



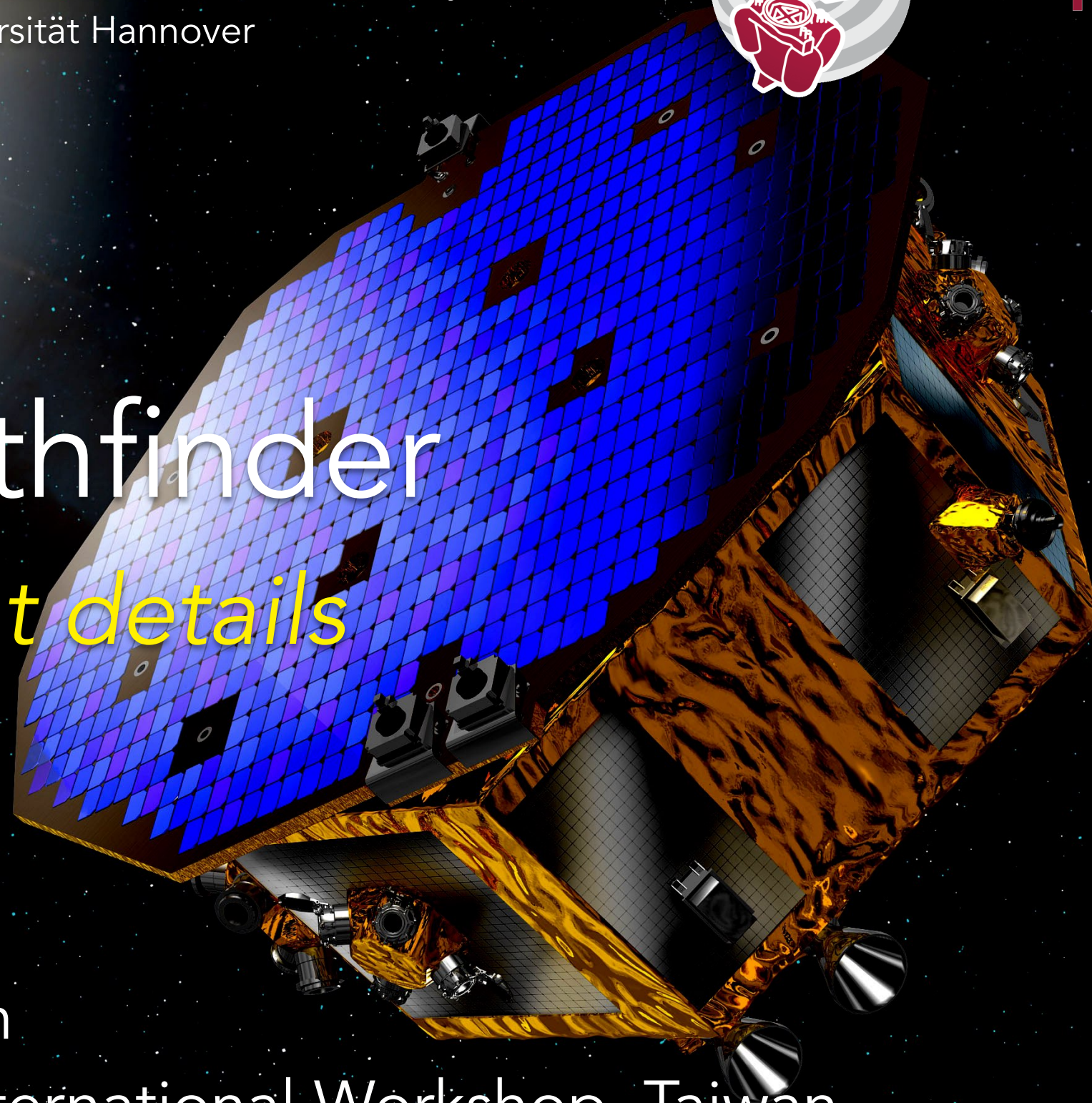
Albert Einstein Institute
Max Planck Institute for Gravitational Physics and
Leibniz Universität Hannover



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LISA Pathfinder

Experiment details & results



M Hewitson
for the LPF Team
ASTROD 2nd International Workshop, Taiwan
May 2017



- Mission aims
- Experiment plan
- Example experiments
- Conclusions

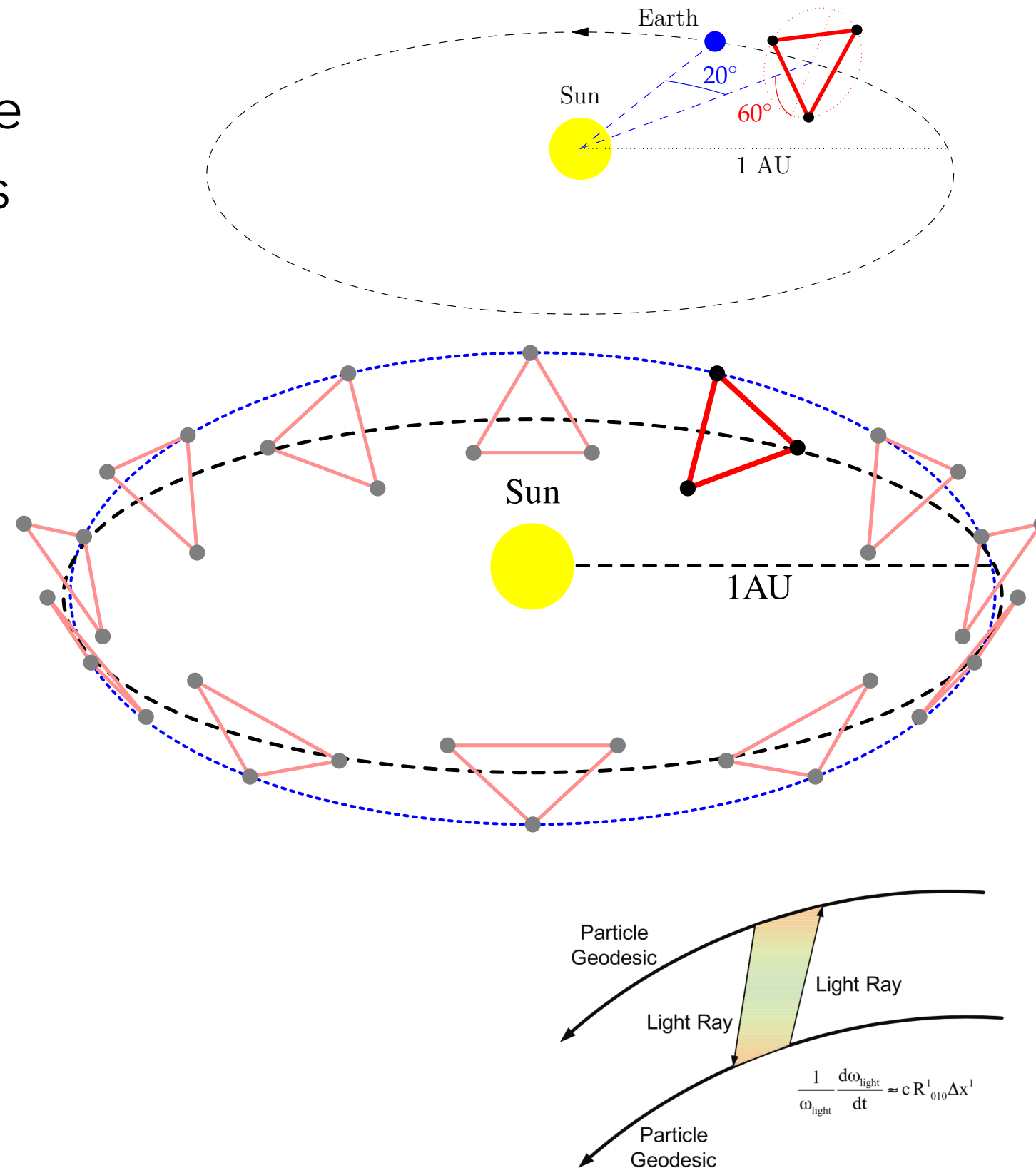
Introduction

The LISA Measurement Concept



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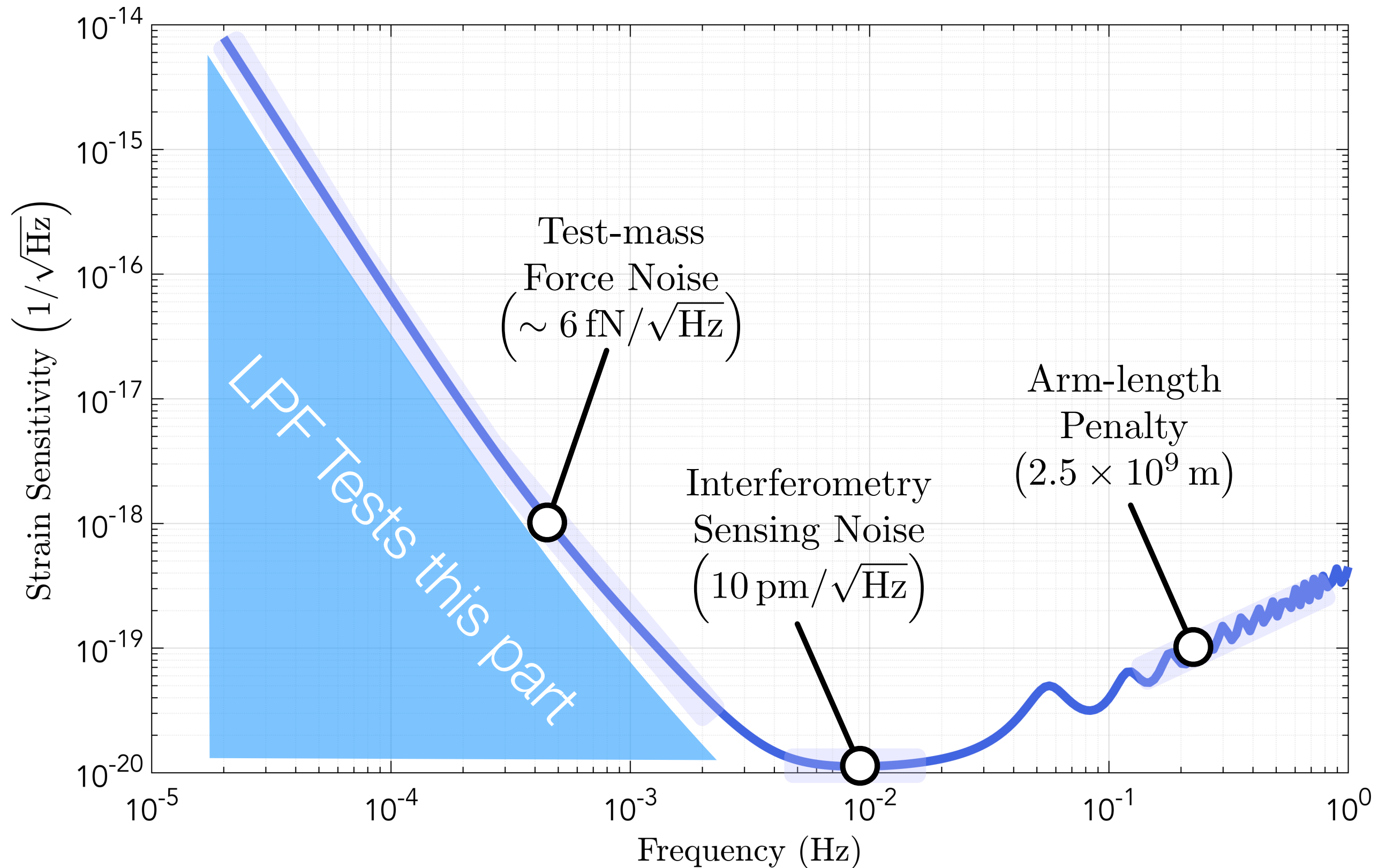
- Probe the change in proper time between free-falling test masses caused by GWs
- Free-falling test masses are housed in drag-free spacecraft separated by a few million km
- Proper time is inferred by the time of flight of photons exchanged between the satellites
- We have multiple transponder links from which we can form a Michelson-like signals



