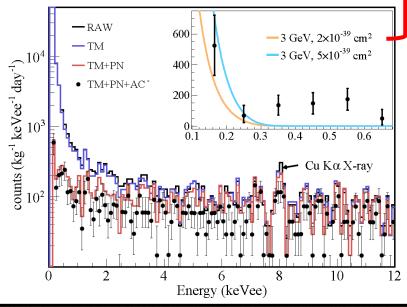
#### CDEX-1 @ CJPL 2014 [PRD13, arXiv: 1404.4946]

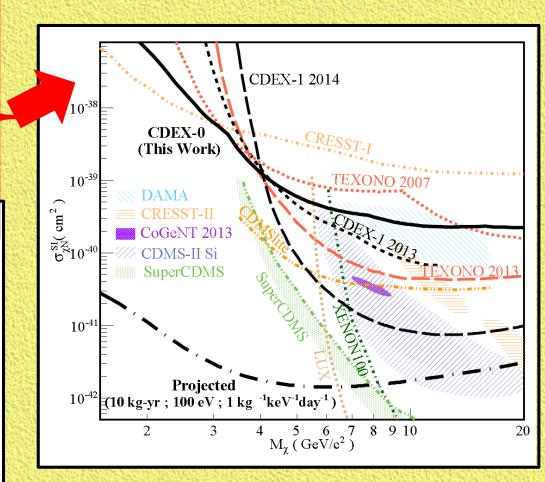
- 🔁 1 kg pPCGe @ 475 eVee threshold
- All events quantitatively accounted for; No Residual Excesses at sub-keV
- Exclude CoGeNT-2013 excess as WIMP-induced, independent of interaction channels



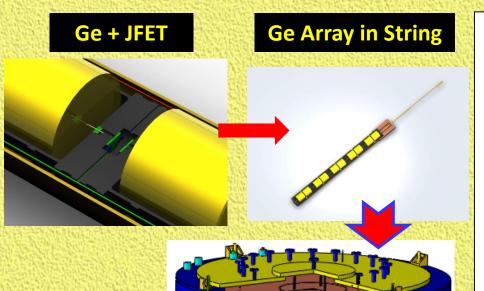
### CDEX-0 [20g prototype] @CJPL 2014 [PRD14] 12g ULEGe e @ 177 eV<sub>ee</sub> Analysis Threshold







### Design of CDEX-10: with LAr Anti-Compton



PCGe in Arrays & Strings

LiqAr (LiqN)as both cryogenics & active anti-Compton

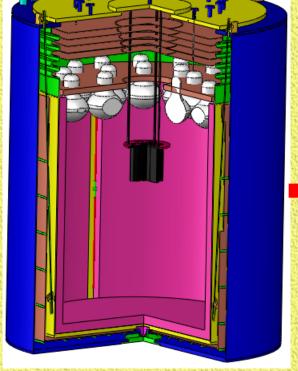
 $\sim$  30-40 cm  $4\pi$  shielding range

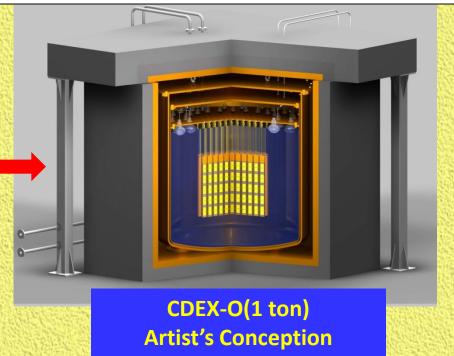
Prototype 2014

Baseline Design for Future O(1 ton) Expt for DM+0 $\nu$ ββ

CDEX-10 (2014+)

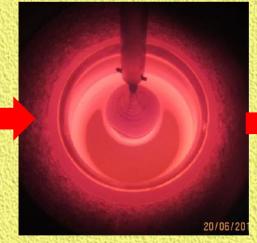






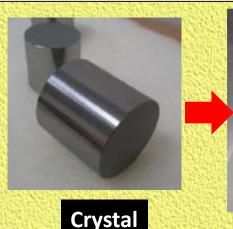
# Ge Processing & Assembly Facility @ THU





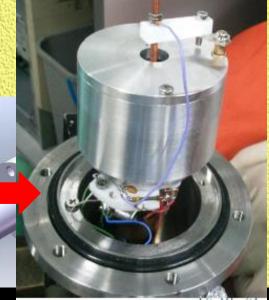


- Growth & Processing of raw Ge crystal
- Application-specific optimized assembly
- R&D on JFETs & Preamps
- Possible Future Commercial Marketing (THU)





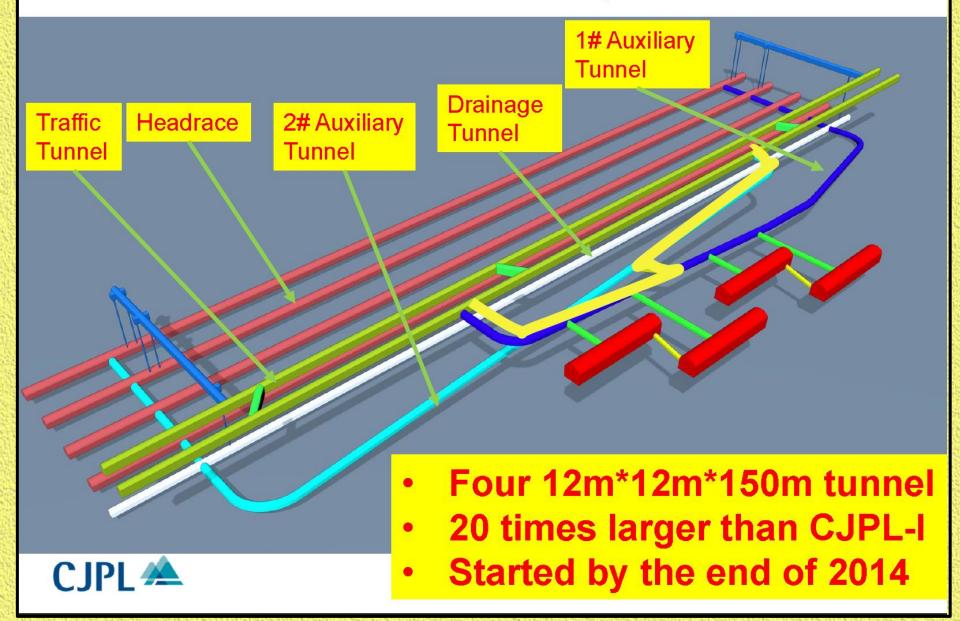




**Processing & Assembling** 

**Detector & Cryogenics** 

# CJPL-II development



### Summary & Prospects



- Competitive results on light WIMPs with sub-keV Ge with TEXONO@KSNL; 10+ improvement at CDEX-1@CJPL
- > Surface leakage to Bulk samples is important to PPCGe at low energy; source of earlier positive signal claims
- Current Theme: lower threshold; modulation studies; background understanding
- CJPL: more space, CDEX: bigger target mass; acquire L-Ar & Ge-techniques
- > KSNL: return to original goal vN coherent scattering ⊕ study exotic electromagnetic effects.