LISA Performance



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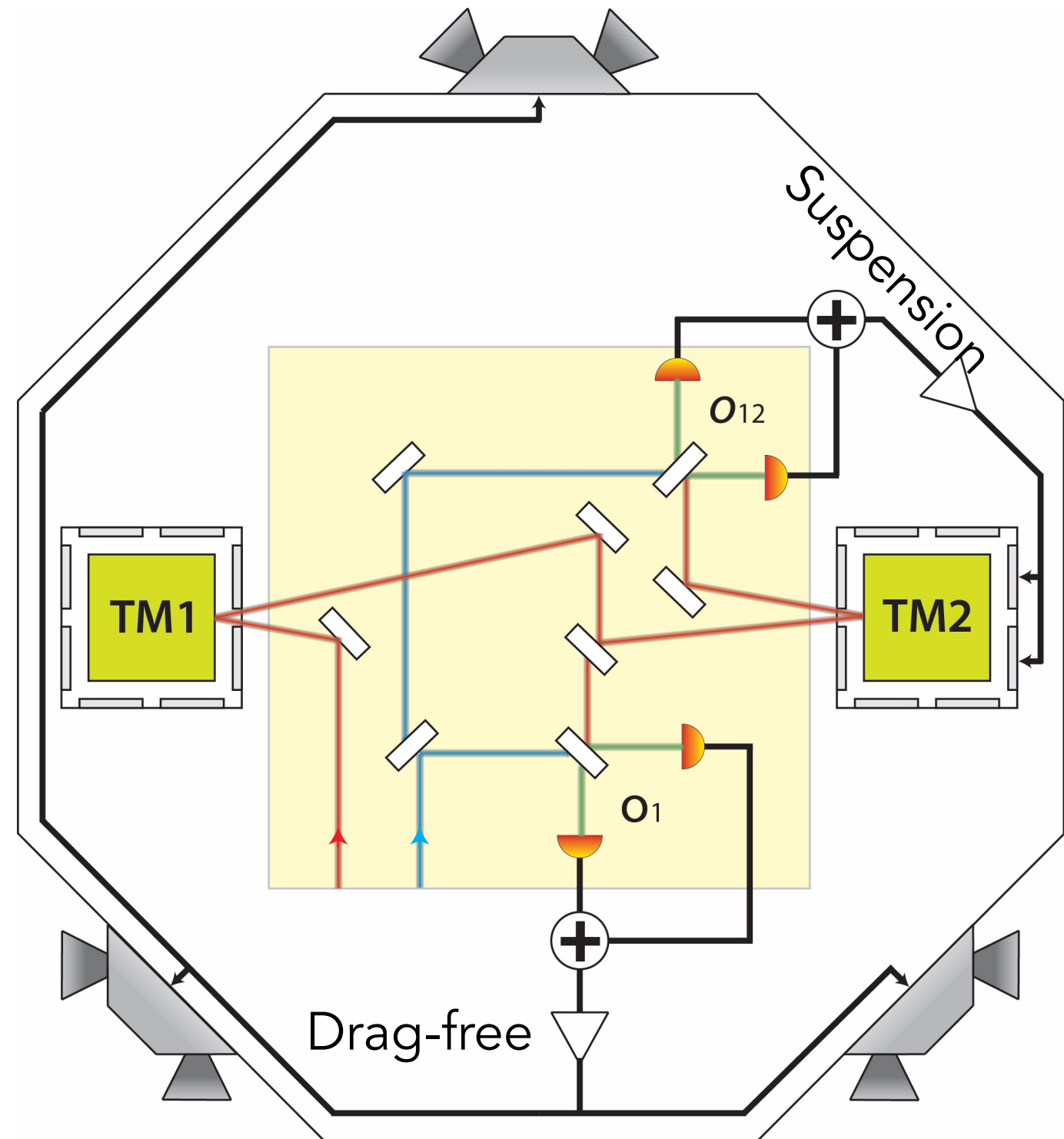
- Aims to test a lot of the LISA performance budget
 - all relevant local test mass disturbance forces
 - local interferometry
- And much of the critical technology
 - drag-free control
 - test mass charge control
 - gravitational balancing
 - SC environment control (thermal, magnetic)
- Additional payload from NASA (ST7) to demonstrate
 - alternative drag-free control
 - alternative micro-Newton thrusters

Control concept on LPF



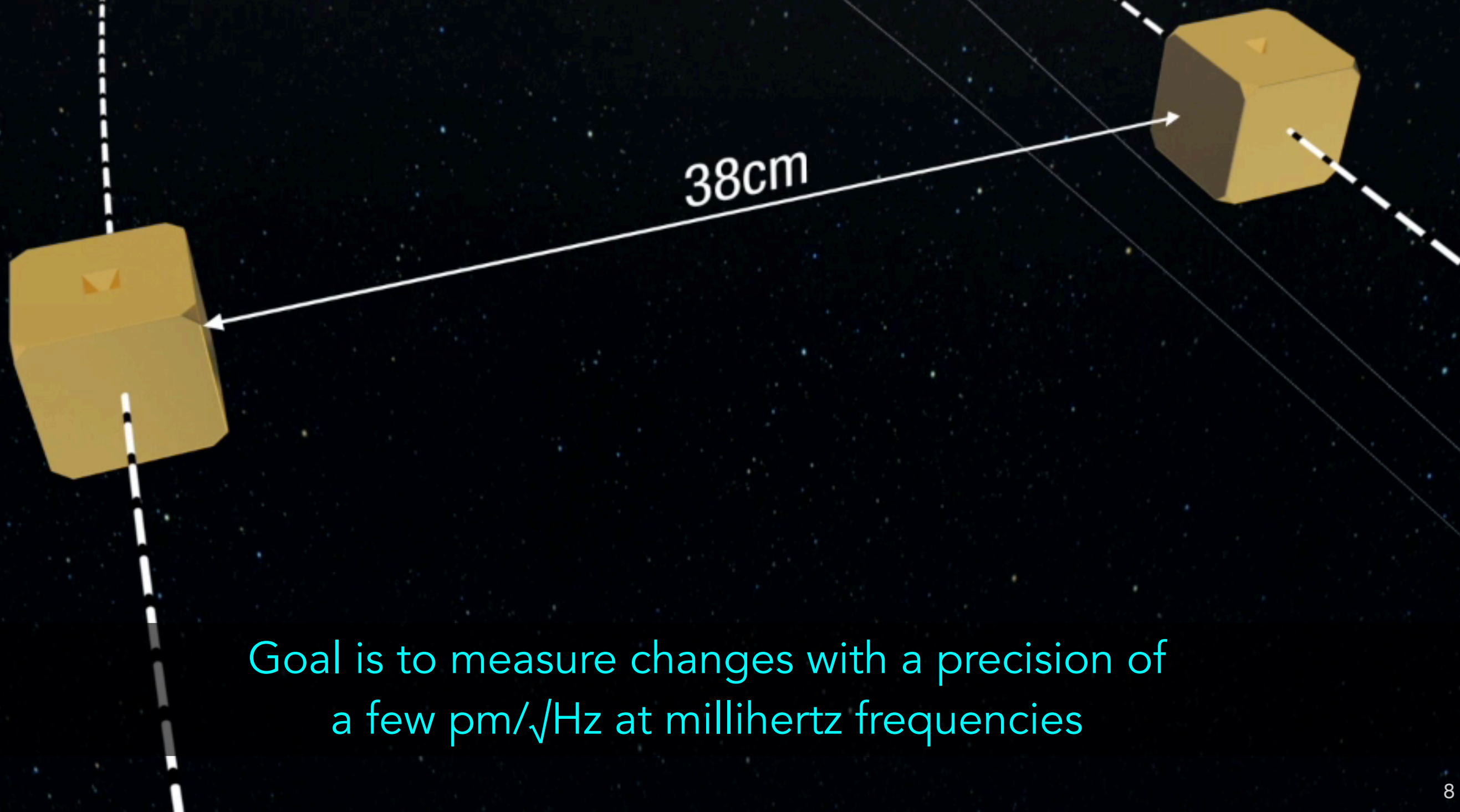
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- Shrink a LISA arm from 2.5 Gm to 38 cm
- Free-falling test mass
- Drag-free satellite
- Suspended witness mass



Metrology

Assess relative acceleration of test masses
by measuring their relative motion using an interferometer

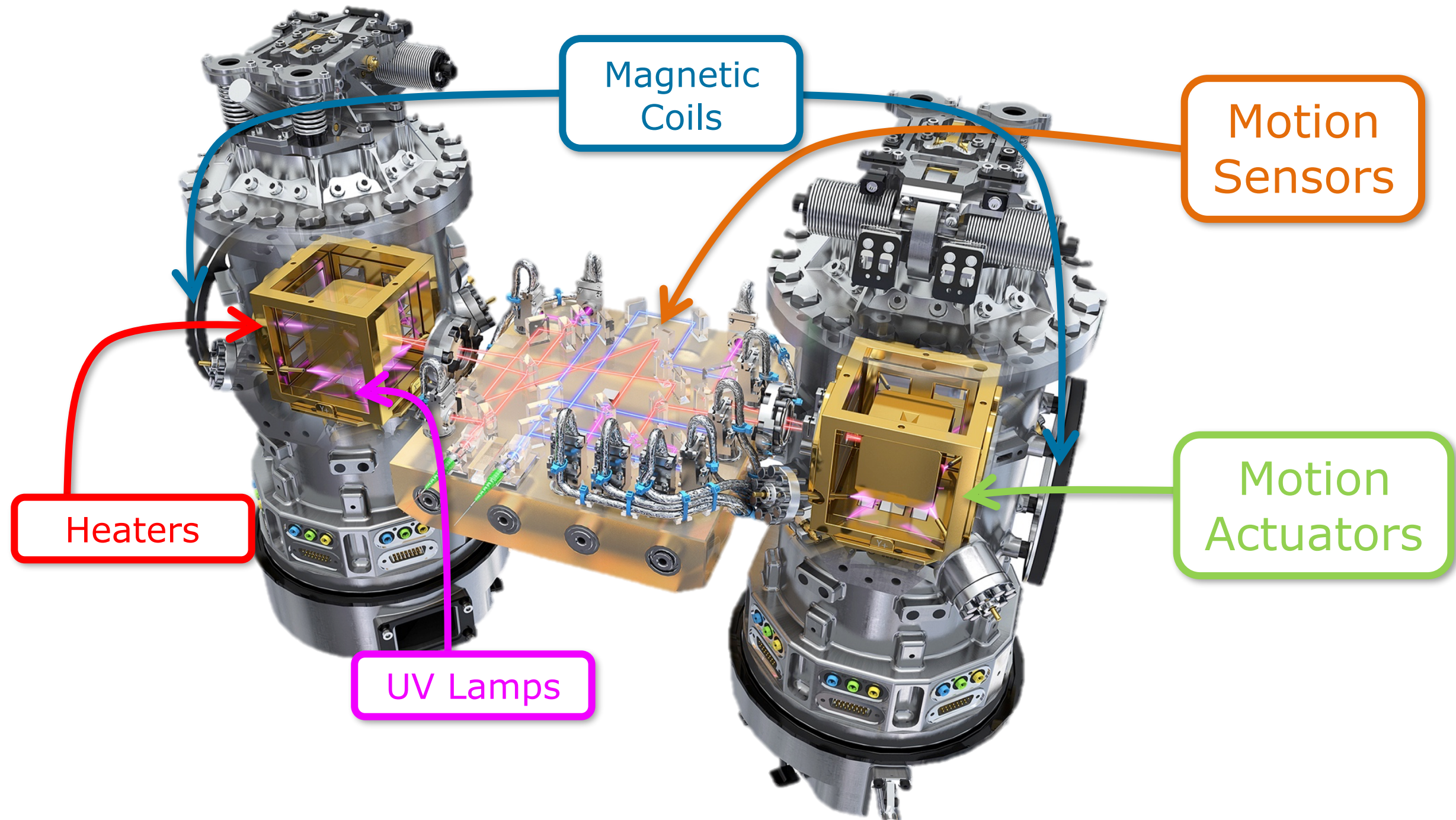


Goal is to measure changes with a precision of
a few $\text{pm}/\sqrt{\text{Hz}}$ at millihertz frequencies

Physics Lab in Space



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Timeline



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- All hardware delivered and integrated mid-2015
- Launch campaign from September to December
- Launch on December 3rd 2015
- Science operations since March 1st 2016

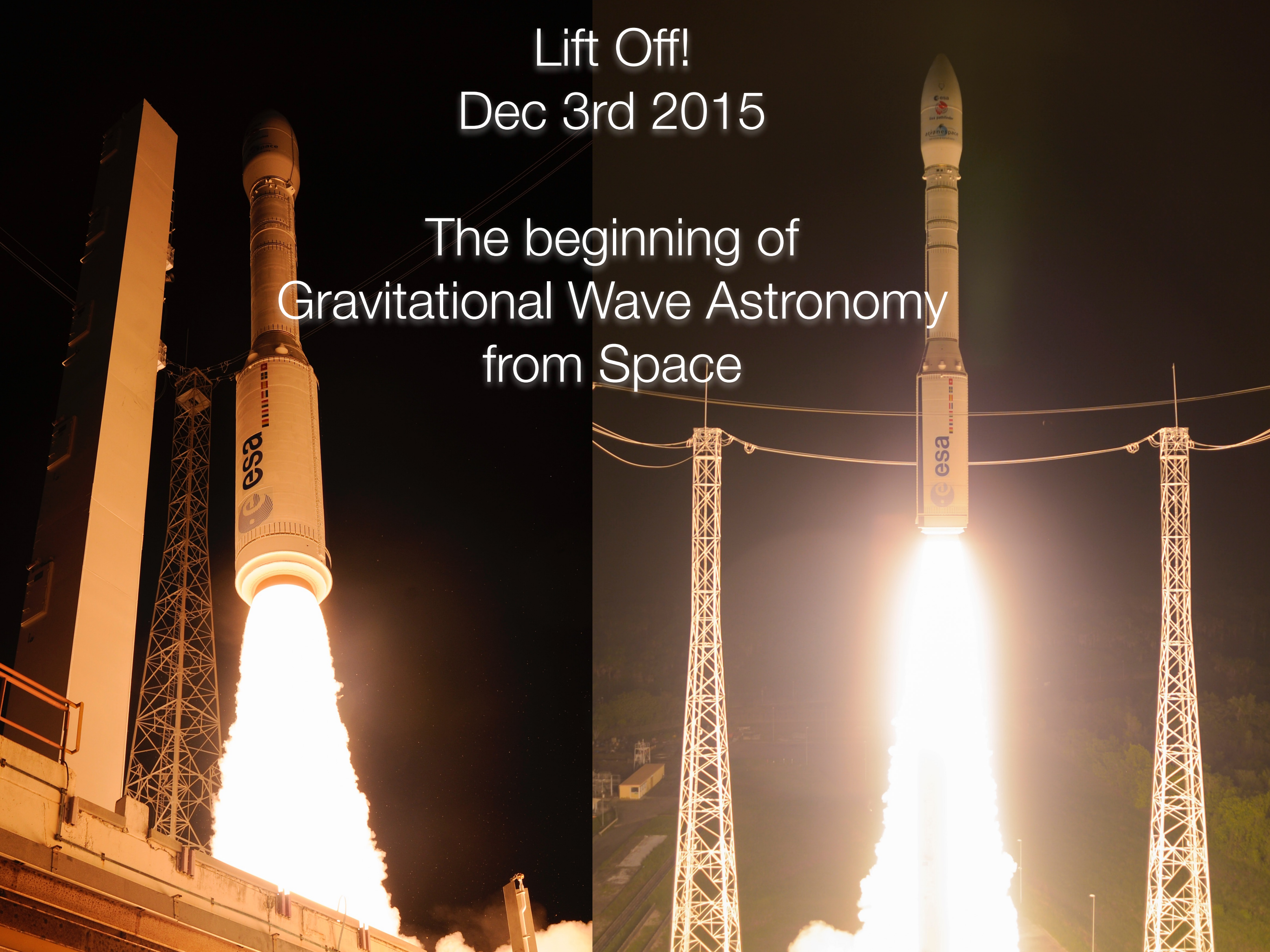






Lift Off!
Dec 3rd 2015

The beginning of
Gravitational Wave Astronomy
from Space



Sequence of events



Launch Dec 3rd 2015

Launch, Transfer

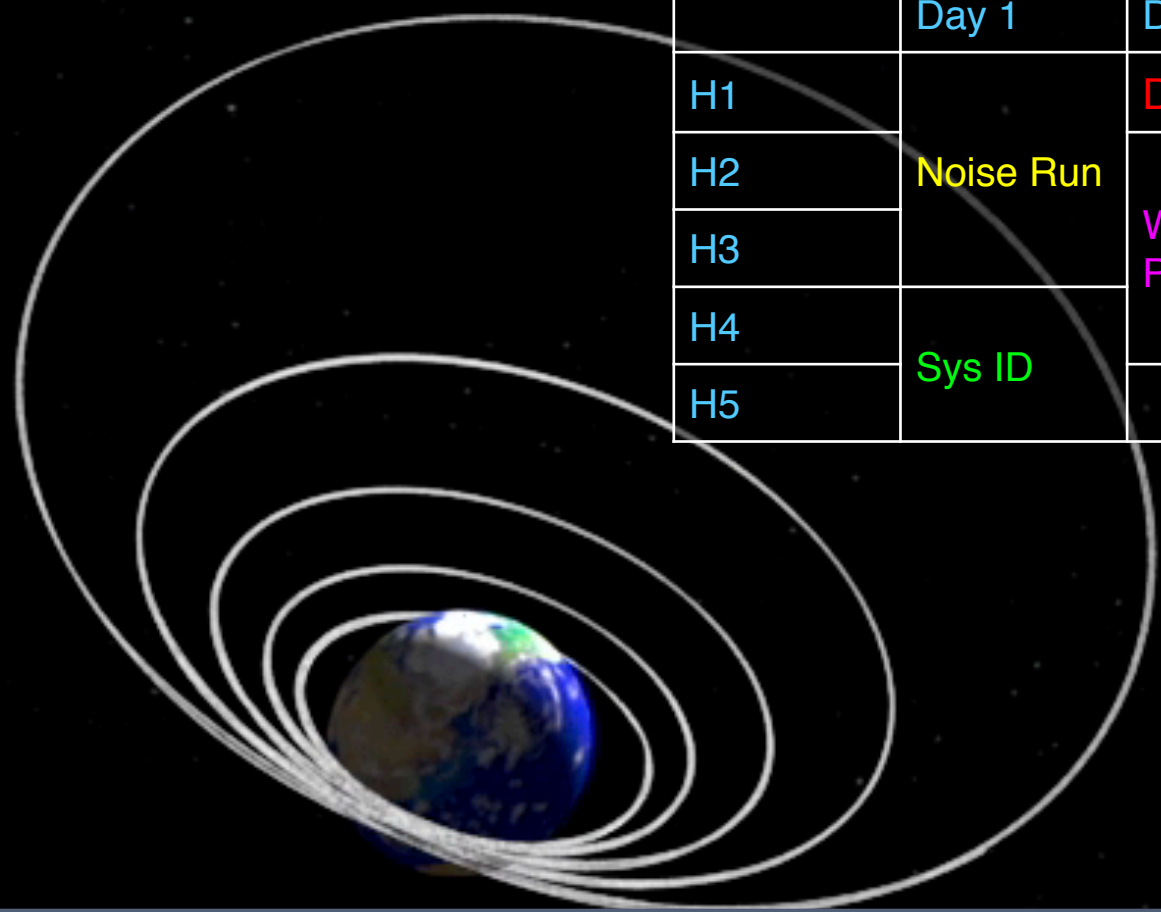
Industrial
Commissioning

IOCR

LTP Science
Ops
03-01 to 06-25

NASA Payload
Commissioning

NASA
Operations



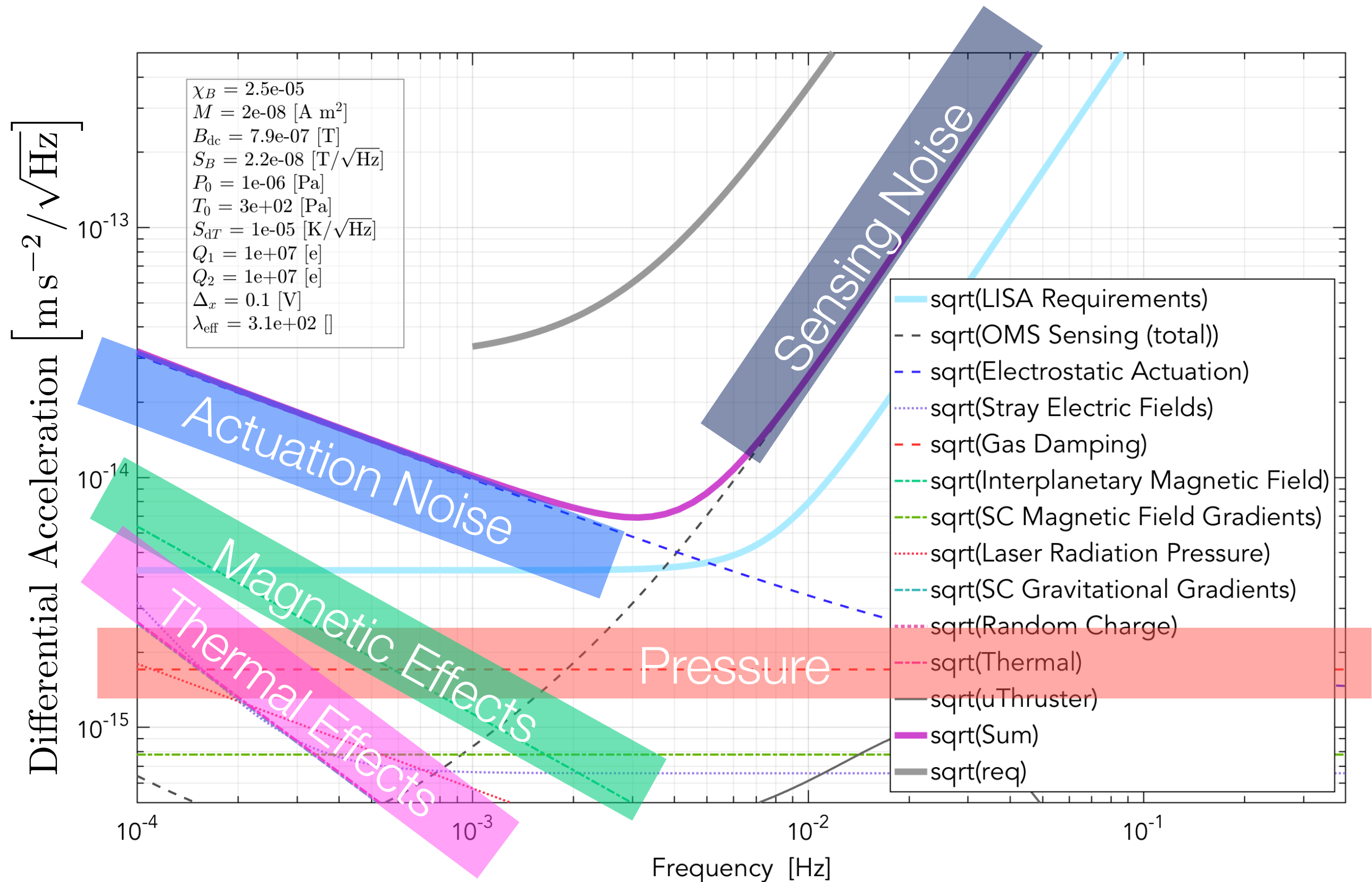
	Day 1	Day 2	Day 3	Day 4
H1	Noise Run	Discharge	Noise Run	Discharge
H2		Working Point		Stray Potentials
H3				
H4	Sys ID			
H5				



Understanding prior to launch



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How did we get there?



- Step 1: Launch
 - successfully completed on December 4th 2015
- Step 2: Commissioning
 - Switch on and check-out all units
 - Discard propulsion module
 - Decage TMs
 - Release TMs
 - Align OMS
 - Commission Drag-free control system
- Step 3: Science Operations

Science Operations & Data Analysis



- In-flight experiments were designed, simulated and tested before flight
- Sequence of experiments evolves with time
 - further simulation of operational weeks run continuously
- Analysis teams are on-duty 7 days a week in ESA's Operations Centre (ESOC, Darmstadt)
- First two weeks of operations were run essentially as planned
- From week 3 onwards, the plan was tailored to fit the observations
 - aiming to achieve a full understanding of the (excellent) measured performance



- 2011: Mock Data Challenges
 - help define algorithms
- 2012-2015: 45 STOC Exercises
 - face-to-face meetings between Scientists and ESA engineers
 - definition of interfaces, workflows in operations
 - technical details of all experiments
- 2012-2015: 5 STOC Simulations
 - simulate the operations scenario
 - from a few days to 2 weeks in length
 - involving full teams, including shifts
 - exercise all main experiments



- March 1st to June 26th
- 12 station keeping manoeuvres
- 31 unique configuration blocks
- 85 unique investigations

Nominal Phase



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~59 days

~4 days

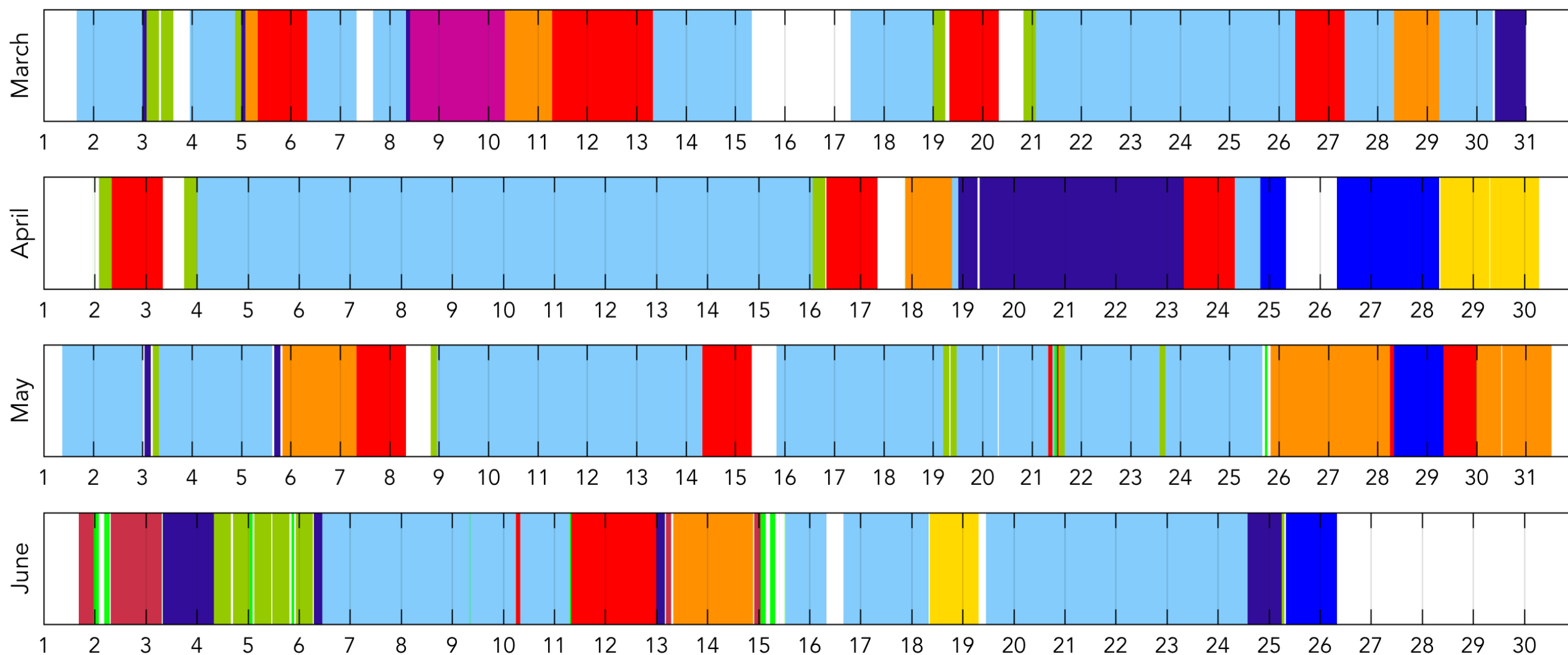
>3 days

~3 days

~4 days

~8 days

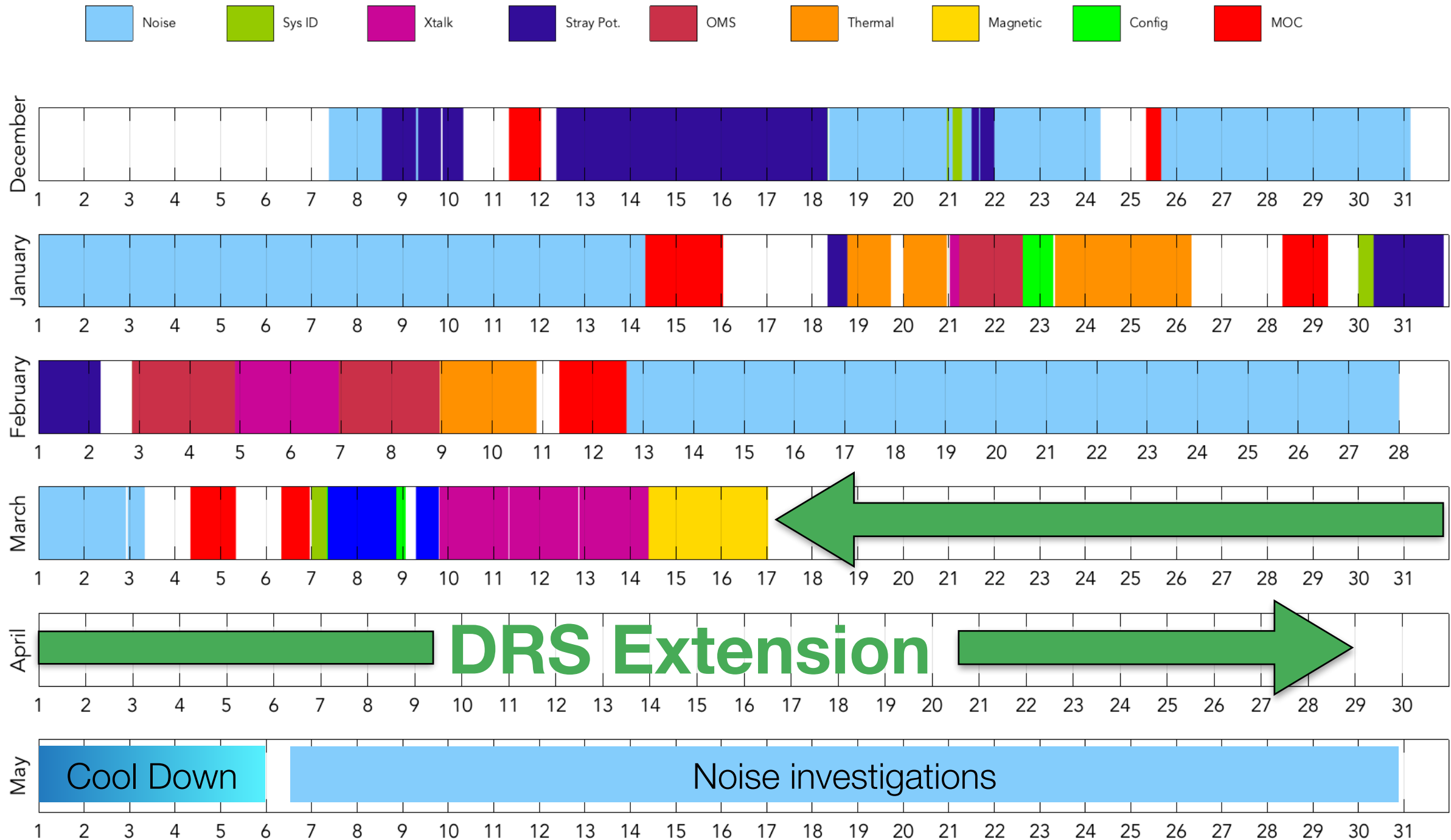
10 days



23



- Hand-back to LTP around Dec 7th



Primary science measurement



- The relative acceleration of the two test masses along the sensitive x-axis
- Representative of the force noise as it appears in a long LISA arm
- Estimated from the observation of the relative motion
 - made using the X12 interferometer
- Different corrections/calibrations made



- Raw Δg
- Inertial correction
 - centrifugal force on TMs due to the rotating and jittering SC
- Further correction
 - Debumped (cross-talk correction)
 - despite TM alignments done in hardware, we can still remove effects of SC jitter leaking into Δg
 - leakage of angular acceleration of SC



$$\Delta g^{\text{raw}} = \ddot{o}_{12} - G_{x_2} \delta(F_{\text{app}_2}, \tau) / M - \omega_2^2 o_{12}$$

- in-loop differential acceleration
- corrected for *applied* forces
 - gain correction factor (estimated from sys id)
 - commanded AC voltages
 - account for quantisation effects in FEE
- corrected for stiffness of TM2
 - estimated from dedicated system identification campaigns



- correct for centrifugal forces arising from the rotation of the observer frame
- steady rotation of SC together with SC jitter gives in-band noise
- not present in LISA

$$\Delta g^{\text{inertial}} = \Delta g^{\text{raw}} + \Delta g^{\text{cent}}$$

$$\Delta g^{\text{cent}} = \vec{\Omega} \times \vec{\Omega} \times \vec{r}$$

$$\vec{\Omega} = \vec{\Omega}_{\text{dc}} + \vec{\Omega}_{\text{noise}}$$

$$\vec{\Omega}_{\text{noise}} = \int G_{\phi} \frac{(N_{\phi_1} + N_{\phi_2})}{2I_{zz}} dt$$

$\vec{\Omega}_{\text{dc}}$ from low frequency SC angular,
estimated from AST quaternions

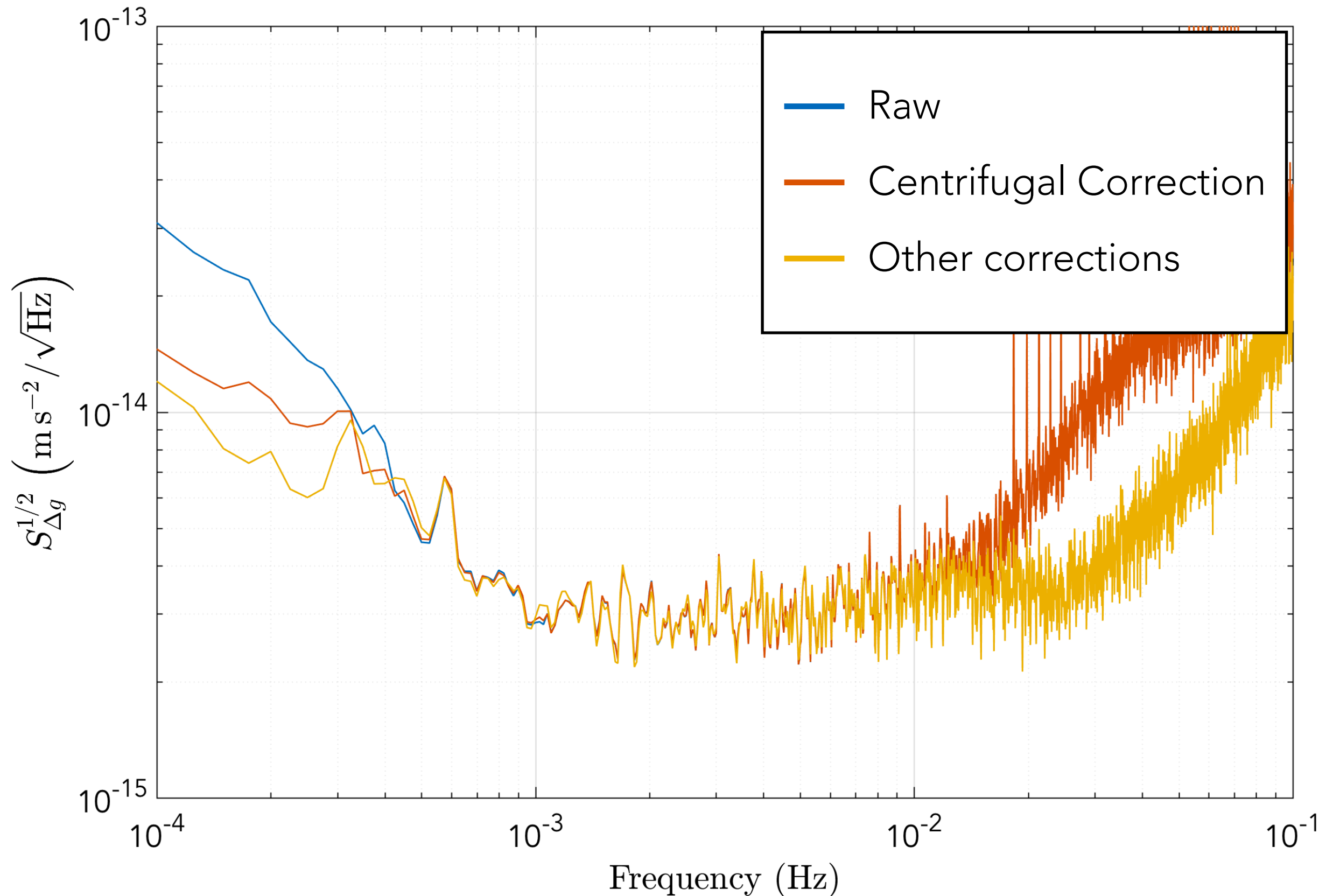


- De-bumping:
 - subtraction of SC jitter leaking into Δg
 - mitigate in hardware and post-processing:
 - Hardware de-bumping via test mass alignment
 - Post-processing: subtract a linear combination of other measured degrees-of-freedom
 - ad-hoc, but effective
- Tangential acceleration and actuation cross-talk
 - mis-alignment of optical axis and the mechanical axis joining the TMs
 - leakage of phi actuation along x (shared electrodes)
- Glitch removal
 - we have some relatively rare ($\sim 1/\text{Day}$) glitches
 - big enough to disturb spectral estimation
 - fit with simple model and subtract

Δg products



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Calibration of the x-axis



$$\Delta g^{\text{raw}} = o_{12}''$$

second derivative of
observed position (from IFO)

- numerical computation
- needs corrected time-stamping

$$- G_{x_2} \delta(F_{\text{app}_2}, \tau) / M$$

$$+ \omega_2^2 o_{12}$$

estimate of applied force

- commanded force not enough
- need modelling of actuation electronics
- delay in actuation
- gain in actuation

spring coupling to S/C

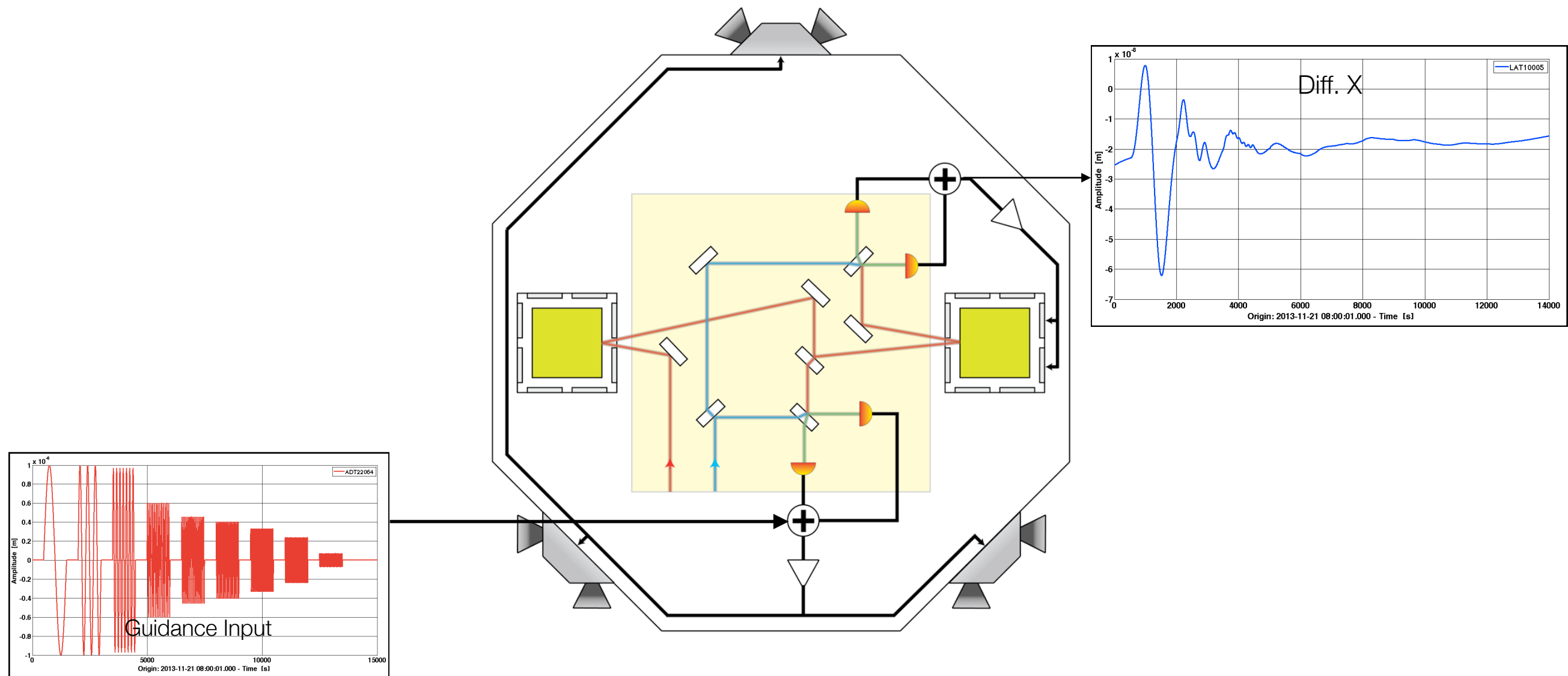
- estimate of the gravitational and electrostatic couplings
- modelled depending on actuation authority

Parameter estimation: the experiment



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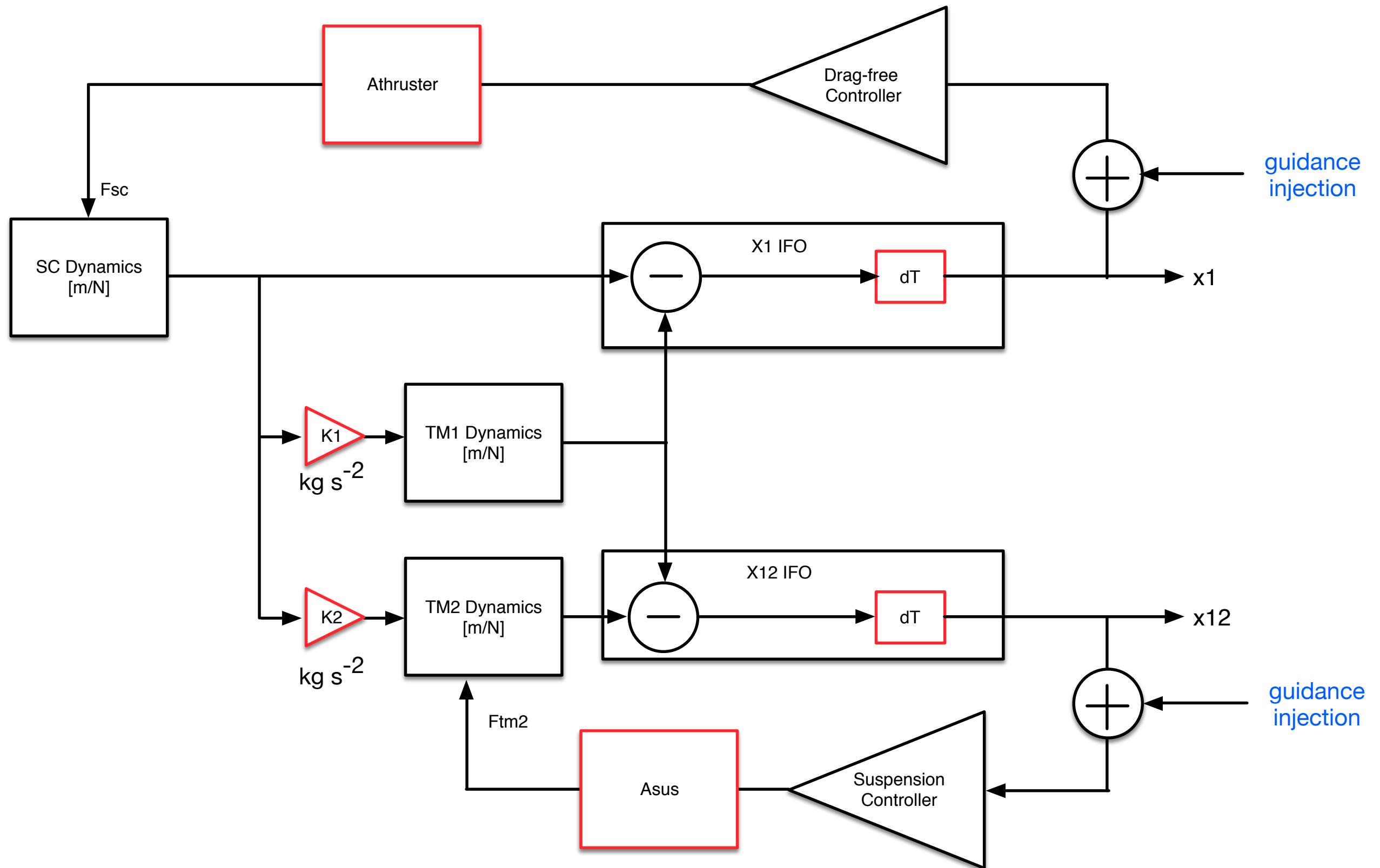
Goal is to measure the key parameters needed for estimating the residual differential acceleration



System control

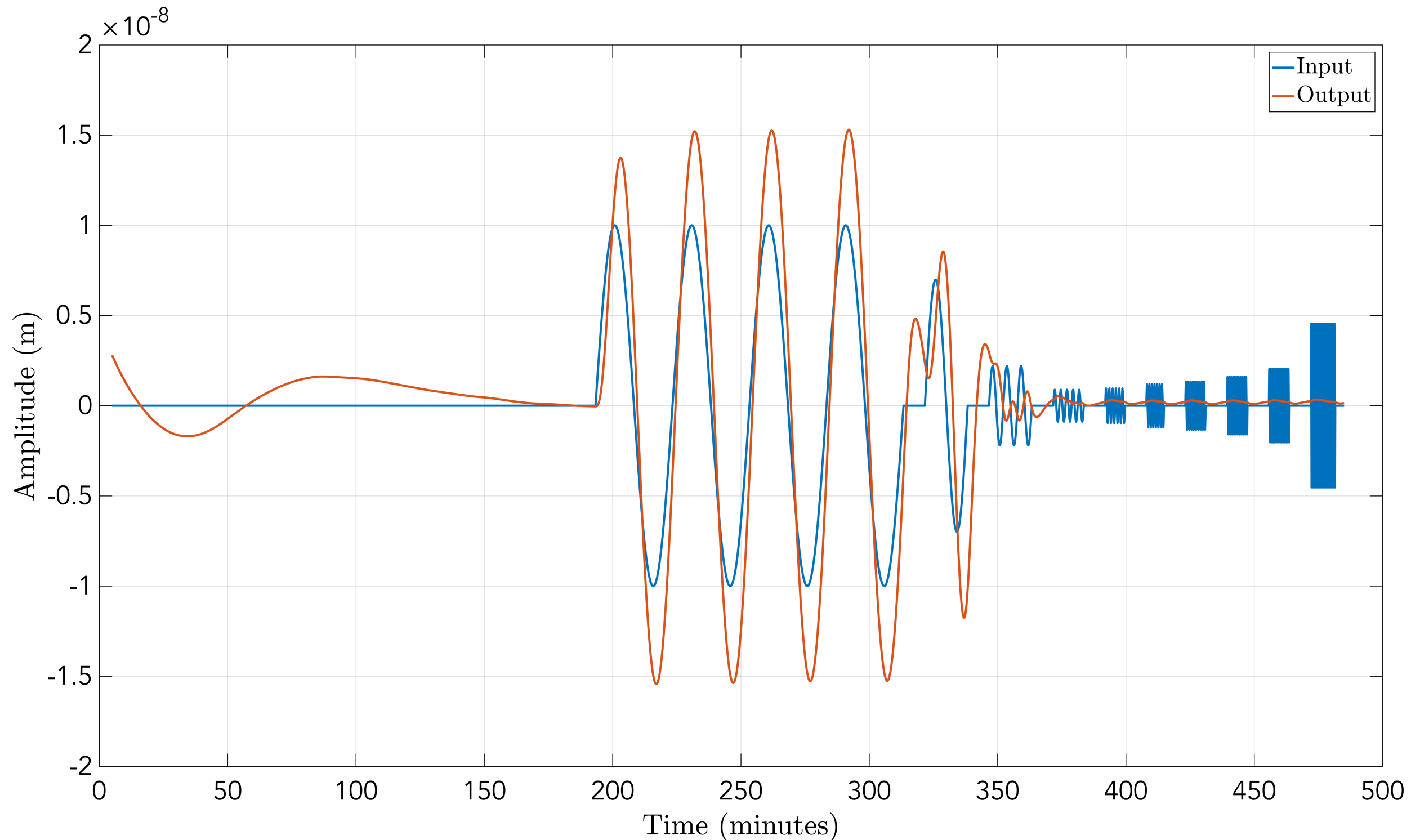


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signals on the x-axis suspension loop



Residuals after model fit



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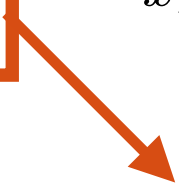
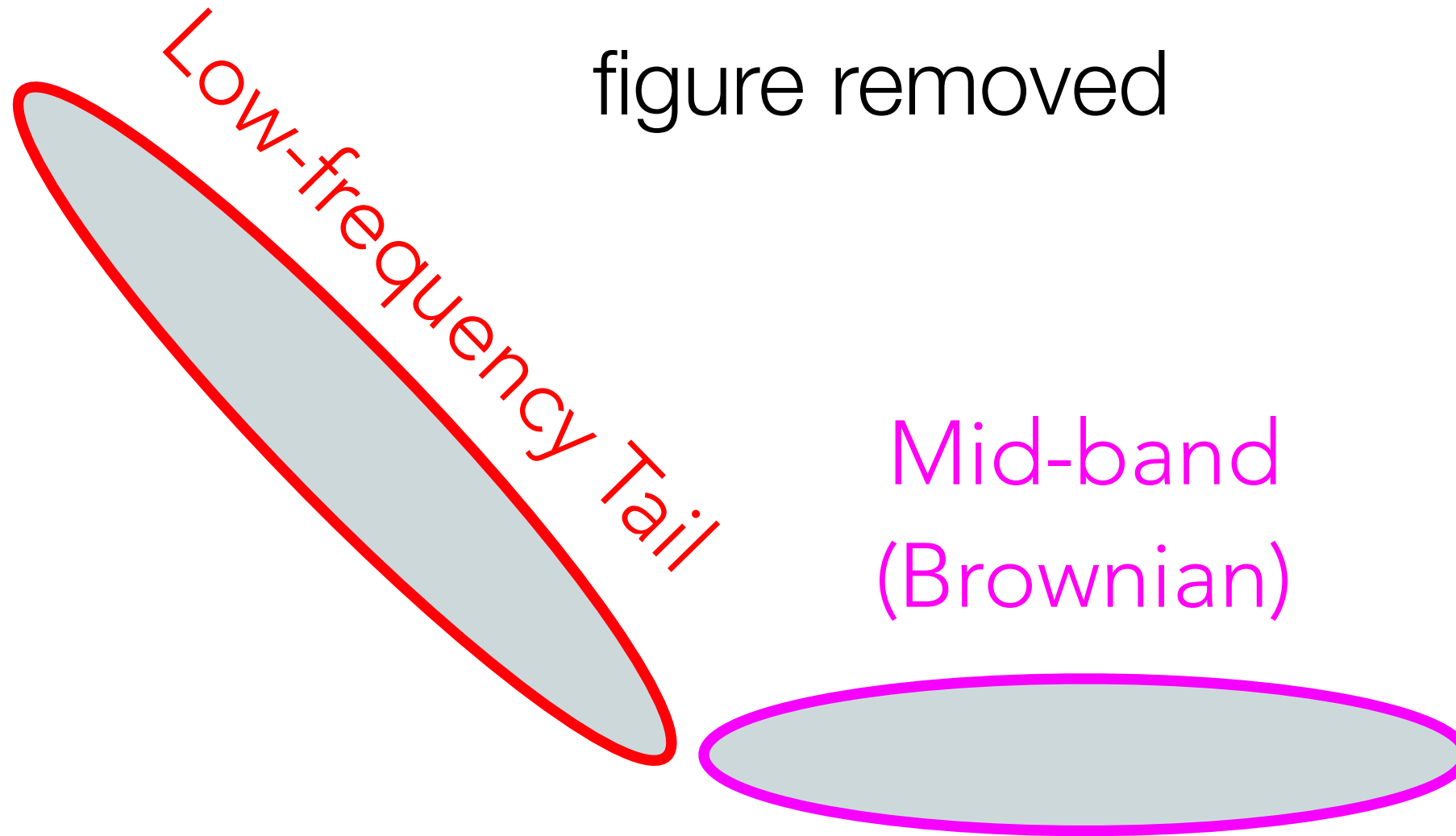
$$\Delta g^{\text{raw}} = \boxed{o_{12}''} - G_{x_2} \delta(F_{\text{app}_2}, \tau) / M - \omega_2^2 o_{12}$$


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Evolution of the noise



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- Mid-band: 1–10 mHz
- Dominated by residual gas around TMs
- Varies with temperature
- Varies with pressure
 - venting, outgassing



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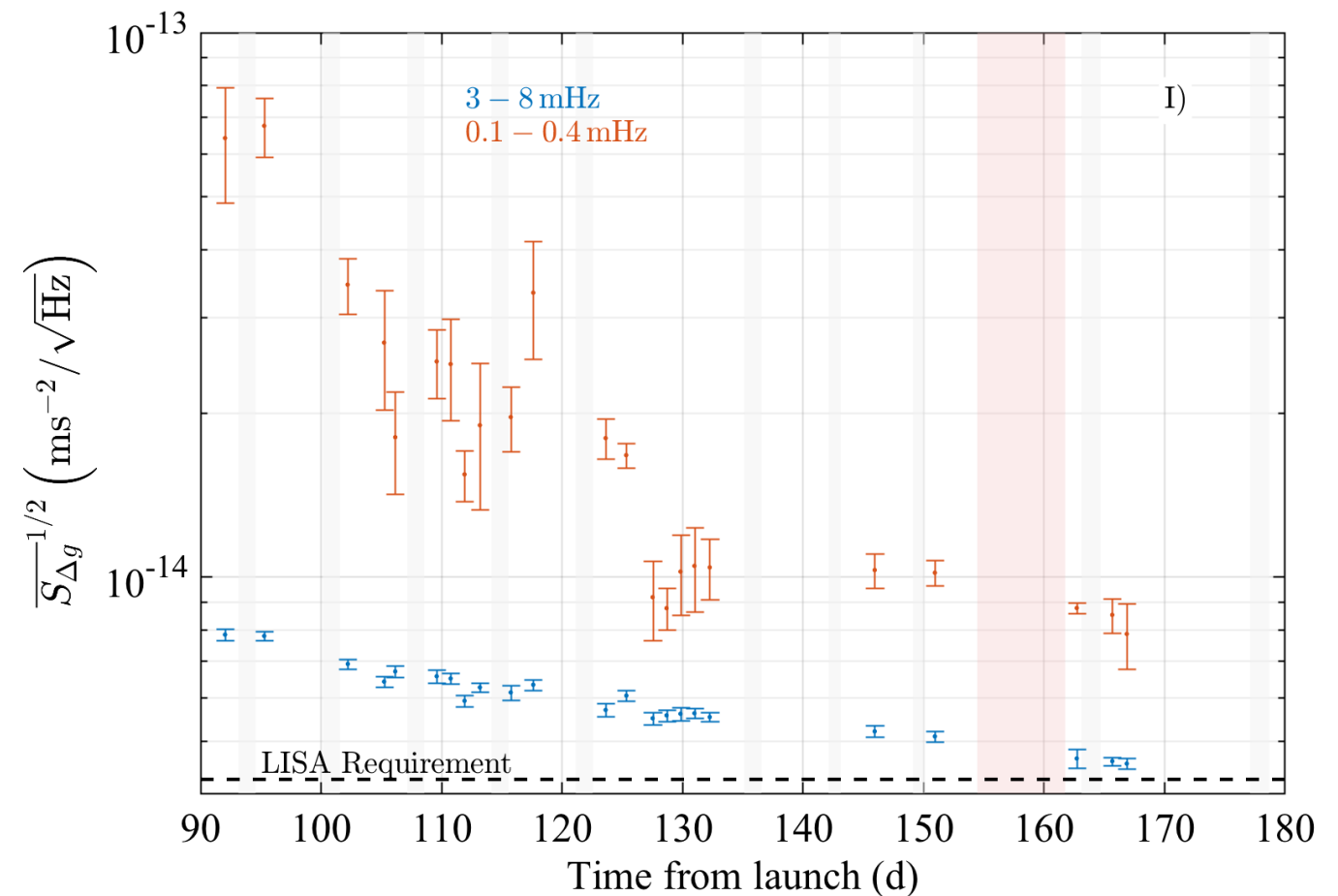


- In the early days, looked like it was improving with time
- In reality, our estimation of applied forces was not sufficient
$$\Delta g^{\text{raw}} = \ddot{o}_{12} - G_{x_2} \delta(F_{\text{app}_2}, \tau) / M - \omega_2^2 o_{12}$$
- We were fooled by the decreasing rate of change of the applied force
 - error in estimate of applied force reduced
- Applied force is properly estimated now
- Data is being fully reanalysed

In the early days...



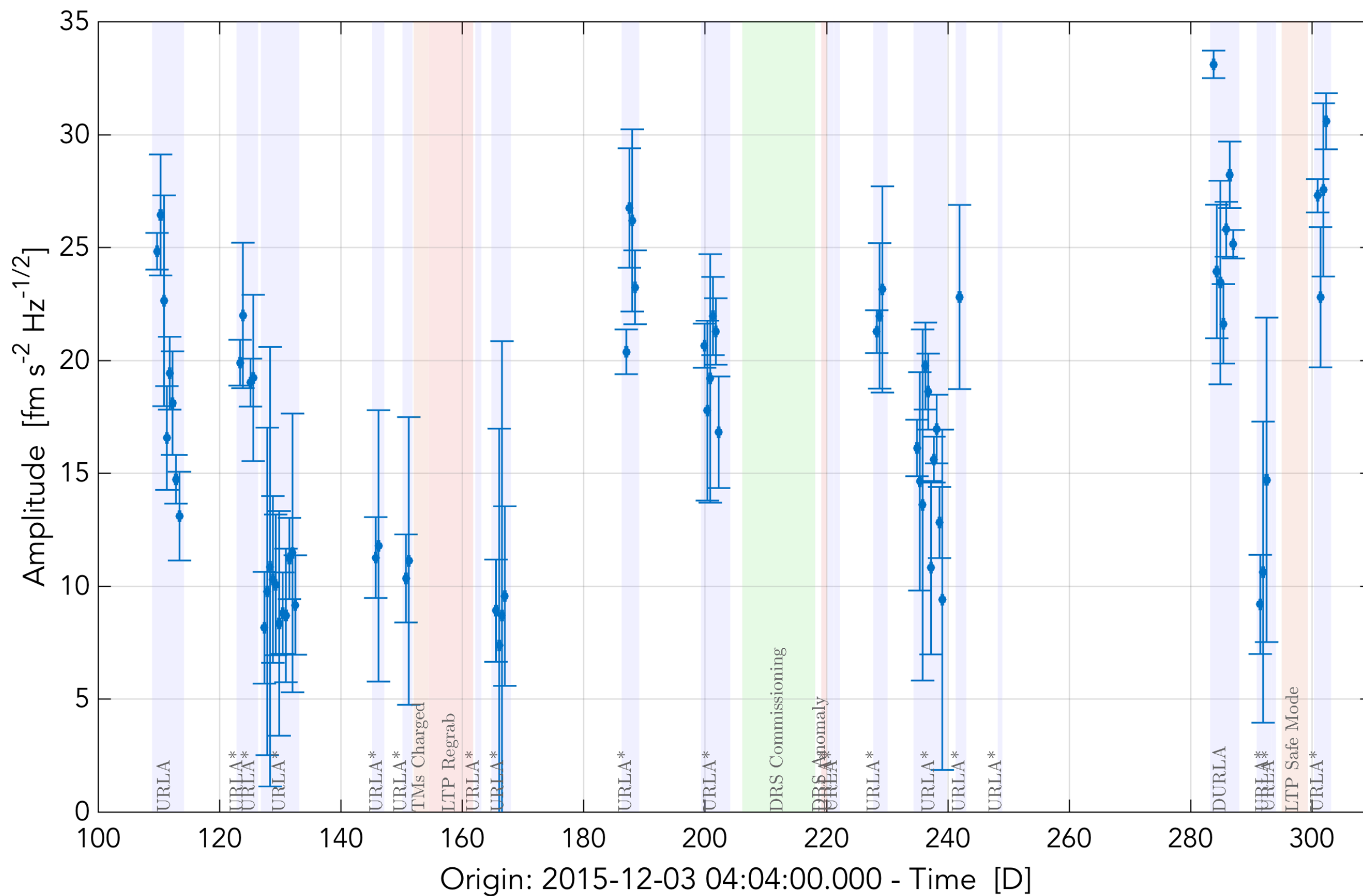
- During nominal operations we observed a decay of a low frequency noise
- Possible causes:
 1. thermo-mechanical relaxation of the system
 2. out-gassing effects creating pressure gradients which lead to forces if temperature drifts
 3. un-modelled actuation noise which depends on *slope* of dc force



with more data...



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Then with correctly estimated applied forces



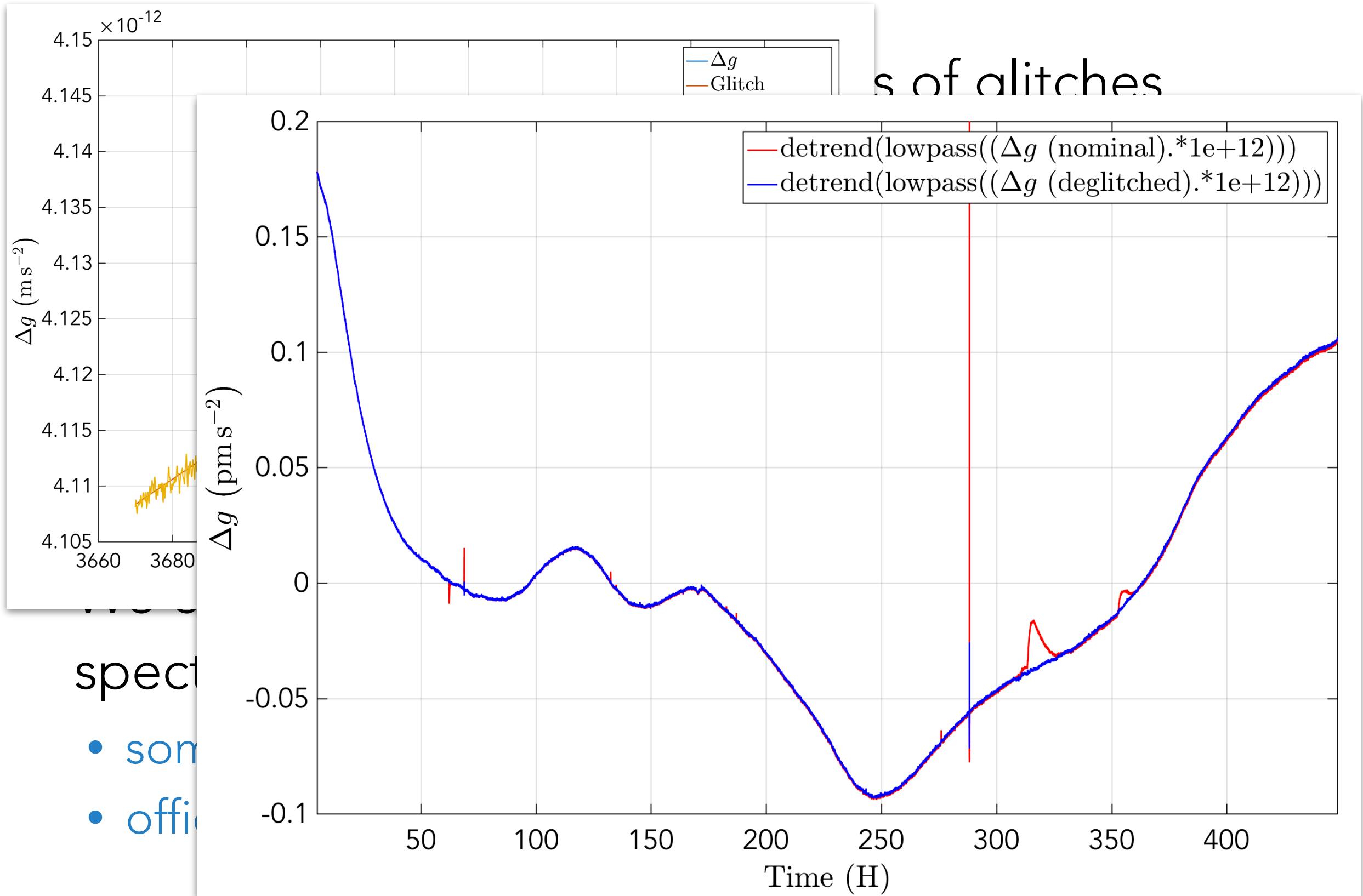
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Glitches, glitches everywhere, ...



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spect

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- Full re-analysis of delta-g segments underway
- Formal de-glitching pipeline is developed
- Final results to be summarised in up-coming publications
- Expectation is that the low frequency tail is rather stationary
 - still some decay
 - maybe associated with pressure and outgassing?



- All data to date is in the process of being reanalysed now our estimate of applied forces has improved
 - significant improvement to low frequency stationarity
- We made good use of slots in DRS phase to monitor performance and prepare for extension
- Extension is now nearly finished
 - began with our best long noise run
 - de-orbiting burn is complete - no more station keeping
 - first cool-down of LCA is complete, everything went well
 - 2nd cool-down yielded lots of glitches
 - return now to higher temperature
- Most other themes are well covered
 - thermal and gas environment of TMs
 - charge behaviour
 - OMS performance and noise budget
 - magnetic couplings
 - ...



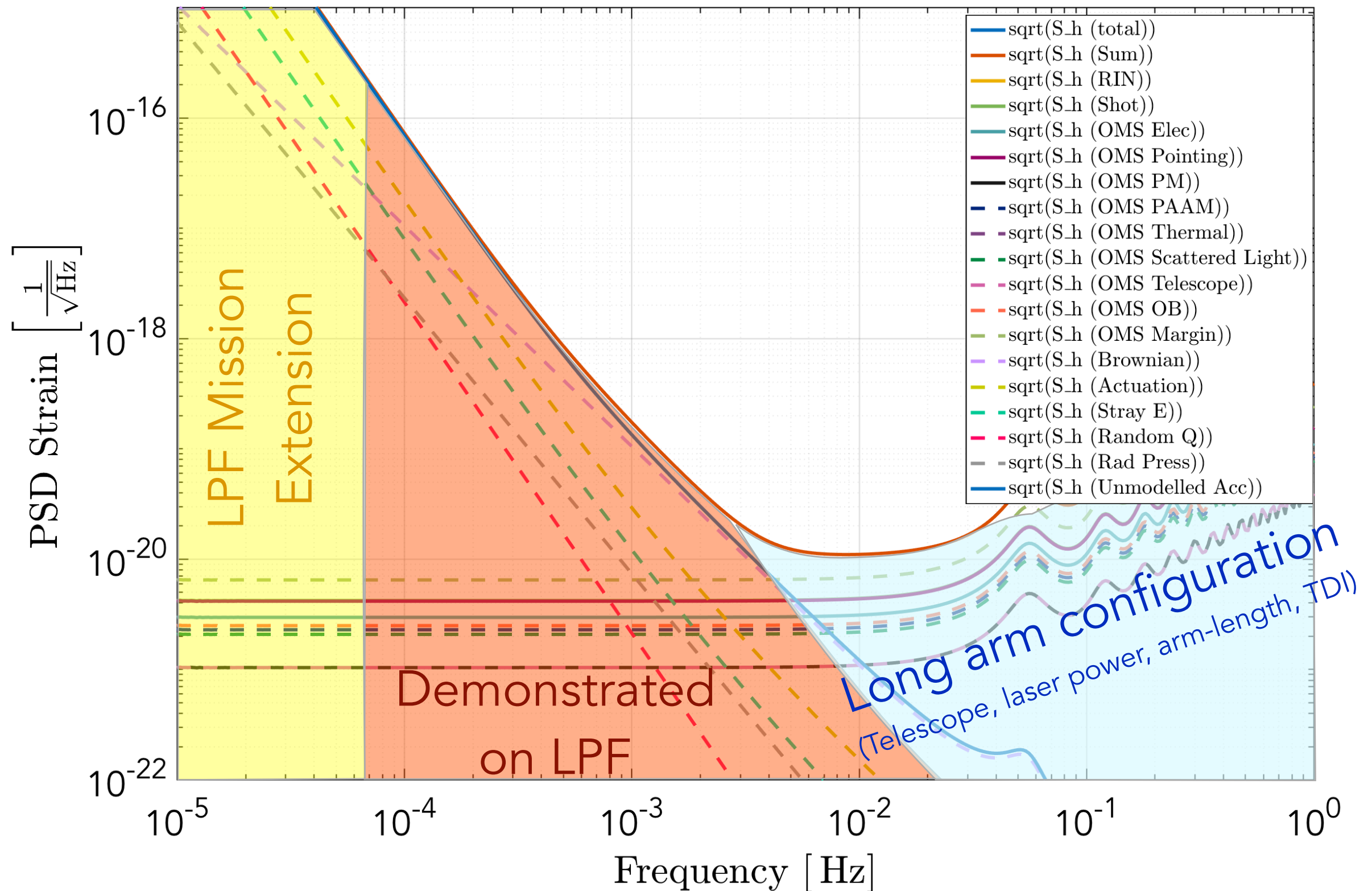
- low frequency excess noise remains top priority
 - seems fairly stationary for whole mission
 - unchanged by changes we've done (?)
 - some slight correlation with time and/or temperature?
 - how do we attack this?
 - cool-down again (done) and measure (underway)
 - heat-up again and measure
 - break correlation of temperature and time
- glitches
 - sources remain unknown
 - rates and zoology still being investigated
 - difficult to affect the rate/populations without hypotheses for the origin(s)
- Satellite to be decommissioned in early July and switched off on July 18th

Projection to LISA sensitivity



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$L=2.500000e+00$ Gm, $D_{\text{telescope}} = 0.30$ m, $P_{\text{wr}} = 1.26$ W





Thanks!